

LECTURE "HIGHER MATHEMATICS"

PART I

Vector Calculus

# 1 Spatial vectors

A vector  $\mathbf{a}$  in physical space is an arrow, pointing from some point  $P$  to another point  $Q$ . The **magnitude** (length)  $|\mathbf{a}| \geq 0$  of  $\mathbf{a}$  is defined as the linear distance between  $P$  and  $Q$ . Generally, a vector  $\mathbf{a}$  is fixed by its magnitude and its direction. As an arrow is shifted in space, without changing magnitude and direction, it still describes the same vector.

## 1.1 Vector algebra

As  $\mathbf{a}$  points from  $P$  to  $Q$  and  $\mathbf{b}$  from  $Q$  to  $R$ , **vector addition** is defined by

$$\mathbf{c} = \mathbf{a} + \mathbf{b}, \quad (1)$$

where  $\mathbf{c}$  is the vector pointing from  $P$  to  $R$ .

Multiplication of  $\mathbf{a}$  with a real number  $\lambda \in \mathbb{R}$  shall, by definition, yield another vector

$$\mathbf{a}' = \lambda \mathbf{a} \quad (2)$$

with length  $|\mathbf{a}'| = |\lambda||\mathbf{a}|$  and direction parallel ( $\lambda > 0$ ) or opposite ( $\lambda < 0$ ) to  $\mathbf{a}$ . For  $\lambda = -1$ , we also write

$$\lambda \mathbf{a} \equiv (-1)\mathbf{a} = -\mathbf{a}, \quad \mathbf{a} + (-1)\mathbf{b} = \mathbf{a} - \mathbf{b}. \quad (3)$$

We choose three mutually orthogonal directions "1", "2", and "3" (which shall form a right-handed system). Then, a typical vector is written as

$$\mathbf{a} = \begin{pmatrix} 2 \text{ m} \\ -5 \text{ m} \\ 4 \text{ m} \end{pmatrix} = \begin{pmatrix} a_1 \\ a_2 \\ a_3 \end{pmatrix} = \underbrace{a_1 \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}}_{=\mathbf{a}_1} + \underbrace{a_2 \begin{pmatrix} 0 \\ 1 \\ 0 \end{pmatrix}}_{=\mathbf{a}_2} + \underbrace{a_3 \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}}_{=\mathbf{a}_3}, \quad (4)$$

where  $\mathbf{a}_1$  is the vector with length  $a_1 = 2 \text{ m}$  pointing in direction 1,  $\mathbf{a}_2$  is the vector with length  $-a_2 = 5 \text{ m}$  pointing opposite to direction 2,  $\mathbf{a}_3$  is the vector with length  $a_3 = 4 \text{ m}$  pointing in direction 3.

The dimensionless objects

$$\mathbf{u}_1 = \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}, \quad \mathbf{u}_2 = \begin{pmatrix} 0 \\ 1 \\ 0 \end{pmatrix}, \quad \mathbf{u}_3 = \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}, \quad (5)$$

with  $|\mathbf{u}_n| = 1$ , are called **unit vectors**,

$$\mathbf{a} = a_1 \mathbf{u}_1 + a_2 \mathbf{u}_2 + a_3 \mathbf{u}_3. \quad (6)$$

The **numbers**  $a_n$  are called the **coordinates** of the vector  $\mathbf{a}$ . Theorem of Pythagoras:

$$|\mathbf{a}| = \sqrt{a_1^2 + a_2^2 + a_3^2}. \quad (7)$$

The **vectors**  $\mathbf{a}_n$  are called the **components** of the vector  $\mathbf{a}$ .

Vectors may have **units other than length**. While the  $a_n$  are carrying the physical units of the vectorial quantity  $\mathbf{a}$ , the  $\mathbf{u}_n$  are "dimensionless". E.g., a **force vector** is

$$\begin{aligned}\mathbf{F} &= F_1 \mathbf{u}_1 + F_2 \mathbf{u}_2 + F_3 \mathbf{u}_3 \\ &= (2.13 \text{ N}) \mathbf{u}_1 + (-1.95 \text{ N}) \mathbf{u}_2 + (4.79 \text{ N}) \mathbf{u}_3 \\ &= (2.13 \text{ N}) \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix} + (-1.95 \text{ N}) \begin{pmatrix} 0 \\ 1 \\ 0 \end{pmatrix} + (4.79 \text{ N}) \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix} = \begin{pmatrix} 2.13 \text{ N} \\ -1.95 \text{ N} \\ 4.79 \text{ N} \end{pmatrix}. \quad (8)\end{aligned}$$

Its **magnitude** is  $|\mathbf{F}| = \sqrt{F_1^2 + F_2^2 + F_3^2} \approx 5.59 \text{ N}$ .

## 1.2 Scalar (dot) and vector (cross) products

The **scalar product** (or **dot product**) of two vectors  $\mathbf{a}$  and  $\mathbf{b}$  is defined as

$$\mathbf{a} \cdot \mathbf{b} := |\mathbf{a}||\mathbf{b}| \cos \gamma \quad (\text{a number !}), \quad (9)$$

where  $\gamma$  is the **angle** between  $\mathbf{a}$  and  $\mathbf{b}$ .

Consequently,  $\mathbf{b} \cdot \mathbf{a} = \mathbf{a} \cdot \mathbf{b}$  and, with some  $\lambda \in \mathbb{R}$ ,  $(\lambda \mathbf{a}) \cdot \mathbf{b} = \lambda(\mathbf{a} \cdot \mathbf{b})$ .

Furthermore,  $(\mathbf{a} + \mathbf{b}) \cdot \mathbf{c} = \mathbf{a} \cdot \mathbf{c} + \mathbf{b} \cdot \mathbf{c}$  (see problem 1.1), and thus

$$\begin{aligned}\mathbf{a} \cdot \mathbf{b} &\equiv (a_1 \mathbf{u}_1 + a_2 \mathbf{u}_2 + a_3 \mathbf{u}_3) \cdot (b_1 \mathbf{u}_1 + b_2 \mathbf{u}_2 + b_3 \mathbf{u}_3) \\ &\equiv \left( \sum_{i=1}^3 a_i \mathbf{u}_i \right) \cdot \left( \sum_{j=1}^3 b_j \mathbf{u}_j \right) \\ &= \sum_{i=1}^3 \sum_{j=1}^3 a_i b_j \underbrace{(\mathbf{u}_i \cdot \mathbf{u}_j)}_{\delta_{ij}} = \sum_{i=1}^3 a_i b_i \equiv a_1 b_1 + a_2 b_2 + a_3 b_3. \quad (10)\end{aligned}$$

Therefore, in terms of the coordinates,

$$\mathbf{a} \cdot \mathbf{b} \equiv \begin{pmatrix} a_1 \\ a_2 \\ a_3 \end{pmatrix} \cdot \begin{pmatrix} b_1 \\ b_2 \\ b_3 \end{pmatrix} = a_1 b_1 + a_2 b_2 + a_3 b_3. \quad (11)$$

Let  $\mathbf{u}_\perp$  be the unit vector perpendicular to both  $\mathbf{a}$  and  $\mathbf{b}$ , such that  $\{\mathbf{a}, \mathbf{b}, \mathbf{u}_\perp\}$  is right-handed. The **vector product** (or **cross product**) of  $\mathbf{a}$  with  $\mathbf{b}$  is defined as

$$\mathbf{a} \times \mathbf{b} := (|\mathbf{a}||\mathbf{b}| \sin \gamma) \mathbf{u}_\perp \quad (\text{a vector !}). \quad (12)$$

Since  $\mathbf{b} \times \mathbf{a} = -\mathbf{a} \times \mathbf{b}$  and  $(\lambda \mathbf{a}) \times \mathbf{b} = \lambda(\mathbf{a} \times \mathbf{b})$ , etc. (exercises), we now have

$$\mathbf{a} \times \mathbf{b} \equiv \dots = (a_2 b_3 - a_3 b_2) \mathbf{u}_1 + (a_3 b_1 - a_1 b_3) \mathbf{u}_2 + (a_1 b_2 - a_2 b_1) \mathbf{u}_3. \quad (13)$$

Therefore, in terms of the coordinates,

$$\begin{pmatrix} a_1 \\ a_2 \\ a_3 \end{pmatrix} \times \begin{pmatrix} b_1 \\ b_2 \\ b_3 \end{pmatrix} = \begin{pmatrix} a_2 b_3 - a_3 b_2 \\ a_3 b_1 - a_1 b_3 \\ a_1 b_2 - a_2 b_1 \end{pmatrix}. \quad (14)$$

## 2 Curves in space: Vector functions of one argument

### 2.1 Orbit of a moving particle as an example

Selecting one point  $O$  in space as origin, any other point  $P$  is fixed by its **position vector**

$$\mathbf{r} = \vec{OP} = x\mathbf{u}_1 + y\mathbf{u}_2 + z\mathbf{u}_3. \quad (15)$$

The position vector of a **moving particle** has time-dependent coordinates  $x, y$ , and  $z$ ,

$$\mathbf{r}(t) = \begin{pmatrix} x(t) \\ y(t) \\ z(t) \end{pmatrix}. \quad (16)$$

This is a **vector function**. As  $t$  grows, the tip of the vector  $\mathbf{r}(t)$  “draws” the **orbit** of the particle which is a **curve in space**.

**Example 1a:** Motion on a helical (helix shaped) orbit is described by the function

$$\mathbf{r}(t) = \begin{pmatrix} R \cos(\omega t) \\ R \sin(\omega t) \\ ut \end{pmatrix}. \quad (17)$$

### 2.2 Derivative of vector functions: velocity vector

The **derivative** of a vector function  $\mathbf{r}(t)$  is another vector function  $\dot{\mathbf{r}}(t)$ , defined as

$$\dot{\mathbf{r}}(t) \equiv \frac{d}{dt} \mathbf{r}(t) := \lim_{\Delta t \rightarrow 0} \frac{\mathbf{r}(t + \Delta t) - \mathbf{r}(t)}{\Delta t} \quad (18)$$

$$\equiv \lim_{\Delta t \rightarrow 0} \frac{1}{\Delta t} \begin{pmatrix} x(t + \Delta t) - x(t) \\ y(t + \Delta t) - y(t) \\ z(t + \Delta t) - z(t) \end{pmatrix} \equiv \begin{pmatrix} \dot{x}(t) \\ \dot{y}(t) \\ \dot{z}(t) \end{pmatrix}. \quad (19)$$

The vectors  $\mathbf{r}(t)$  and  $\mathbf{r}(t + \Delta t)$  (blue) and their difference (green) are shown in Fig. 1.

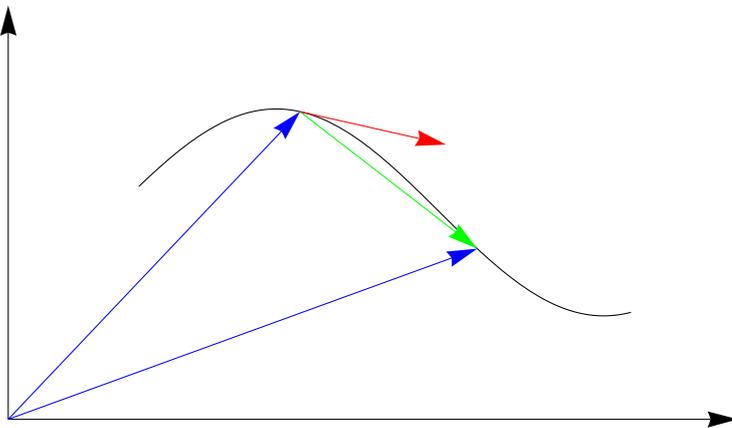


Figure 1: .

At any given time  $t$ , the **magnitude** of the vector  $\dot{\mathbf{r}}(t)$  is

$$\begin{aligned} |\dot{\mathbf{r}}(t)| &= \left| \lim_{\Delta t \rightarrow 0} \frac{\mathbf{r}(t + \Delta t) - \mathbf{r}(t)}{\Delta t} \right| \\ &= \lim_{\Delta t \rightarrow 0} \frac{|\mathbf{r}(t + \Delta t) - \mathbf{r}(t)|}{\Delta t}. \end{aligned} \quad (20)$$

For sufficiently small  $\Delta t$  (when orbital curvature becomes negligible), the numerator  $|\mathbf{r}(t + \Delta t) - \mathbf{r}(t)|$  approaches the distance  $\Delta s$  on the orbit, travelled by the particle between times  $t$  and  $t + \Delta t$ . Consequently, the limit (20) is the temporary **speed**  $v(t)$  of the particle,

$$|\dot{\mathbf{r}}(t)| \equiv \sqrt{\dot{x}(t)^2 + \dot{y}(t)^2 + \dot{z}(t)^2} = \frac{\Delta s}{\Delta t} = v(t). \quad (21)$$

As is obvious from the figure, the **direction** of the vector  $\dot{\mathbf{r}}(t)$ , indicated in red color, is **tangential** to the orbit at  $\mathbf{r}(t)$ . Therefore, the vector

$$\mathbf{v}(t) \equiv \dot{\mathbf{r}}(t) \quad (22)$$

is called the **velocity (vector)** of the particle.

**Example 1b:** In Example 1a, we obtain

$$\dot{\mathbf{r}}(t) = \begin{pmatrix} -R\omega \sin(\omega t) \\ +R\omega \cos(\omega t) \\ u \end{pmatrix}, \quad v(t) = \sqrt{(R\omega)^2 + u^2} \equiv v. \quad (23)$$

### 2.3 Parametrization of a curve

Instead of  $t$ , any monotonic function  $p(t)$  may serve as a parameter  $p$  of a curve,

$$\mathbf{r}(p) = \begin{pmatrix} x(p) \\ y(p) \\ z(p) \end{pmatrix}. \quad (24)$$

A typical **example** is the **arc length**  $p(t) = s(t)$  along the curve,

$$s(t) = \int_0^t dt' |\dot{\mathbf{r}}(t')| \equiv \int_0^t dt' \sqrt{\dot{x}(t')^2 + \dot{y}(t')^2 + \dot{z}(t')^2}. \quad (25)$$

**Example 1c:** For the curve of Example 1a, we obtain

$$s(t) = \int_0^t dt' \sqrt{(R\omega)^2 + u^2} \equiv vt \quad \Rightarrow \quad \mathbf{r}(s) = \begin{pmatrix} R \cos\left(\frac{\omega}{v}s\right) \\ R \sin\left(\frac{\omega}{v}s\right) \\ \frac{u}{v}s \end{pmatrix}. \quad (26)$$

**Exercise:** Writing  $\frac{d\mathbf{r}(p)}{dp} = \dot{\mathbf{r}}(p)$ , we find **chain** and **product rules** for vector functions

$$\frac{d}{dt} \mathbf{r}(p(t)) = \left[ \dot{\mathbf{r}}(p) \right] \Big|_{p=p(t)} \dot{p}(t), \quad (27)$$

$$\frac{d}{dp} \left[ \mathbf{r}_1(p) \cdot \mathbf{r}_2(p) \right] = \dot{\mathbf{r}}_1(p) \cdot \mathbf{r}_2(p) + \mathbf{r}_1(p) \cdot \dot{\mathbf{r}}_2(p). \quad (28)$$

## 2.4 Unit tangent vector

Since the vector  $\dot{\mathbf{r}}(p)$  is tangential to the orbit, the ("normalized") vector

$$\mathcal{T}(p) = \frac{\dot{\mathbf{r}}(p)}{|\dot{\mathbf{r}}(p)|} = \begin{pmatrix} \mathcal{T}_1(p) \\ \mathcal{T}_2(p) \\ \mathcal{T}_3(p) \end{pmatrix} \quad (29)$$

is the **unit tangent vector** of the orbit.

As a consequence,  $\mathcal{T}(p) \cdot \mathcal{T}(p) = 1$  and the **product rule** yields

$$\begin{aligned} 0 &= \frac{d}{dp} [\mathcal{T}(p) \cdot \mathcal{T}(p)] \\ &= 2 \mathcal{T}(p) \cdot \frac{d\mathcal{T}(p)}{dp} \quad \Rightarrow \quad \frac{d\mathcal{T}(p)}{dp} \perp \mathcal{T}(p). \end{aligned} \quad (30)$$

Consequently, the vectors  $\mathcal{T}(p)$  and  $\frac{d\mathcal{T}(p)}{dp}$  are **orthogonal** to each other.

**Remark 1:** When  $p = s$  is the arc length along the curve (with  $|\dot{\mathbf{r}}(s)| = 1$ ), we write

$$\frac{d\mathcal{T}(s)}{ds} =: \frac{1}{\rho(s)} \mathcal{N}(s) \quad \begin{cases} |\mathcal{N}(s)| = 1, \\ \rho(s) \geq 0. \end{cases} \quad (31)$$

- $\mathcal{N}(s)$ , the **unit normal vector**, points toward the center of that circle in space which best approximates the curve at the point  $\mathbf{r}(s)$ .
- $\rho(s)$  is the radius of that circle ("**radius of curvature**").

**Remark 2:** Choosing again the time  $t$  as parameter, with  $v(t) = \frac{ds(t)}{dt} \equiv \dot{s}(t)$ , we have

$$\mathbf{v}(t) \equiv \dot{\mathbf{r}}(t) = v(t) \mathcal{T}(s(t)). \quad (32)$$

Taking here the derivative  $\frac{d}{dt}$  (for the second time), we obtain another vector function of time  $t$  which is called the vector  $\mathbf{a}(t)$  of the particle's **acceleration**,

$$\begin{aligned} \mathbf{a}(t) \equiv \ddot{\mathbf{r}}(t) &= \dot{v}(t) \mathcal{T}(s(t)) + v(t) \frac{d}{dt} \mathcal{T}(s(t)) \\ &= \dot{v}(t) \mathcal{T}(s(t)) + v(t) \frac{d}{ds} \mathcal{T}(s) \Big|_{s=s(t)} \underbrace{\dot{s}(t)}_{v(t)} \\ &= \ddot{s}(t) \mathcal{T}(s(t)) + \frac{\dot{s}(t)^2}{\rho(s(t))} \mathcal{N}(s(t)). \end{aligned} \quad (33)$$

We see that  $\mathbf{a}(t)$  has a component  $\ddot{s}(t) \mathcal{T}(s(t))$  **tangential** to the orbit of the moving particle, but, in addition, also has a component  $\frac{\dot{s}(t)^2}{\rho(s(t))} \mathcal{N}(s(t))$  **normal** to the orbit. Consequently, while  $|\mathbf{v}(t)| = |\dot{s}(t)|$ , we have

$$|\mathbf{a}(t)| = \sqrt{\ddot{s}(t)^2 + \left[ \frac{\dot{s}(t)^2}{\rho(s(t))} \right]^2} \geq |\ddot{s}(t)|. \quad (34)$$

### 3 Scalar fields

A **scalar field** is a function  $f$  that assigns a number

$$f(\mathbf{r}) \equiv f(x, y, z) \tag{35}$$

to each point  $\mathbf{r} = (x, y, z)$  in space.

Examples:  $T(x, y, z)$  can be the air temperature (in K) in a room or the temperature distribution in a solid,  $\rho(x, y, z)$  the density (in  $\text{mg}/\text{m}^3$ ) of  $\text{CO}_2$  in air (e.g., above a city). [In addition, these values are typically changing with time  $t$ ,  $T = T(x, y, z, t)$ , etc.]

#### 3.1 Graphical representations

##### 3.1.1 2D fields: 3D plot

An example for a 2D scalar field  $f(x, y)$  is the temperature  $T(x, y)$  on a planar metal surface ( $xy$ -plane). As a model, we consider the simple analytical function

$$T(x, y) = \frac{y}{1 + x^2}. \tag{36}$$

Such a function with only **two** variables can be illustrated by a graph in 3D space: For each point  $(x, y)$  on the  $xy$ -plane consider the number  $z = T(x, y)$  as the third coordinate of a **point**  $(x, y, z)$  in **3D space** above (or, for  $z < 0$ , below) the  $xy$ -plane.

$x$	0	1	2	3	0	1	2	3
$y$	1	1	1	1	2	2	2	2
$z$	1.0	0.5	0.2	0.1	2.0	1.0	0.4	0.2

For a few selected points  $(x, y)$  in the  $xy$ -plane (black dots), the resulting points  $(x, y, z)$  are plotted in Fig. 2 (red dots). The union of all these image points forms a curved surface in 3D space, indicated by blue curves in Fig. 2, called the **3D plot** of  $f$ .

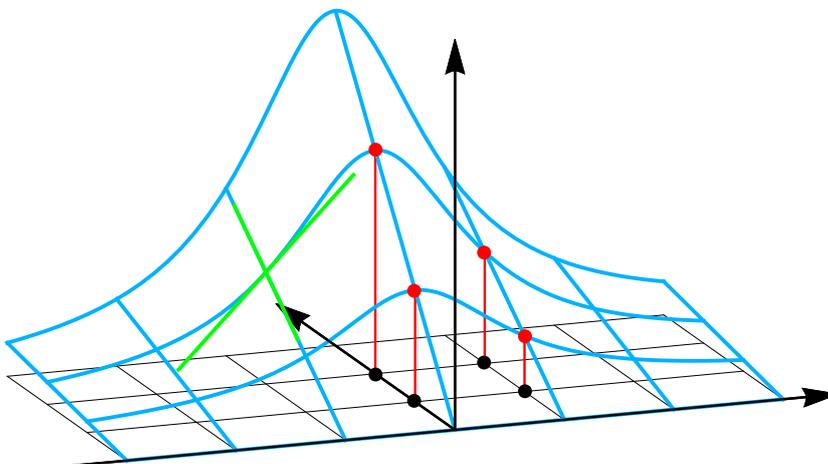


Figure 2: 3D plot of the function  $T(x, y) = \frac{y}{1+x^2}$  (blue solid curves).

### 3.1.2 2D fields: Contour plot

Alternatively, we can plot curves of constant temperature in the  $xy$ -plane: Resolving

$$T(x, y) \equiv \frac{y}{1+x^2} = z \quad (37)$$

for  $y$  (we could resolve for  $x$  as well), we find

$$\begin{aligned} y &= z(1+x^2) \\ &\equiv y_z(x). \end{aligned} \quad (38)$$

For each fixed value of  $z$ , this equation describes a curve in the  $xy$ -plane, joining points with constant temperature  $T(x, y) = z$ . For  $z \in \{0.5, 1.0, 1.5, \dots\}$ , these "isothermals" on the  $xy$ -plane are shown as blue curves in Fig. 3.

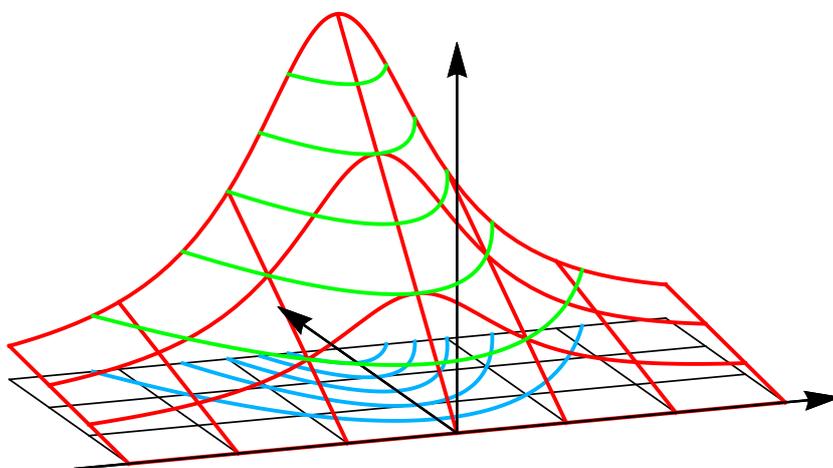


Figure 3: Contour plot of the function  $T(x, y) = \frac{y}{1+x^2}$  (blue solid curves).

Such a set of curves in the  $xy$ -plane is called a **contour plot** of the scalar field  $T(x, y)$ .

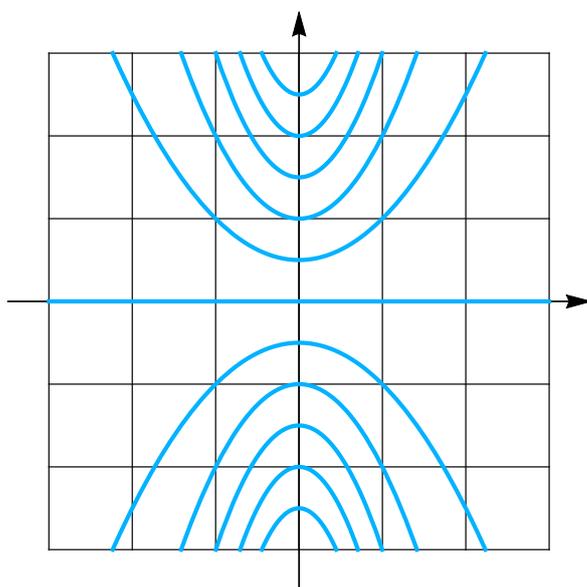


Figure 4: Contour plot of the function  $T(x, y) = \frac{y}{1+x^2}$  in the  $xy$ -plane (blue solid curves).

### 3.2 Partial derivatives

For a given scalar field  $f(x, y, z)$ , the three functions (scalar fields as well!)

$$\begin{aligned} f_x(x, y, z) &\equiv \frac{\partial}{\partial x} f(x, y, z) := \lim_{h \rightarrow 0} \frac{f(x+h, y, z) - f(x, y, z)}{h}, \\ f_y(x, y, z) &\equiv \frac{\partial}{\partial y} f(x, y, z), \\ f_z(x, y, z) &\equiv \frac{\partial}{\partial z} f(x, y, z) \end{aligned} \quad (39)$$

are called its **partial derivatives**. They are evaluated in the same way as the derivative of a function with only **one** variable, treating all the other variables simply as **parameters**.

**Example:** The partial derivatives of  $f(x, y, z) = x^2y \cos(z^2)$  are

$$\frac{\partial f}{\partial x} = 2xy \cos(z^2), \quad \frac{\partial f}{\partial y} = x^2 \cos(z^2), \quad \frac{\partial f}{\partial z} = -2x^2yz \sin(z^2). \quad (40)$$

For a 2D scalar field  $f(x, y)$ , the partial derivatives are visible in the 3D plot of Fig. 2:

- $f_x(x_0, y_0)$  is the slope of the tangent (green) to the curve  $z = f(x, y_0)$  in the  $xz$ -plane,
- $f_y(x_0, y_0)$  is the slope of the tangent (green) to the curve  $z = f(x_0, y)$  in the  $yz$ -plane, both tangents drawn at the point  $(x_0, y_0, z_0)$ , where  $z_0 = f(x_0, y_0)$ .

### 3.3 Gradient

To find an alternative way of illustrating, applicable also to a 3D field  $f(x, y, z)$ , we take the quantities  $\frac{\partial f}{\partial x}$  and  $\frac{\partial f}{\partial y}$  as the coordinates of a column vector

$$\mathbf{G}_f(x, y) = \begin{pmatrix} \frac{\partial}{\partial x} f(x, y) \\ \frac{\partial}{\partial y} f(x, y) \end{pmatrix} \equiv \begin{pmatrix} f_x(x, y) \\ f_y(x, y) \end{pmatrix}. \quad (41)$$

This vector  $\mathbf{G}_f(x, y)$ , actually a vector field, is called the **gradient** of the scalar field  $f$  at the point  $\mathbf{r} = (x, y)$ . Introducing the **nabla operator**  $\nabla$ ,

$$\nabla = \begin{pmatrix} \frac{\partial}{\partial x} \\ \frac{\partial}{\partial y} \end{pmatrix}, \quad (42)$$

the gradient is usually written as

$$\mathbf{G}_f(x, y) = \nabla f(x, y) \quad \Leftrightarrow \quad \mathbf{G}_f(\mathbf{r}) = \nabla f(\mathbf{r}). \quad (43)$$

At the point  $(x_0, y_0)$  on the planar metal surface ( $xy$  plane), the vector

$$\mathbf{G}_T(x_0, y_0) = \nabla T(x, y) \Big|_{x=x_0, y=y_0} \quad (44)$$

points into that direction where the temperature is increasing at the strongest rate. Consequently, it is exactly perpendicular to the corresponding curve in the contour plot.

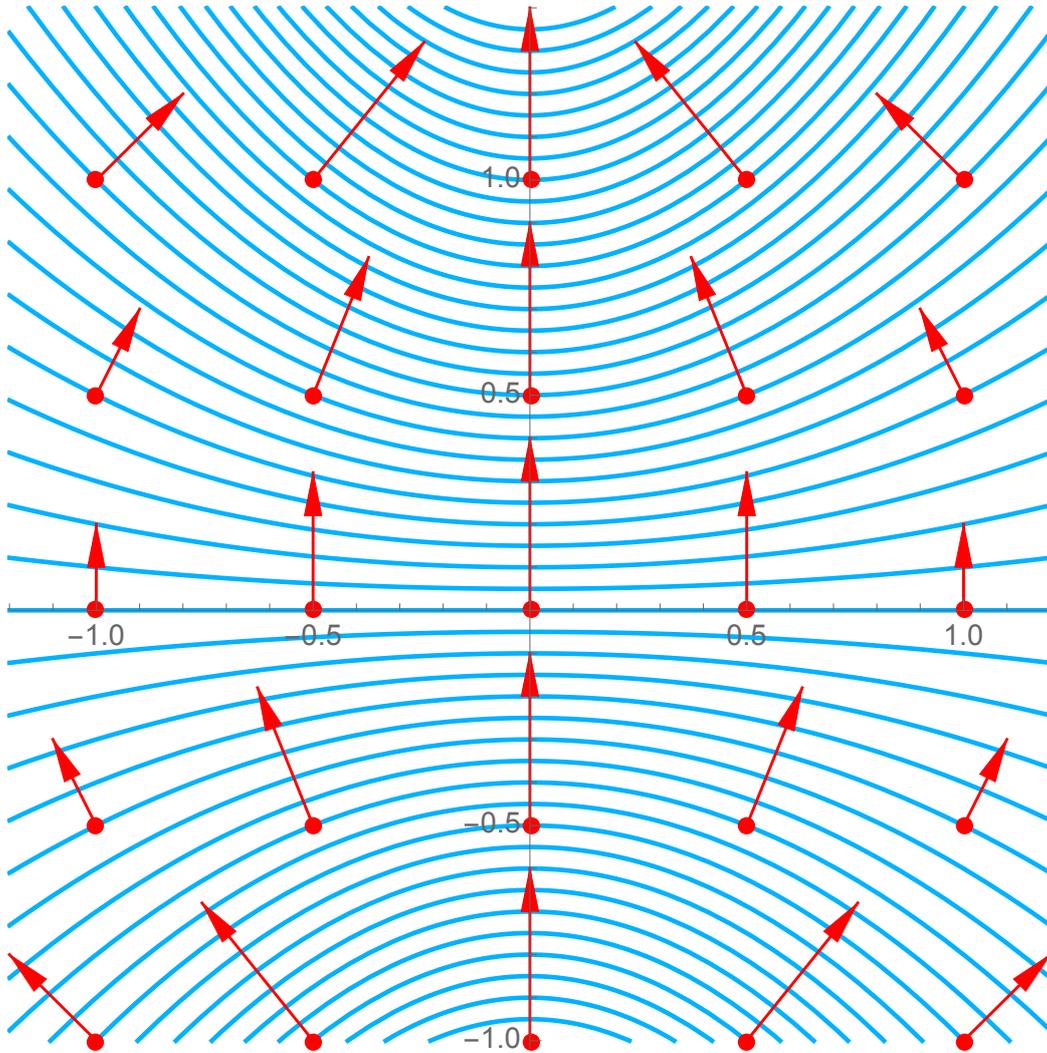


Figure 5: Gradient field of the function  $T(x, y) = \frac{y}{1+x^2}$  (red arrows).

A 3D scalar field  $f(x, y, z)$  such as the air temperature  $T(x, y, z)$ , cannot be represented graphically (since there is no such thing as a “4D plot”!). In this case, the gradient is a 3D column vector,

$$\mathbf{G}_f(x, y, z) := \nabla f(x, y, z) \equiv \begin{pmatrix} f_x(x, y, z) \\ f_y(x, y, z) \\ f_z(x, y, z) \end{pmatrix}, \quad \nabla := \begin{pmatrix} \frac{\partial}{\partial x} \\ \frac{\partial}{\partial y} \\ \frac{\partial}{\partial z} \end{pmatrix}. \quad (45)$$

Still, now in 3D space, this vector points into the direction where the function  $f(x, y, z)$ , e.g., the air temperature  $T(x, y, z)$ , increases at the strongest rate.

## 3.4 Multiple integrals

### 3.4.1 Double or surface integrals

On a water surface ( $xy$ -plane), let  $\rho(x, y)$  be the mass density (in  $\mu\text{g}/\text{cm}^2$ ) of an oil film, representing a 2D scalar field. For simplicity, we assume the function

$$\rho(x, y) = \left\{ \begin{array}{ll} \rho_0 \left(1 - \frac{x^2 + y^2}{R^2}\right) & (x^2 + y^2 \leq R^2) \\ 0 & (x^2 + y^2 \geq R^2) \end{array} \right\}, \quad \rho_0 = 2.7 \frac{\mu\text{g}}{\text{cm}^2}, \quad R = 0.50 \text{ m}. \quad (46)$$

From the function  $\rho(x, y)$ , we wish to compute the amount  $m_\Omega$  of oil (in  $\mu\text{g}$ ) distributed within a given piece  $\Omega$  of water surface, e.g., the right triangle (red in Fig. 6)

$$\Omega = \left\{ (x, y) \in \mathbb{R}^2 \mid 0 \leq x \leq a, \quad 0 \leq y \leq \frac{bx}{a} \right\} \quad (a, b > 0, \quad a^2 + b^2 < R^2). \quad (47)$$

For an approximate solution to this problem, we introduce small step sizes  $\Delta x$  and  $\Delta y$ , generating a mesh of discrete points  $\mathbf{r}_{ij}$  in the  $xy$ -plane (black dots in Fig. 6),

$$\mathbf{r}_{ij} = (x_i, y_j) = (i \cdot \Delta x, j \cdot \Delta y) \quad (i, j \in \mathbb{Z}). \quad (48)$$

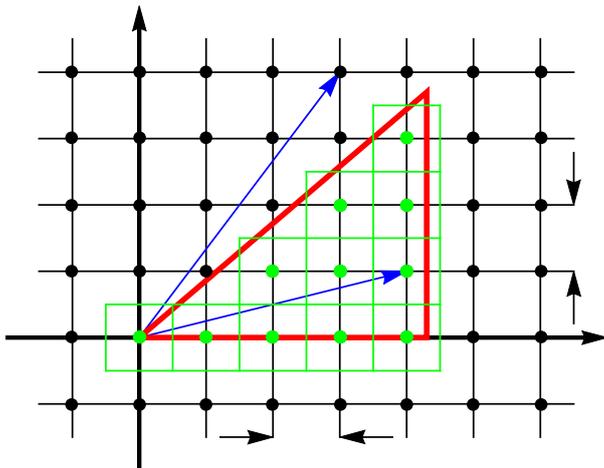


Figure 6: The right triangle  $\Omega$  (red) and the position vectors  $\mathbf{r}_{41}$  and  $\mathbf{r}_{34}$  (blue).

For sufficiently small  $\Delta x$  and  $\Delta y$ , the amount  $\Delta m_{ij}$  of oil within the small rectangle with side lengths  $\Delta x$  and  $\Delta y$ , centered at the point  $\mathbf{r}_{ij}$ , is approximately given by

$$\Delta m_{ij} \approx \Delta x \cdot \Delta y \cdot \rho(x_i, y_j). \quad (49)$$

Considering all points  $\mathbf{r}_{ij} \in \Omega$  (green in Fig. 6) and taking the sum, we have

$$m_\Omega \approx \sum_{\mathbf{r}_{ij} \in \Omega} \Delta m_{ij} = \sum_{i=i_{\min}}^{i_{\max}} \Delta x \left[ \sum_{j=j_{\min}(i)}^{j_{\max}(i)} \Delta y \cdot \rho(x_i, y_j) \right]. \quad (50)$$

For the situation in Fig. 6, e.g., the outer sum has the fixed limits  $i_{\min} = 0$  and  $i_{\max} = 4$ , while the limits  $j_{\min}$  and  $j_{\max}$  of the inner sum are depending on the current value of  $i$ :

$$j_{\min}(i) = 0 \quad (\forall i), \quad j_{\max}(0) = j_{\max}(1) = 0, \quad j_{\max}(2) = 1, \quad j_{\max}(3) = 2, \quad j_{\max}(4) = 3.$$

At this point, it is easy to see<sup>1</sup> that the limit  $\Delta x, \Delta y \rightarrow 0$  yields the **double integral**

$$m_\Omega = \int_{x_{\min}}^{x_{\max}} dx \int_{y_{\min}(x)}^{y_{\max}(x)} dy \rho(x, y), \quad (58)$$

where  $x_{\min}$  and  $x_{\max}$  are the extreme values of the  $x$ -coordinates of all points  $(x, y)$  in  $\Omega$ , while  $y_{\min}(x)$  and  $y_{\max}(x)$  are, for a given value of  $x$ , the ones of the  $y$ -coordinates,

$$m_\Omega = \int_0^a dx \int_0^{bx/a} dy \rho(x, y). \quad (59)$$

---

<sup>1</sup>Fixed by the boundary of the domain  $\Omega \in \mathbb{R}^2$ , the numbers  $i_{\min}$  and  $i_{\max}$  depend on  $\Delta x$  and, for each  $i$ , the numbers  $j_{\min}(i)$  and  $j_{\max}(i)$  depend on  $\Delta y$ , with the limits

$$\lim_{\Delta x \rightarrow 0} [i_{\min} \cdot \Delta x] = x_{\min}, \quad \lim_{\Delta x \rightarrow 0} [i_{\max} \cdot \Delta x] = x_{\max}; \quad (51)$$

$$\lim_{\Delta y \rightarrow 0} [j_{\min}(i) \cdot \Delta y] = y_{\min}(x_i), \quad \lim_{\Delta y \rightarrow 0} [j_{\max}(i) \cdot \Delta y] = y_{\max}(x_i), \quad (52)$$

where  $x_{\min}/x_{\max}$  are the smallest/largest  $x$ -coordinates of any  $\mathbf{r} \in \Omega$ , while  $y_{\min}(x)/y_{\max}(x)$  are the smallest/largest  $y$ -coordinates of any  $\mathbf{r} \in \Omega$  with given  $x$ -coordinate  $x$ .

E.g., for the triangle  $\Omega$  in Fig. 6/Eq. (47), we have

$$x_{\min} = 0, \quad x_{\max} = a, \quad y_{\min}(x) \equiv 0, \quad y_{\max}(x) = \frac{bx}{a}. \quad (53)$$

The approximation (50) becomes exact in the limit  $\Delta x, \Delta y \rightarrow 0$ ,

$$m_\Omega = \lim_{\Delta x \rightarrow 0} \sum_{i=i_{\min}}^{i_{\max}} \Delta x \left[ \lim_{\Delta y \rightarrow 0} \sum_{j=j_{\min}(i)}^{j_{\max}(i)} \Delta y \rho(x_i, y_j) \right]. \quad (54)$$

Obviously, for any fixed value of  $x_i$ , the limit [...] in squared brackets is the ordinary integral of the function  $f(y) \equiv \rho(x_i, y)$ ,

$$\lim_{\Delta y \rightarrow 0} \sum_{j=j_{\min}(i)}^{j_{\max}(i)} \Delta y f(y_j) = \int_{y_{\min}(x_i)}^{y_{\max}(x_i)} dy f(y). \quad (55)$$

It can be evaluated in the familiar way,

$$\begin{aligned} \int_{y_{\min}(x_i)}^{y_{\max}(x_i)} dy \rho(x_i, y) &= \rho_0 \int_0^{bx_i/a} dy (A - By^2) \quad \left( A = 1 - \frac{x_i^2}{R^2}, \quad B = \frac{1}{R^2} \right) \\ &= \rho_0 \left[ Ay - \frac{B}{3} y^3 \right]_{y=0}^{y=bx_i/a} \\ &= \rho_0 \left[ \frac{b}{a} x_i - \frac{b}{aR^2} \left( 1 + \frac{b^2}{3a^2} \right) x_i^3 \right] \equiv F(x_i). \end{aligned} \quad (56)$$

In a second step, the limit  $\Delta x \rightarrow 0$  in Eq. (54) yields the ordinary integral of  $F(x)$ ,

$$\begin{aligned} m_\Omega &= \lim_{\Delta x \rightarrow 0} \sum_{i=i_{\min}}^{i_{\max}} \Delta x F(x_i) \\ &= \int_{x_{\min}}^{x_{\max}} dx F(x) \\ &= \int_0^a dx F(x) = \rho_0 \left[ \frac{ab}{2} - \frac{ab}{12R^2} (3a^2 + b^2) \right] = \rho_0 \frac{ab}{2} \left[ 1 - \frac{3a^2 + b^2}{6R^2} \right]. \end{aligned} \quad (57)$$

These two subsequent integrations are usually written as the double integral of Eq. (59).

In evaluating a double integral, we first perform the “inner“ (here:  $y$ ) integration (note that the upper limit  $bx/a$  of the  $y$  integral depends on the variable  $x$  of the ”outer“ integral!), treating the other variable  $x$  as a **constant parameter**,

$$\begin{aligned}
m_\Omega &= \int_0^a dx \int_0^{bx/a} dy \rho_0 \left(1 - \frac{x^2 + y^2}{R^2}\right) \\
&= \int_0^a dx \int_0^{bx/a} dy \frac{\rho_0}{R^2} \left(R^2 - (x^2 + y^2)\right) \\
&= \frac{\rho_0}{R^2} \int_0^a dx \int_0^{bx/a} dy (R^2 - x^2 - y^2) \\
&= \frac{\rho_0}{R^2} \int_0^a dx \left[ (R^2 - x^2)y - \frac{y^3}{3} \right]_{y=0}^{y=bx/a} \\
&= \frac{\rho_0}{R^2} \int_0^a dx \left[ \frac{bR^2}{a}x - \left(\frac{b}{a} + \frac{b^3}{3a^3}\right)x^3 \right]. \tag{60}
\end{aligned}$$

Remaining is an ordinary integral over **one** variable  $x$ ,

$$m_\Omega = \frac{\rho_0}{R^2} \left[ \frac{bR^2}{a} \frac{x^2}{2} - \left(\frac{b}{a} + \frac{b^3}{3a^3}\right) \frac{x^4}{4} \right]_0^a = \rho_0 \frac{ab}{2} \left[ 1 - \frac{3a^2 + b^2}{6R^2} \right]. \tag{61}$$

**Remark:** It is also possible, to take the  $y$  integral as the outer one. Then, however, the integration limits must be re-defined,

$$\begin{aligned}
m_\Omega &= \frac{\rho_0}{R^2} \int_{y_{\min}}^{y_{\max}} dy \int_{x_{\min}(y)}^{x_{\max}(y)} dx (R^2 - y^2 - x^2) \\
&= \frac{\rho_0}{R^2} \int_0^b dy \int_{ay/b}^a dx (R^2 - y^2 - x^2) \\
&= \frac{\rho_0}{R^2} \int_0^b dy \left[ (R^2 - y^2)x - \frac{x^3}{3} \right]_{x=ay/b}^{x=a} \\
&= \frac{\rho_0}{R^2} \int_0^b dy \left[ (R^2 - y^2) \left(a - \frac{a}{b}y\right) - \frac{1}{3} \left(a^3 - \frac{a^3}{b^3}y^3\right) \right] \\
&= \frac{\rho_0}{R^2} \int_0^b dy \left[ \left(aR^2 - \frac{a^3}{3}\right) - \frac{aR^2}{b}y - ay^2 + \left(\frac{a}{b} + \frac{a^3}{3b^3}\right)y^3 \right] \tag{62}
\end{aligned}$$

Again, we are left with an ordinary integral over **one** variable  $y$ ,

$$\begin{aligned}
m_\Omega &= \frac{\rho_0}{R^2} \left[ \left(aR^2 - \frac{a^3}{3}\right)b - \frac{abR^2}{2} - \frac{ab^3}{3} + \left(\frac{ab^3}{4} + \frac{a^3b}{12}\right) \right] \\
&= \rho_0 \left[ \frac{ab}{2} - \frac{ab}{12R^2}(3a^2 + b^2) \right]. \tag{63}
\end{aligned}$$

The result, of course, cannot depend on the particular way chosen for evaluating the integrals. It is completely fixed by the domain  $\Omega \subseteq \mathbb{R}^2$  and by the integrand (scalar field)  $\rho(x, y) \equiv \rho(\mathbf{r})$ . Therefore, we use the notation

$$m_\Omega = \int_\Omega d^2r \rho(\mathbf{r}), \tag{64}$$

and call this the **double integral** of the scalar field  $\rho(\mathbf{r})$  over the domain  $\Omega \subseteq \mathbb{R}^2$ .

**Example:** We consider the **circle**  $\Omega$  with radius  $a$  and center at  $\mathbf{r} = \mathbf{0}$ ,

$$\Omega = \left\{ (x, y) \in \mathbb{R}^2 \mid x^2 + y^2 \leq a^2 \right\}. \quad (65)$$

In this case,  $x_{\min}, x_{\max} = \mp a$  and  $y_{\min}(x), y_{\max}(x) = \mp \sqrt{a^2 - x^2}$ ,

$$\begin{aligned} \int_{\Omega} d^2r \rho(\mathbf{r}) &= \frac{\rho_0}{R^2} \int_{-a}^a dx \int_{-\sqrt{a^2-x^2}}^{\sqrt{a^2-x^2}} dy (R^2 - x^2 - y^2) \\ &= \frac{\rho_0}{R^2} \int_{-a}^a dx \left[ (R^2 - x^2)y - \frac{y^3}{3} \right]_{y=-\sqrt{a^2-x^2}}^{y=\sqrt{a^2-x^2}} \\ &= \frac{\rho_0}{R^2} \int_{-a}^a dx \left[ (R^2 - x^2)2\sqrt{a^2 - x^2} - \frac{2}{3}(a^2 - x^2)^{3/2} \right] \\ &= \frac{\rho_0}{R^2} \int_{-a}^a dx \left[ 2\left(R^2 - \frac{a^2}{3}\right) - \frac{4}{3}x^2 \right] \sqrt{a^2 - x^2}. \end{aligned} \quad (66)$$

From standard integral tables (e.g.: Bronstein, integrals nos. 157 and 159), we extract

$$\int dx \sqrt{a^2 - x^2} = \frac{1}{2} \left( x\sqrt{a^2 - x^2} + a^2 \arcsin \frac{x}{a} \right), \quad (67)$$

$$\int dx x^2 \sqrt{a^2 - x^2} = \frac{a^2}{8} \left( x\sqrt{a^2 - x^2} + a^2 \arcsin \frac{x}{a} \right) - \frac{x}{4} (a^2 - x^2)^{3/2}, \quad (68)$$

to find

$$\int_{\Omega} d^2r \rho(\mathbf{r}) = \rho_0 \left( \pi a^2 - \frac{\pi a^4}{2R^2} \right). \quad (69)$$

This derivation is simplified considerably by using polar coordinates, see Eq. (94) below.

### 3.4.2 Triple or volume integrals

Integrals of scalar fields  $f(\mathbf{r})$  over a 3D domain  $\Omega$  are called **triple** or **volume integrals**,

$$\int_{\Omega} d^3r f(\mathbf{r}) = \int_{x_{\min}}^{x_{\max}} dx \int_{y_{\min}(x)}^{y_{\max}(x)} dy \int_{z_{\min}(x,y)}^{z_{\max}(x,y)} dz f(x, y, z). \quad (70)$$

Again, the order of these three integrations may be chosen differently, when the limits of each integral are re-defined properly.

**Example:** Let  $\Omega$  be a pyramid with height  $H$  and quadratic base with side  $a$ ,

$$\Omega = \left\{ (x, y, z) \mid 0 \leq z \leq H, \quad 0 \leq |x|, |y| \leq \frac{a}{2} \left( 1 - \frac{z}{H} \right) \right\}. \quad (71)$$

In this case, we choose the  $z$  integral as the outer one to obtain

$$\begin{aligned} \int_{\Omega} d^3r f(\mathbf{r}) &= \int_{z_{\min}}^{z_{\max}} dz \int_{x_{\min}(z)}^{x_{\max}(z)} dx \int_{y_{\min}(z)}^{y_{\max}(z)} dy f(x, y, z) \\ &= \int_0^H dz \int_{-\frac{a}{2}(1-\frac{z}{H})}^{\frac{a}{2}(1-\frac{z}{H})} dx \int_{-\frac{a}{2}(1-\frac{z}{H})}^{\frac{a}{2}(1-\frac{z}{H})} dy f(x, y, z). \end{aligned} \quad (72)$$

For the simple function  $f(x, y, z) = az$ , this integral will be evaluated in Eq. (78) below.

### 3.4.3 Average value of a function

A solid rod, lying on the  $x$ -axis (with its endpoints at  $x = a$  and at  $x = b$ ), may have a non-uniform ( $x$ -dependent) temperature distribution, given by a function  $T(x)$ .

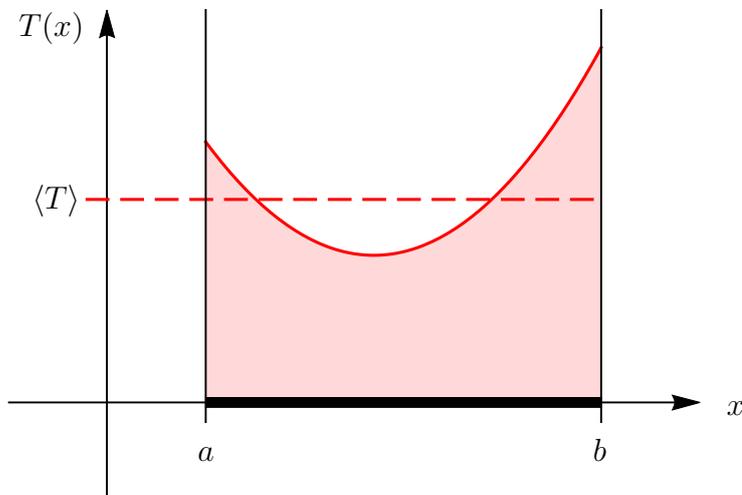


Figure 7: The rod on the  $x$ -axis (black), its non-uniform temperature distribution (red curve), and its average temperature (dashed red line).

From the figure, it is obvious that the **average temperature** of the rod is given by

$$\langle T \rangle = \frac{1}{L} \int_a^b dx T(x), \quad (73)$$

where  $L = b - a$  is the length of the rod.

Instead of the rod, covering the 1D interval  $\Omega = [a, b] \subset \mathbb{R}$  with **length**  $L_\Omega = b - a$ , we may also consider a thin metal sheet, covering a 2D surface  $\Omega \subset \mathbb{R}^2$  with **area**  $A_\Omega$ , or a piece of solid material, covering a 3D region  $\Omega \subset \mathbb{R}^3$  with **volume**  $V_\Omega$ .

Area  $A_\Omega$  and length  $L_\Omega$  can each be called a "volume"  $V_\Omega$  in  $D$ -dimensional space (with  $D = 2$  and  $D = 1$ , respectively). In this spirit, we generally define:

**Def.:** The **average value** of a scalar field  $T(\mathbf{r})$  in a  $D$ -dimensional domain  $\Omega \subset \mathbb{R}^D$  is

$$\langle T(\mathbf{r}) \rangle_{\mathbf{r} \in \Omega} \equiv \frac{1}{V_\Omega} \int_\Omega d^D r T(\mathbf{r}), \quad (74)$$

where  $V_\Omega$  is the  $D$ -dimensional "volume" of  $\Omega$  (with  $D = 3$ ,  $D = 2$  or  $D = 1$ ).

**Example 1:** Consider a metal sheet covering the "upper half"  $\Omega$  of a disk,

$$\Omega = \left\{ (x|y) \in \mathbb{R}^2 \mid x^2 + y^2 \leq a^2, y \geq 0 \right\} \quad (V_\Omega \equiv A_\Omega = \frac{\pi}{2} a^2). \quad (75)$$

Given the function  $T(x, y) = T_0 \cdot \frac{y}{a}$ , the average temperature on the sheet is

$$\begin{aligned} \langle T(\mathbf{r}) \rangle_{\mathbf{r} \in \Omega} &= \frac{1}{\frac{\pi}{2} a^2} \int_{-a}^a dx \int_0^{\sqrt{a^2 - x^2}} dy T_0 \cdot \frac{y}{a} \\ &= \frac{2}{\pi a^2} \frac{T_0}{a} \int_{-a}^a dx \left[ \frac{y^2}{2} \right]_0^{\sqrt{a^2 - x^2}} = \frac{4}{3\pi} T_0. \end{aligned} \quad (76)$$

**Example 2:** Inside the pyramid  $\Omega$  of Eq. (71) the temperature shall increase linearly from  $0^\circ\text{C}$  at the floor ( $z = 0$ ) to some maximum value  $T_0$  at the tip ( $z = H$ ),

$$T(\mathbf{r}) = \frac{T_0}{H} z \equiv T(x, y, z). \quad (77)$$

Given the volume  $V_\Omega = \frac{1}{3}a^2H$  of this pyramid, the **average temperature** inside is

$$\begin{aligned} \langle T(\mathbf{r}) \rangle_{\mathbf{r} \in \Omega} &\equiv \frac{1}{V_\Omega} \int_\Omega d^3r T(\mathbf{r}) \\ &= \frac{3}{a^2H} \int_0^H dz \int_{-\frac{a}{2}(1-\frac{z}{H})}^{\frac{a}{2}(1-\frac{z}{H})} dx \int_{-\frac{a}{2}(1-\frac{z}{H})}^{\frac{a}{2}(1-\frac{z}{H})} dy \frac{T_0}{H} z \\ &= \frac{3T_0}{a^2H^2} \int_0^H dz \int_{-\frac{a}{2}(1-\frac{z}{H})}^{\frac{a}{2}(1-\frac{z}{H})} dx \left[ zy \right]_{y=-\frac{a}{2}(1-\frac{z}{H})}^{y=\frac{a}{2}(1-\frac{z}{H})} \\ &= \frac{3T_0}{a^2H^2} \int_0^H dz \int_{-\frac{a}{2}(1-\frac{z}{H})}^{\frac{a}{2}(1-\frac{z}{H})} dx za \left( 1 - \frac{z}{H} \right) \\ &= \frac{3T_0}{a^2H^2} \int_0^H dz \left[ za \left( 1 - \frac{z}{H} \right) x \right]_{x=-\frac{a}{2}(1-\frac{z}{H})}^{x=\frac{a}{2}(1-\frac{z}{H})} \\ &= \frac{3T_0}{a^2H^2} \int_0^H dz za^2 \left( 1 - \frac{z}{H} \right)^2 \\ &= \frac{3T_0}{H^2} \int_0^H dz \left( \frac{z^3}{H^2} - \frac{2z^2}{H} + z \right) = \frac{T_0}{4}. \end{aligned} \quad (78)$$

**Remark A: Computing a volume (in 3D or in 2D).**

For the trivial function  $T(\mathbf{r}) \equiv 1$ , we have  $\langle T(\mathbf{r}) \rangle_\Omega = 1$ , implying the formula

$$V_\Omega = \int_\Omega d^D r 1. \quad (79)$$

**Example 3:** Find the volume  $V_\Omega$  of the tetrahedron  $\Omega$ , bounded by the four planes  $x = 0$ ,  $y = 0$ ,  $z = 0$  and  $\frac{x}{a} + \frac{y}{b} + \frac{z}{c} = 1$  (with  $a, b, c > 0$ ; SKETCH!),

$$\begin{aligned} V_\Omega &= \int_\Omega d^3r 1 = \int_0^a dx \int_0^{b(1-\frac{x}{a})} dy \int_0^{c(1-\frac{x}{a}-\frac{y}{b})} dz 1 \\ &= c \int_0^a dx \int_0^{b(1-\frac{x}{a})} dy \left( 1 - \frac{x}{a} - \frac{y}{b} \right) \\ &= c \int_0^a dx \left[ \left( 1 - \frac{x}{a} \right) y - \frac{y^2}{2b} \right]_{y=0}^{y=b(1-\frac{x}{a})} \\ &= bc \int_0^a dx \left[ \left( 1 - \frac{x}{a} \right)^2 - \frac{1}{2} \left( 1 - \frac{x}{a} \right)^2 \right] \\ &= \frac{bc}{2} \int_0^a dx \left( 1 - \frac{x}{a} \right)^2 = -\frac{bc}{2} \int_1^0 adu u^2 = \frac{abc}{6}. \end{aligned} \quad (80)$$

### Remark B: Double integral as a 3D volume

Let  $\Sigma$  be a domain in the  $xy$ -plane, fixed by  $x_{\min}$ ,  $x_{\max}$ ,  $y_{\min}(x)$ , and  $y_{\max}(x)$ . Then, the double integral of a 2D scalar field  $f(\mathbf{r}) = f(x, y)$  is

$$\begin{aligned} \int_{\Sigma} d^2r f(\mathbf{r}) &= \int_{x_{\min}}^{x_{\max}} dx \int_{y_{\min}(x)}^{y_{\max}(x)} dy f(x, y) \\ &\equiv \int_{x_{\min}}^{x_{\max}} dx \int_{y_{\min}(x)}^{y_{\max}(x)} dy \int_0^{f(x,y)} dz 1, \end{aligned} \quad (81)$$

where we have utilized the trivial identity  $f(x, y) = \int_0^{f(x,y)} dz 1$ . Consequently, provided that  $f(\mathbf{r}) \geq 0$  for  $\mathbf{r} \in \Sigma$ , the double integral  $\int_{\Sigma} d^2r f(\mathbf{r})$  equals the **volume** of the 3D region  $\Omega$  bounded by  $\Sigma$  and the (curved) 3D-plot of  $f$ ,

$$\int_{\Sigma} d^2r f(\mathbf{r}) = \int_{\Omega} d^3r 1. \quad (82)$$

**Example 4:** The tetrahedral volume from example 2 can also be obtained as the double integral of the 2D scalarfield  $f(x, y) = c(1 - \frac{x}{a} - \frac{y}{b}) \equiv z$  over  $\Sigma = \{(x, y) \in \mathbb{R}^2 \mid \dots\}$ ,

$$V_{\Omega} = \int_{\Sigma} d^2r f(\mathbf{r}) = c \int_0^a dx \int_0^{b(1-x/a)} dy \left(1 - \frac{x}{a} - \frac{y}{b}\right) = \dots = \frac{abc}{6}. \quad (83)$$

## 4 Curvilinear coordinates

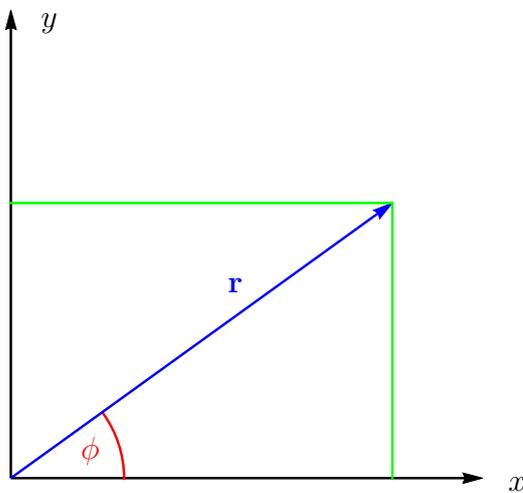
### 4.1 Planar polar coordinates

The position vector  $\mathbf{r}$  of a point  $P$  in the  $xy$ -plane with **cartesian coordinates**  $(x, y)$ ,

$$\mathbf{r} = x\mathbf{u}_1 + y\mathbf{u}_2 = \begin{pmatrix} x \\ y \end{pmatrix}, \quad (84)$$

is fixed as well by its **planar polar coordinates**  $(r, \phi)$ :

$r =  \mathbf{r} $	$0 \leq r < +\infty$	<b>Length of <math>\mathbf{r}</math></b>	Distance of $P$ from the origin $O$ .
$\phi$	$0 \leq \phi < 2\pi$	<b>Azimuth angle of <math>\mathbf{r}</math></b>	Angle from the $x$ -axis to $\mathbf{r}$ (ccw)



#### 4.1.1 Coordinate lines

We express the “old” coordinates  $(x, y)$  in terms of the “new” ones  $(r, \phi)$ ,

$$\mathbf{r}(r, \phi) \equiv \begin{pmatrix} x(r, \phi) \\ y(r, \phi) \end{pmatrix} = \begin{pmatrix} r \cos \phi \\ r \sin \phi \end{pmatrix}. \quad (85)$$

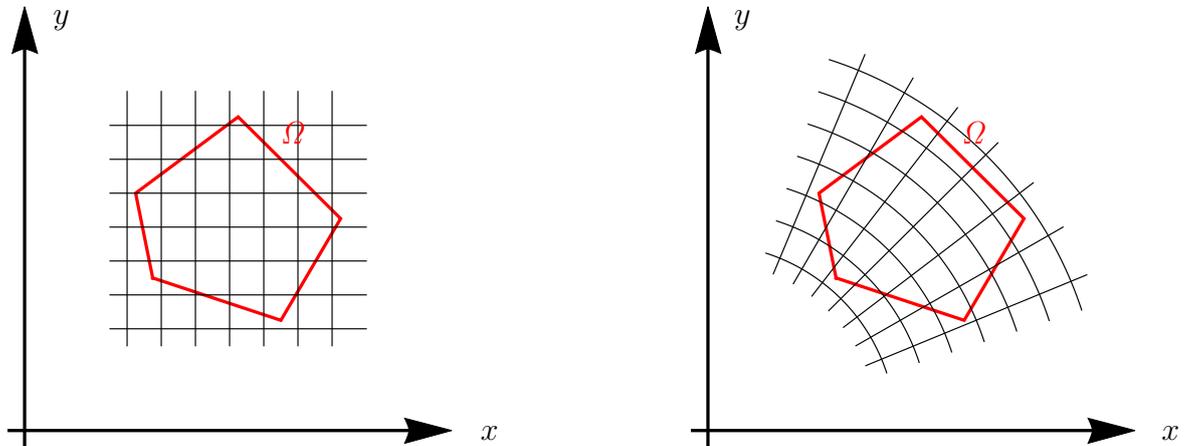
This is a vector function of **two** variables ( $r$  and  $\phi$ ).

- For fixed  $\phi = \phi_0$  and  $r$  growing from  $r = 0$  to  $r = \infty$ , the point  $\mathbf{r}(r, \phi_0)$  in the  $xy$ -plane moves on a **ray** to infinity, starting at the origin.
- For fixed  $r = r_0$  and  $\phi$  growing from  $\phi = 0$  to  $\phi = 2\pi$ , the point  $\mathbf{r}(r_0, \phi)$  moves ccw on a **circle** of radius  $r_0$  centered at the origin.

These rays and circles are the **coordinate lines** of planar polar coordinates.

### 4.1.2 Double integrals

As in section 3.4.1, let  $\rho(x, y)$  be the 2D mass density of a thin oil film on a water surface ( $xy$ -plane  $\mathbb{R}^2$ ). To compute the amount  $m_\Omega$  of oil distributed within a given piece  $\Omega$  of water surface (red pentagon in the figures below), we introduce a grid of coordinate lines.



**Cartesian** coordinate lines (left part of the figure) cut the  $xy$ -plane into small rectangles ("tiles") with side lengths  $\Delta x$ ,  $\Delta y$ , and with the uniform surface area (per tile)

$$\Delta A = \Delta x \cdot \Delta y. \quad (86)$$

By proper choice of the coordinate lines, these cartesian tiles have their midpoints at

$$\mathbf{r}_{ij} = \begin{pmatrix} x_i \\ y_j \end{pmatrix}, \quad (87)$$

where  $x_i = i \cdot \Delta x$  and  $y_j = j \cdot \Delta y$  (with  $i, j \in \mathbb{Z}$ ). Consequently, the amount  $m_{ij}$  of oil distributed within the tile at  $\mathbf{r}_{ij}$  is approximately given by  $\rho(\mathbf{r}_{ij}) \cdot \Delta A$ ,

$$m_{ij} \approx \rho(\mathbf{r}_{ij}) \cdot \Delta x \cdot \Delta y, \quad (88)$$

which becomes exact in the limit  $\Delta x, \Delta y \rightarrow 0$ .

**Polar** coordinate lines (right part of the figure) cut the  $xy$ -plane into curved tiles which, spanning a small angle  $\Delta\phi$ , are sections of concentric rings with thickness  $\Delta r$ . For small  $\Delta\phi$ , such a curved tile, with outer and inner radii  $r \pm \frac{1}{2}\Delta r$ , is approximately rectangular, with side lengths  $\Delta r$  and  $r\Delta\phi$ . Consequently, the surface area per tile is  $r$ -dependent,

$$\Delta\tilde{A}(r) \approx r \cdot \Delta r \cdot \Delta\phi. \quad (89)$$

In terms of  $x(r, \phi)$  and  $y(r, \phi)$  from Eq. (85), these tiles have their midpoints at

$$\mathbf{s}_{\alpha\beta} = \begin{pmatrix} x(r_\alpha, \phi_\beta) \\ y(r_\alpha, \phi_\beta) \end{pmatrix}, \quad (90)$$

where  $r_\alpha = \alpha \cdot \Delta r$  and  $\phi_\beta = \beta \cdot \Delta\phi$  (with  $\alpha, \beta = 1, 2, 3, \dots$ ). Consequently, the amount  $\tilde{m}_{\alpha\beta}$  of oil distributed within the tile at  $\mathbf{s}_{\alpha\beta}$  is approximately given by  $\rho(\mathbf{s}_{\alpha\beta}) \cdot \Delta\tilde{A}(r_\alpha)$ ,

$$\tilde{m}_{\alpha\beta} \approx \rho(\mathbf{s}_{\alpha\beta}) r_\alpha \cdot \Delta r \cdot \Delta\phi, \quad (91)$$

which becomes exact in the limit  $\Delta r, \Delta\phi \rightarrow 0$ .

Eq. (88) for **cartesian** coordinates implies [cf. Eq. (50) in section 3.4.1]

$$\begin{aligned}
m_\Omega &\approx \sum_{\mathbf{r}_{ij} \in \Omega} m_{ij} \approx \sum_{i=i_{\min}}^{i_{\max}} \Delta x \sum_{j=j_{\min}(i)}^{j_{\max}(i)} \Delta y \cdot \rho(x_i, y_j) \\
&\rightarrow \int_{x_{\min}}^{x_{\max}} dx \int_{y_{\min}(x)}^{y_{\max}(x)} dy \rho(x, y).
\end{aligned} \tag{92}$$

Similarly, writing  $\rho(\mathbf{s}_{\alpha\beta}) \equiv \rho(x(r_\alpha, \phi_\beta), y(r_\alpha, \phi_\beta)) = \tilde{\rho}(r_\alpha, \phi_\beta)$ , Eq. (91) yields

$$\begin{aligned}
m_\Omega &\approx \sum_{\mathbf{s}_{\alpha\beta} \in \Omega} \tilde{m}_{\alpha\beta} = \sum_{\alpha=\alpha_{\min}}^{\alpha_{\max}} \Delta r \sum_{\beta=\beta_{\min}(\alpha)}^{\beta_{\max}(\alpha)} \Delta \phi \cdot r_\alpha \cdot \tilde{\rho}(r_\alpha, \phi_\beta) \\
&\rightarrow \int_{r_{\min}}^{r_{\max}} dr \int_{\phi_{\min}(r)}^{\phi_{\max}(r)} d\phi r \tilde{\rho}(r, \phi).
\end{aligned} \tag{93}$$

The extra factor  $J(r, \phi) = r$  in the integrand is the **Jacobian** of planar **polar** coordinates.

**Example 1:** For the disk  $\Omega$  with radius  $a$  of Eq. (65), Eq. (66) now becomes

$$\begin{aligned}
m_\Omega &= \int_\Omega d^2r \rho(\mathbf{r}) = \int_{-a}^a dx \int_{-\sqrt{a^2-x^2}}^{\sqrt{a^2-x^2}} dy \rho(x, y), \quad \rho(x, y) = \frac{\rho_0}{R^2} [R^2 - (x^2 + y^2)] \\
&= \int_0^a dr \int_0^{2\pi} d\phi r \tilde{\rho}(r, \phi) \\
&= \frac{\rho_0}{R^2} \int_0^a dr \int_0^{2\pi} d\phi r (R^2 - r^2) \\
&= \frac{\rho_0}{R^2} \int_0^a dr [(R^2 r - r^3) \phi]_{\phi=0}^{\phi=2\pi} \\
&= \frac{\rho_0}{R^2} \int_0^a dr 2\pi (R^2 r - r^3) \\
&= \frac{\rho_0}{R^2} 2\pi \left[ R^2 \frac{r^2}{2} - \frac{r^4}{4} \right]_{r=0}^{r=a} = \rho_0 \left( \pi a^2 - \frac{\pi a^4}{2R^2} \right).
\end{aligned} \tag{94}$$

**Example 2:** If  $\Omega$ , more generally, is a **section of a circular ring**, we have

$$\int_\Omega d^2r f(\mathbf{r}) = \int_{r_1}^{r_2} dr \int_{\phi_1}^{\phi_2} d\phi r \tilde{f}(r, \phi). \tag{95}$$

**Remark:** For general curvilinear coordinates  $(v, w)$ , the surface area  $\Delta\tilde{A}$  of the curved tile depends on both  $v$  and  $w$ . The corresponding **Jacobian**  $J(v, w)$  is defined by

$$J(v, w) = \lim_{\Delta v, \Delta w \rightarrow 0} \frac{\Delta\tilde{A}(v, w)}{\Delta v \cdot \Delta w}. \tag{96}$$

For small but **finite**  $\Delta v$  and  $\Delta w$ , we then have  $\Delta\tilde{A}(v, w) \approx J(v, w) \cdot \Delta v \cdot \Delta w$  and

$$m_\Omega = \int_{v_{\min}}^{v_{\max}} dv \int_{w_{\min}(v)}^{w_{\max}(v)} dw J(v, w) \tilde{\rho}(r, \phi). \tag{97}$$

Generally, the Jacobian is the determinant of the **functional matrix**,

$$J(v, w) = \det \begin{pmatrix} \partial x(v, w)/\partial v & \partial x(v, w)/\partial w \\ \partial y(v, w)/\partial v & \partial y(v, w)/\partial w \end{pmatrix}. \tag{98}$$

### 4.1.3 Scale factors

Consider the partial derivatives of the two-argument vector function  $\mathbf{r}(r, \phi)$ ,

$$\begin{aligned}\mathbf{h}_r(r, \phi) &= \frac{\partial \mathbf{r}(r, \phi)}{\partial r} = \frac{\partial}{\partial r} \begin{pmatrix} r \cos \phi \\ r \sin \phi \end{pmatrix} = \begin{pmatrix} \cos \phi \\ \sin \phi \end{pmatrix}, \\ \mathbf{h}_\phi(r, \phi) &= \frac{\partial \mathbf{r}(r, \phi)}{\partial \phi} = \frac{\partial}{\partial \phi} \begin{pmatrix} r \cos \phi \\ r \sin \phi \end{pmatrix} = \begin{pmatrix} -r \sin \phi \\ +r \cos \phi \end{pmatrix}.\end{aligned}\quad (99)$$

At each point  $\mathbf{r}(r, \phi)$ , these are **tangential vectors** to the two coordinate lines. The magnitudes

$$h_r(r, \phi) \equiv |\mathbf{h}_r(r, \phi)| = 1, \quad h_\phi(r, \phi) \equiv |\mathbf{h}_\phi(r, \phi)| = r, \quad (100)$$

are called **scale factors**. For orthogonal coordinates, their product yields the **Jacobian**,

$$J(r, \phi) = |\mathbf{h}_r(r, \phi)| |\mathbf{h}_\phi(r, \phi)| = r. \quad (101)$$

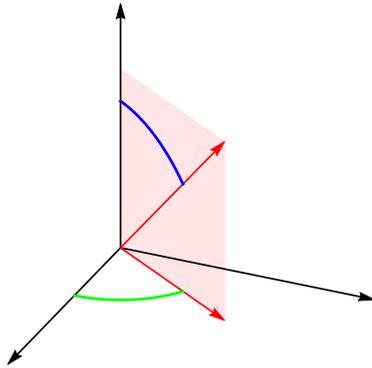
## 4.2 Spherical (polar) coordinates

The position vector  $\mathbf{r}$  of a point  $P$  in  $xyz$ -space with **cartesian coordinates**  $(x, y, z)$ ,

$$\mathbf{r} = x\mathbf{u}_1 + y\mathbf{u}_2 + z\mathbf{u}_3 = \begin{pmatrix} x \\ y \\ z \end{pmatrix}, \quad (102)$$

is fixed as well by its **spherical (polar) coordinates**  $(r, \theta, \phi)$ :

$r$	$0 \leq r < \infty$	<b>Length</b> $ \mathbf{r} $ of $\mathbf{r}$	Distance of $P$ from the origin $O$ .
$\theta$	$0 \leq \theta \leq \pi$	<b>Polar angle</b> of $\mathbf{r}$	Angle between $\mathbf{r}$ and the $z$ -axis (polar axis) $\mathbf{u}_3$
$\phi$	$0 \leq \phi < 2\pi$	<b>Azimuth angle</b> of $\mathbf{r}$	Angle in the $xy$ -plane from the $x$ -axis $\mathbf{u}_1$ (ccw) to the projection $\mathbf{r}_{xy}$ of $\mathbf{r}$ onto the $xy$ -plane



### 4.2.1 Coordinate lines and surfaces

We express the “old” coordinates  $(x, y, z)$  in terms of the “new” ones  $(r, \theta, \phi)$ ,

$$\mathbf{r}(r, \theta, \phi) \equiv \begin{pmatrix} x(r, \theta, \phi) \\ y(r, \theta, \phi) \\ z(r, \theta, \phi) \end{pmatrix} = \begin{pmatrix} r \sin \theta \cos \phi \\ r \sin \theta \sin \phi \\ r \cos \theta \end{pmatrix}. \quad (103)$$

This is a vector function of **three** arguments ( $r$ ,  $\theta$ , and  $\phi$ ).

Keeping **two** arguments fixed, we are generating the **coordinate lines**:

- The point  $\mathbf{r}(r, \theta_0, \phi_0)$ , with growing  $r$ , moves on a **ray**, starting at  $\mathbf{r} = \mathbf{0}$ .
- The point  $\mathbf{r}(r_0, \theta, \phi_0)$ , with growing  $\theta$ , moves on a vertical **semi circle**.
- The point  $\mathbf{r}(r_0, \theta_0, \phi)$ , with growing  $\phi$ , moves on a horizontal **circle**.

Keeping only **one** argument fixed, we are generating the **coordinate surfaces**:

- The point  $\mathbf{r}(r_0, \theta, \phi)$  moves on a **spherical surface** with radius  $r_0$ .
- The point  $\mathbf{r}(r, \theta_0, \phi)$  moves on a **conical surface** with opening angle  $\theta_0$ .
- The point  $\mathbf{r}(r, \theta, \phi_0)$  moves on a **semi plane**, bounded by the  $z$ -axis.

Except for points on the  $z$ -axis, the functions  $x(r, \theta, \phi)$ ,  $y(\dots)$ ,  $z(\dots)$  can be **inverted**:

- Generally (including points on the  $z$ -axis), we have

$$r(x, y, z) = \sqrt{x^2 + y^2 + z^2}. \quad (104)$$

- For  $\mathbf{r} \neq \mathbf{0} \Leftrightarrow r \neq 0$ , we have

$$\begin{aligned} \theta(x, y, z) &= \arccos \frac{z}{r} \\ &\equiv \arccos \frac{z}{\sqrt{x^2 + y^2 + z^2}}, \end{aligned} \quad (105)$$

while  $\theta$  is not defined for  $r = 0$ .

- Eventually, for  $x^2 + y^2 \neq 0$ , we have

$$\phi(x, y) = \begin{cases} \arctan \frac{y}{x} & (x > 0, y \geq 0), \\ \pi/2 & (x = 0, y > 0), \\ \pi + \arctan \frac{y}{x} & (x < 0), \\ 3\pi/2 & (x = 0, y < 0), \\ 2\pi + \arctan \frac{y}{x} & (x > 0, y < 0), \end{cases} \quad (106)$$

while  $\phi$  is not defined for points on the  $z$ -axis (with  $x^2 + y^2 = 0$ ).

#### 4.2.2 Triple integrals

A sphere (radius  $R$ , center at the origin) has the **electric dipole** charge density

$$\rho(\mathbf{r}) \equiv \rho(x, y, z) = a \cdot z \quad (a > 0). \quad (107)$$

(What is the physical dimension of the constant  $a > 0$  ?)

The amount  $q_\Omega$  of charge on the upper half sphere  $\Omega$  ( $z \geq 0$ ),

$$\Omega = \left\{ (x, y, z) \mid x^2 + y^2 + z^2 \leq R^2, \quad z \geq 0 \right\}, \quad (108)$$

is in **cartesian coordinates** given by the triple integral

$$\int_{\Omega} d^3r \rho(\mathbf{r}) = \int_{-R}^R dx \int_{-\sqrt{R^2-x^2}}^{\sqrt{R^2-x^2}} dy \int_0^{\sqrt{R^2-x^2-y^2}} dz \rho(x, y, z). \quad (109)$$

In terms of **spherical polar coordinates**  $(r, \theta, \phi)$ , writing

$$\tilde{\rho}(r, \theta, \phi) = \rho(r \sin \theta \cos \phi, r \sin \theta \sin \phi, r \cos \theta), \quad (110)$$

the limits of this integral are simplified considerably,

$$\int_{\Omega} d^3r \rho(\mathbf{r}) = \int_0^R dr \int_0^{\pi/2} d\theta \int_0^{2\pi} d\phi J(r, \theta, \phi) \tilde{\rho}(r, \theta, \phi), \quad (111)$$

with the **Jacobian of spherical polar coordinates**,

$$J(r, \theta, \phi) = r^2 \sin \theta. \quad (112)$$

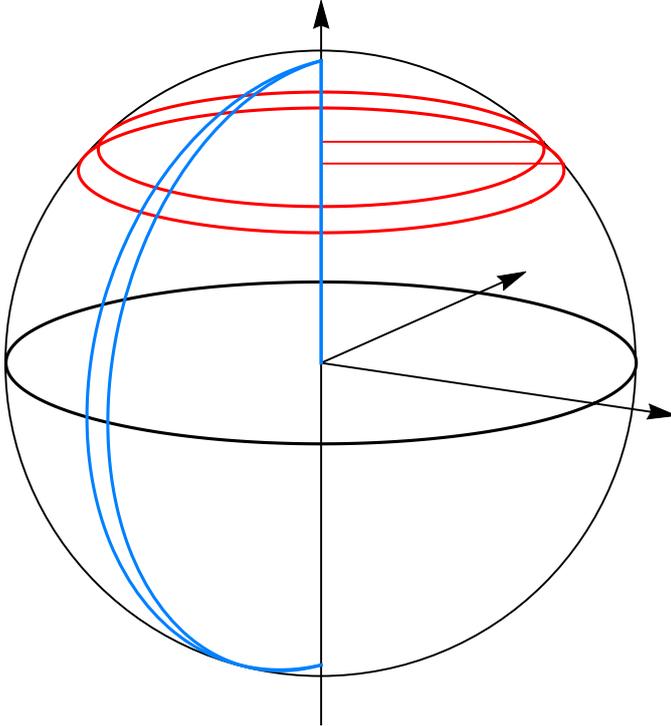
This extra factor arises, since

$$\Delta V_0 = \Delta r \cdot \Delta \theta \cdot \Delta \phi \cdot J(r_0, \theta_0, \phi_0) = \Delta r \cdot (r_0 \Delta \theta) \cdot (r_0 \sin \theta_0 \Delta \phi) \quad (113)$$

is the volume of the small almost-cuboid

$$\Delta \Omega = \left\{ \mathbf{r}(r, \theta, \phi) \left| \begin{array}{l} r_0 - \frac{1}{2} \Delta r \leq r \leq r_0 + \frac{1}{2} \Delta r, \\ \theta_0 - \frac{1}{2} \Delta \theta \leq \theta \leq \theta_0 + \frac{1}{2} \Delta \theta, \\ \phi_0 - \frac{1}{2} \Delta \phi \leq \phi \leq \phi_0 + \frac{1}{2} \Delta \phi \end{array} \right. \right\} \quad (114)$$

surrounding the point  $\mathbf{r}_0 = \mathbf{r}(r_0, \theta_0, \phi_0)$ .



Now, using  $\tilde{\rho}(r, \theta, \phi) = a \cdot r \cos \theta$ , we can evaluate the integral,

$$\begin{aligned} \int_{\Omega} d^3r \rho(\mathbf{r}) &= \int_0^R dr \int_0^{\pi/2} d\theta \int_0^{2\pi} d\phi r^2 \sin \theta (a \cdot r \cos \theta) \\ &= 2\pi a \int_0^R dr \int_0^{\pi/2} d\theta r^3 \sin \theta \cos \theta \\ &= 2\pi a \int_0^R dr \left[ \frac{r^3}{2} \sin^2 \theta \right]_0^{\pi/2} = \pi a \left[ \frac{r^4}{4} \right]_0^R = \frac{\pi}{4} a R^4. \end{aligned} \quad (115)$$

### 4.2.3 Scale factors

Consider the partial derivatives of the three-argument vector function  $\mathbf{r}(r, \theta, \phi)$ ,

$$\begin{aligned} \mathbf{h}_r(r, \theta, \phi) &= \frac{\partial \mathbf{r}(r, \theta, \phi)}{\partial r} = \frac{\partial}{\partial r} \begin{pmatrix} r \sin \theta \cos \phi \\ r \sin \theta \sin \phi \\ r \cos \theta \end{pmatrix} = \begin{pmatrix} \sin \theta \cos \phi \\ \sin \theta \sin \phi \\ \cos \theta \end{pmatrix}, \\ \mathbf{h}_\theta(r, \theta, \phi) &= \frac{\partial \mathbf{r}(r, \theta, \phi)}{\partial \theta} = \frac{\partial}{\partial \theta} \begin{pmatrix} r \sin \theta \cos \phi \\ r \sin \theta \sin \phi \\ r \cos \theta \end{pmatrix} = \begin{pmatrix} r \cos \theta \cos \phi \\ r \cos \theta \sin \phi \\ -r \sin \theta \end{pmatrix}, \\ \mathbf{h}_\phi(r, \theta, \phi) &= \frac{\partial \mathbf{r}(r, \theta, \phi)}{\partial \phi} = \frac{\partial}{\partial \phi} \begin{pmatrix} r \sin \theta \cos \phi \\ r \sin \theta \sin \phi \\ r \cos \theta \end{pmatrix} = \begin{pmatrix} -r \sin \theta \sin \phi \\ r \sin \theta \cos \phi \\ 0 \end{pmatrix}, \end{aligned} \quad (116)$$

At each point  $\mathbf{r}(r, \theta, \phi)$ , these are **tangential vectors** to the three coordinate lines. The magnitudes  $h_u(r, \theta, \phi) \equiv |\mathbf{h}_u(r, \theta, \phi)|$

$$h_r(r, \theta, \phi) = 1, \quad h_\theta(r, \theta, \phi) = r, \quad h_\phi(r, \theta, \phi) = r \sin \theta, \quad (117)$$

are called **scale factors**. Again, their product yields the **Jacobian**,

$$J(r, \theta, \phi) = |\mathbf{h}_r| |\mathbf{h}_\theta| |\mathbf{h}_\phi| = r^2 \sin \theta. \quad (118)$$

## 4.3 General (orthogonal) curvilinear coordinates

Generalizing  $(r, \theta, \phi)$ , we now consider **any** set  $\alpha = (u, v, w)$  of three coordinates which uniquely specify the cartesian coordinates  $(x, y, z)$  of each point  $\mathbf{r}$  in space

$$\mathbf{r}(u, v, w) \equiv \begin{pmatrix} x(u, v, w) \\ y(u, v, w) \\ z(u, v, w) \end{pmatrix} = \mathbf{r}(\alpha), \quad (119)$$

with **some** functions  $x(u, v, w)$ ,  $y(u, v, w)$ , and  $z(u, v, w)$ .

### 4.3.1 Scale factors

**Tangential vectors** to the three coo.-lines through  $\mathbf{r}(\alpha)$  are given by

$$\mathbf{h}_1(\alpha) = \frac{\partial \mathbf{r}(\alpha)}{\partial u}, \quad \mathbf{h}_2(\alpha) = \frac{\partial \mathbf{r}(\alpha)}{\partial v}, \quad \mathbf{h}_3(\alpha) = \frac{\partial \mathbf{r}(\alpha)}{\partial w}. \quad (120)$$

**Def.:** The coordinates  $\alpha = (u, v, w)$  are called **orthogonal**, when

$$\mathbf{h}_i(\alpha) \cdot \mathbf{h}_j(\alpha) = 0 \quad (i \neq j), \quad \text{at each point } \mathbf{r}(\alpha). \quad (121)$$

The set  $\{\mathbf{h}_1(\alpha), \mathbf{h}_2(\alpha), \mathbf{h}_3(\alpha)\}$  is called the **local basis** at  $\mathbf{r}(\alpha)$ .

The magnitudes  $h_i(\alpha) := |\mathbf{h}_i(\alpha)|$  are called the **scale factors** at  $\mathbf{r}(\alpha)$ .

**Recall:** For spherical coordinates  $(u_1, u_2, u_3) = (r, \theta, \phi)$ , the scale factors are

$$h_r = 1, \quad h_\theta = r, \quad h_\phi = r \sin \theta. \quad (122)$$

### 4.3.2 Multiple integrals and Jacobian

For a given function (scalar field)  $f(\mathbf{r}) = f(x, y, z)$ , we wish to evaluate the triple integral

$$\int_{\Omega} d^3r f(\mathbf{r}) \quad (123)$$

over some volume region  $\Omega$ .

Suppose we can find proper curvilinear coo's  $(u, v, w)$  such that

$$\Omega = \left\{ \mathbf{r}(u, v, w) \mid a_1 \leq u \leq a_2, \quad b_1 \leq v \leq b_2, \quad c_1 \leq w \leq c_2 \right\}, \quad (124)$$

with some **constants**  $a_{1,2}, b_{1,2}, c_{1,2}$ . Then, we have

$$\int_{\Omega} d^3r f(\mathbf{r}) = \int_{a_1}^{a_2} du \int_{b_1}^{b_2} dv \int_{c_1}^{c_2} dw J(u, v, w) \tilde{f}(u, v, w), \quad (125)$$

where

$$\tilde{f}(\alpha) = f(x(\alpha), y(\alpha), z(\alpha)) \quad (126)$$

and the **Jacobian** is generally given by

$$J(\alpha) = [\mathbf{h}_1(\alpha) \times \mathbf{h}_2(\alpha)] \cdot \mathbf{h}_3(\alpha) \equiv \det \begin{pmatrix} \partial x / \partial u & \partial x / \partial v & \partial x / \partial w \\ \partial y / \partial u & \partial y / \partial v & \partial y / \partial w \\ \partial z / \partial u & \partial z / \partial v & \partial z / \partial w \end{pmatrix}. \quad (127)$$

When the coo's  $(u, v, w)$  are **orthogonal**, this expression is simplified to

$$J(\alpha) = h_1(\alpha) h_2(\alpha) h_3(\alpha). \quad (128)$$

**Ideas for a proof:** With  $\mathbf{r}(\alpha_0) = \mathbf{r}(u_0, v_0, w_0)$ , we have for sufficiently small  $\Delta u$

$$\begin{aligned} \Delta u \cdot h_1(\alpha_0) &\equiv \Delta u \cdot \lim_{\delta u \rightarrow 0} \frac{1}{\delta u} \left| \mathbf{r}(u_0 + \delta u, v_0, w_0) - \mathbf{r}(u_0, v_0, w_0) \right| \\ &\approx \Delta s_u, \end{aligned} \quad (129)$$

with the **arc length**  $\Delta s_u$  on the  $u$ -line between  $\mathbf{r}(\alpha_0) \equiv \mathbf{r}(u_0, v_0, w_0)$  and  $\mathbf{r}(u_0 + \Delta u, v_0, w_0)$ .

Consequently, for sufficiently small  $\Delta u$ ,  $\Delta v$ , and  $\Delta w$ , the product

$$\Delta u \Delta v \Delta w h_1(\alpha_0) h_2(\alpha_0) h_3(\alpha_0) \approx \Delta s_u \Delta s_v \Delta s_w \approx \Delta V \quad (130)$$

in case of orthogonal coo.'s becomes the volume of the small region (almost-cuboid)

$$\Delta \Omega = \left\{ \mathbf{r}(u, v, w) \mid |u - u_0| \leq \frac{1}{2} \Delta u, \quad |v - v_0| \leq \frac{1}{2} \Delta v, \quad |w - w_0| \leq \frac{1}{2} \Delta w \right\} \quad (131)$$

surrounding the point  $\mathbf{r}(\alpha_0)$ , cf. Eq. (114) and red marking in the figure there.

**For later purposes:** For sufficiently small  $\Delta u, \Delta v > 0$ , the product

$$\Delta u \Delta v [\mathbf{h}_1(\alpha_0) \times \mathbf{h}_2(\alpha_0)] \approx \Delta \mathbf{A} \quad (132)$$

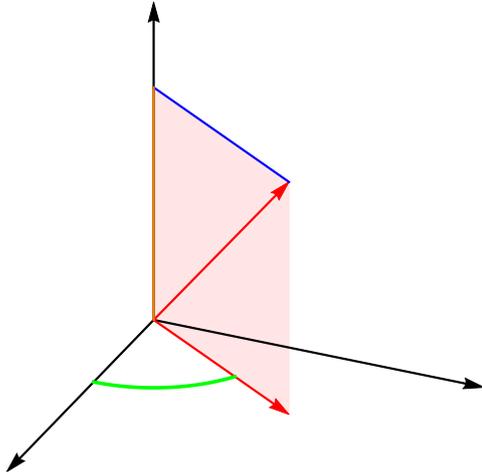
is that **vector**, which is orthogonal to the coordinate surface  $w = w_0$  at the point  $\mathbf{r}(\alpha_0)$  and whose magnitude  $|\Delta \mathbf{A}|$  equals the area of the piece of surface, given by

$$\Delta \Sigma = \left\{ \mathbf{r}(u, v, w_0) \mid |u - u_0| \leq \frac{1}{2} \Delta u, \quad |v - v_0| \leq \frac{1}{2} \Delta v \right\}. \quad (133)$$

### 4.3.3 Further example: Cylindrical (polar) coordinates

These are  $\alpha = (s, \phi, z)$ , defined as

$s =  \mathbf{r}_{xy} $	$0 \leq s < \infty$		Distance of $P$ from the $z$ -axis.
$\phi$	$0 \leq \phi < 2\pi$	<b>Azimuth angle</b>	(the same as for spherical polars)
$z$	$-\infty < z < \infty$		the cartesian $z$ coordinate



Again, we express the “old” coordinates  $(x, y, z)$  in terms of the “new” ones  $(s, \phi, z)$ ,

$$\mathbf{r}(s, \phi, z) \equiv \begin{pmatrix} x(s, \phi, z) \\ y(s, \phi, z) \\ z(s, \phi, z) \end{pmatrix} = \begin{pmatrix} s \cos \phi \\ s \sin \phi \\ z \end{pmatrix}. \quad (134)$$

This is a vector function of **three** arguments ( $s$ ,  $\phi$ , and  $z$ ).

- Keeping **two** arguments fixed, we are generating the **coordinate lines**:  
 $s$ -lines: Horizontal **rays**,  $\phi$ -lines: Horizontal **circles**,  $z$ -lines: Vertical **lines**.
- Keeping only **one** argument fixed, we are generating the **coordinate surfaces**:  
 Vertical cylindrical surfaces, vertical semi planes, and horizontal planes.

In terms of the **tangential vectors**

$$\mathbf{h}_s(\alpha) = \begin{pmatrix} \cos \phi \\ \sin \phi \\ 0 \end{pmatrix}, \quad \mathbf{h}_\phi(\alpha) = \begin{pmatrix} -s \sin \phi \\ +s \cos \phi \\ 0 \end{pmatrix}, \quad \mathbf{h}_z(\alpha) = \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix} \quad (135)$$

(which are mutually **orthogonal!**), the **Jacobian** is

$$J(s, \phi, z) \equiv |\mathbf{h}_s(\alpha)| |\mathbf{h}_\phi(\alpha)| |\mathbf{h}_z(\alpha)| = s. \quad (136)$$

## 5 Vector fields

A **vector field** assigns to each point  $\mathbf{r}$  in space a **vector**

$$\mathbf{v}(\mathbf{r}) \equiv \mathbf{v}(x, y, z) = \begin{pmatrix} v_1(x, y, z) \\ v_2(x, y, z) \\ v_3(x, y, z) \end{pmatrix}. \quad (137)$$

$\mathbf{v}(\mathbf{r})$  may be, at windy weather for example, the **velocity of flowing air** at position  $\mathbf{r}$ . Another example is the **electrostatic field**  $\mathbf{E}(\mathbf{r})$  **due to** a charge distribution.

We also consider 2D vector fields (on the  $xy$ -plane),

$$\mathbf{v}(\mathbf{r}) \equiv \mathbf{v}(x, y) = \begin{pmatrix} v_1(x, y) \\ v_2(x, y) \end{pmatrix}. \quad (138)$$

### 5.1 Gradient fields als a special case

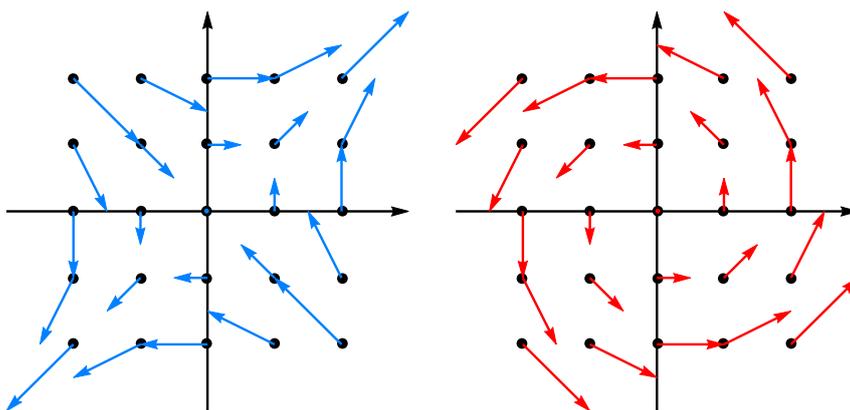
A particularly important example is the **gradient of a scalar field**  $f(\mathbf{r}) \equiv f(x, y, z)$ ,

$$\mathbf{G}_f(\mathbf{r}) = \nabla f(\mathbf{r}) \equiv \begin{pmatrix} f_x(x, y, z) \\ f_y(x, y, z) \\ f_z(x, y, z) \end{pmatrix}. \quad (139)$$

**Example:** Consider the 2D vector field (blue in the figure below, NEW 2019 !!!)

$$\mathbf{p}(\mathbf{r}) \equiv \begin{pmatrix} p_1(x, y) \\ p_2(x, y) \end{pmatrix} = \begin{pmatrix} y \\ x \end{pmatrix}. \quad (140)$$

Obviously,  $\mathbf{p}(\mathbf{r}) = \nabla f(\mathbf{r})$ , with  $f(x, y) = xy$ .



In contrast, the quite similar 2D vector field (red in the figure)

$$\mathbf{q}(\mathbf{r}) \equiv \begin{pmatrix} q_1(x, y) \\ q_2(x, y) \end{pmatrix} = \begin{pmatrix} -y \\ x \end{pmatrix} \quad (141)$$

can obviously **not** be the gradient of any continuous 2D scalar field  $f(\mathbf{r}) = f(x, y)$ : The value  $f(\mathbf{r})$  would increase constantly as the point  $\mathbf{r}$  is moving on a circle centered at the origin  $\mathbf{r} = \mathbf{0}$ . After one full turn, this would conflict with the continuity of  $f$ .  $\square$

- When a 2D vector field  $\mathbf{v}(\mathbf{r})$  shall be gradient of some scalar field  $f(\mathbf{r})$ ,

$$\mathbf{v}(\mathbf{r}) \equiv \begin{pmatrix} v_1(x, y) \\ v_2(x, y) \end{pmatrix} = \nabla f(\mathbf{r}) \equiv \begin{pmatrix} \frac{\partial}{\partial x} f(x, y) \\ \frac{\partial}{\partial y} f(x, y) \end{pmatrix}, \quad (142)$$

then its components  $v_1$  and  $v_2$  must satisfy the **necessary condition**

$$\frac{\partial}{\partial y} v_1(x, y) = \frac{\partial}{\partial y} \left[ \frac{\partial}{\partial x} f(x, y) \right] \equiv \frac{\partial}{\partial x} \left[ \frac{\partial}{\partial y} f(x, y) \right] = \frac{\partial}{\partial x} v_2(x, y). \quad (143)$$

**Example:** For the above fields  $\mathbf{p}(\mathbf{r})$  and  $\mathbf{q}(\mathbf{r})$ , we have

$$\frac{\partial}{\partial y} p_1(x, y) = 1 = \frac{\partial}{\partial x} p_2(x, y), \quad \frac{\partial}{\partial y} q_1(x, y) = -1 \neq 1 = \frac{\partial}{\partial x} q_2(x, y), \quad (144)$$

reveiling immediately that  $\mathbf{p}(\mathbf{r})$  is a gradient field, while  $\mathbf{q}(\mathbf{r})$  is not.

- For a 3D vector field, Eq. (137), we obtain **three necessary conditions**,

$$\frac{\partial v_2}{\partial x} = \frac{\partial v_1}{\partial y}, \quad \frac{\partial v_3}{\partial y} = \frac{\partial v_2}{\partial z}, \quad \frac{\partial v_1}{\partial z} = \frac{\partial v_3}{\partial x}. \quad (145)$$

Using the nabla operator  $\nabla$  and writing  $(x, y, z) = (x_1, x_2, x_3)$ ,  $\frac{\partial}{\partial x_n} = \partial_n$ , these conditions are summarized to a single **vector condition**,

$$\nabla \times \mathbf{v}(\mathbf{r}) \equiv \begin{pmatrix} \partial_1 \\ \partial_2 \\ \partial_3 \end{pmatrix} \times \begin{pmatrix} v_1 \\ v_2 \\ v_3 \end{pmatrix} \equiv \begin{pmatrix} \partial_2 v_3 - \partial_3 v_2 \\ \partial_3 v_1 - \partial_1 v_3 \\ \partial_1 v_2 - \partial_2 v_1 \end{pmatrix} = \mathbf{0}. \quad (146)$$

The **vector field**  $\nabla \times \mathbf{v}(\mathbf{r})$  is called the **curl** of the vector field  $\mathbf{v}(\mathbf{r})$ .

**Remark 1:** A 2D vector field is a 3D one with a zero third component,

$$\begin{pmatrix} v_1(x, y) \\ v_2(x, y) \end{pmatrix} \rightarrow \begin{pmatrix} v_1(x, y) \\ v_2(x, y) \\ 0 \end{pmatrix}, \quad \nabla \times \mathbf{v}(\mathbf{r}) = \begin{pmatrix} 0 \\ 0 \\ \partial_1 v_2 - \partial_2 v_1 \end{pmatrix}. \quad (147)$$

**Remark 2:** In section 6.4, we shall see that, under mild assumptions, the condition  $\nabla \times \mathbf{v}(\mathbf{r}) \equiv \mathbf{0}$  is even **sufficient** for  $\mathbf{v}(\mathbf{r})$  to be a gradient field.

## 5.2 Line integrals

In a uniform electric field  $\mathbf{E}$  (plate capacitor), a point charge  $q$  (mass  $m$ ) experiences the force  $\mathbf{F} = q\mathbf{E}$ . As  $q$  is moving without friction on a **straight** track from position  $\mathbf{r}_A$  to position  $\mathbf{r}_B$ , its velocities  $v_{A,B}$  at these positions are related by

$$\frac{m}{2} v_B^2 = \frac{m}{2} v_A^2 + W_{AB}, \quad (148)$$

where  $W_{AB}$  is the **work** done by the force  $\mathbf{F}$ ,

$$W_{AB} = \left( |\mathbf{F}| \cos \gamma \right) |\mathbf{r}_B - \mathbf{r}_A| \equiv \mathbf{F} \cdot [\mathbf{r}_B - \mathbf{r}_A]. \quad (149)$$

Note:  $|\mathbf{F}| \cos \gamma$  is the component of  $\mathbf{F}$  in direction of the vector  $\Delta \mathbf{r} = \mathbf{r}_B - \mathbf{r}_A$ .

More generally, let  $\mathbf{E}(\mathbf{r})$  be a **non-uniform vector field**. Then, both magnitude and direction of the force vector  $\mathbf{F} = q\mathbf{E}(\mathbf{r}) = \mathbf{F}(\mathbf{r})$  will change as  $q$  is moving. Furthermore, the track may have the shape of some **curve**  $\Gamma$ , given by the **vector function**

$$\mathbf{r}(u) \equiv \begin{pmatrix} x(u) \\ y(u) \\ z(u) \end{pmatrix}, \quad \Gamma = \left\{ \mathbf{r}(u) \mid u_A \leq u \leq u_B \right\}. \quad (150)$$

We partition the interval  $[u_A, u_B]$  into  $N$  pieces with equal lengths  $\Delta u = \frac{u_B - u_A}{N}$ ,

$$u_n = u_A + n\Delta u \quad (n = 0, 1, 2, \dots, N). \quad (151)$$

Consequently, as  $q$  is moving on  $\Gamma$  from  $\mathbf{r}_A = \mathbf{r}(u_A)$  to  $\mathbf{r}_B = \mathbf{r}(u_B)$ , the work done by the force field  $\mathbf{F}(\mathbf{r})$  is

$$W_{AB} = \lim_{N \rightarrow \infty} \sum_{n=1}^N \mathbf{F}(\mathbf{r}(u_n)) \cdot [\mathbf{r}(u_n) - \mathbf{r}(u_{n-1})]. \quad (152)$$

As  $\Delta u \rightarrow 0$  ( $N \rightarrow \infty$ ), we obtain

$$\mathbf{r}(u_n) - \mathbf{r}(u_{n-1}) \equiv \begin{pmatrix} x(u_n) - x(u_{n-1}) \\ y(u_n) - y(u_{n-1}) \\ z(u_n) - z(u_{n-1}) \end{pmatrix} \rightarrow \begin{pmatrix} \dot{x}(u_n) \\ \dot{y}(u_n) \\ \dot{z}(u_n) \end{pmatrix} \Delta u = \dot{\mathbf{r}}(u_n) \Delta u, \quad (153)$$

where the derivatives of  $x(u)$ , etc. are denoted as  $\dot{x}(u)$ . Eventually,

$$W_{AB} = \lim_{N \rightarrow \infty} \sum_{n=1}^N [\mathbf{F}(\mathbf{r}(u_n)) \cdot \dot{\mathbf{r}}(u_n)] \Delta u = \int_{u_A}^{u_B} du \underbrace{[\dot{\mathbf{r}}(u) \cdot \mathbf{F}(\mathbf{r}(u))]}_{f(u)}. \quad (154)$$

This is an **ordinary integral** over one real variable  $u$ .

**Def.:** The **line integral** of the vector field  $\mathbf{F}(\mathbf{r})$  along the curve  $\Gamma$  is defined as

$$\int_{\Gamma} d\mathbf{r} \cdot \mathbf{F}(\mathbf{r}) := \int_{u_A}^{u_B} du \dot{\mathbf{r}}(u) \cdot \mathbf{F}(\mathbf{r}(u)) \quad (u_A < u_B). \quad (155)$$

While its **absolute value** is independent of the parametrization  $\mathbf{r}(u)$  of  $\Gamma$ , its **sign** depends on which one of the two endpoints of  $\Gamma$  is chosen as the "starting point  $A$ ".

**Remark:** When  $\Gamma$  is a **closed loop**, with  $\mathbf{r}_B = \mathbf{r}_A$ , the integral symbol may be decorated with a circle,

$$\int_{\Gamma} d\mathbf{r} \cdot \mathbf{F}(\mathbf{r}) \equiv \oint_{\Gamma} d\mathbf{r} \cdot \mathbf{F}(\mathbf{r}). \quad (156)$$

**Example:** We consider the line integral of the vector field

$$\mathbf{F}(\mathbf{r}) = \begin{pmatrix} ax + by \\ cx + dy \end{pmatrix} \quad (157)$$

(with constants  $a, b, c, d \in \mathbb{R}$ ) along a circle  $\Gamma$  of radius  $R$  in the  $xy$ -plane,

$$\Gamma = \left\{ \mathbf{r}(\phi) \mid 0 \leq \phi \leq 2\pi \right\}, \quad \mathbf{r}(\phi) = \begin{pmatrix} x(\phi) \\ y(\phi) \end{pmatrix} = \begin{pmatrix} R \cos \phi \\ R \sin \phi \end{pmatrix}. \quad (158)$$

In this case (closed loop!), we have

$$\dot{\mathbf{r}}(u) = \begin{pmatrix} -R \sin \phi \\ R \cos \phi \end{pmatrix}. \quad (159)$$

Consequently, with  $u_A \equiv \phi_A = 0$  and  $u_B \equiv \phi_B = 2\pi$ , Eq. (155) yields

$$\begin{aligned} \oint_{\Gamma} \mathbf{dr} \cdot \mathbf{F}(\mathbf{r}) &= \int_{\phi_A}^{\phi_B} d\phi \dot{\mathbf{r}}(\phi) \cdot \mathbf{F}(\mathbf{r}(\phi)) \\ &= \int_0^{2\pi} d\phi \begin{pmatrix} -R \sin \phi \\ R \cos \phi \end{pmatrix} \cdot \begin{pmatrix} aR \cos \phi + bR \sin \phi \\ cR \cos \phi + dR \sin \phi \end{pmatrix} \\ &= R^2 \int_0^{2\pi} d\phi \left\{ (d-a) \sin \phi \cos \phi - b \sin^2 \phi + c \cos^2 \phi \right\} \\ &= R^2 \left\{ (d-a) \left[ \frac{\sin^2 \phi}{2} \right]_0^{2\pi} - b \left[ \frac{\phi - \sin^2 \phi \cos \phi}{2} \right]_0^{2\pi} + c \left[ \frac{\phi + \sin^2 \phi \cos \phi}{2} \right]_0^{2\pi} \right\} \\ &= R^2 \pi (c-b). \end{aligned} \quad (160)$$

Remarkably, this result does not depend on the values of  $a$  and  $d$ .

Choosing  $a = d = 0$ ,  $c = 1$  and  $b = \pm 1$ , respectively, the vector fields  $\mathbf{p}(\mathbf{r}) = \begin{pmatrix} y \\ x \end{pmatrix}$  and  $\mathbf{q}(\mathbf{r}) = \begin{pmatrix} -y \\ x \end{pmatrix}$  of section 5.1 are recovered [where  $\mathbf{p}(\mathbf{r}) = \nabla xy$  is a gradient], implying that

$$\oint_{\Gamma} \mathbf{dr} \cdot \mathbf{p}(\mathbf{r}) = 0, \quad \oint_{\Gamma} \mathbf{dr} \cdot \mathbf{q}(\mathbf{r}) = 2R^2\pi. \quad (161)$$

The first one of these two special results has a simple reason:

**Theorem:** For the gradient  $\mathbf{G}_f(\mathbf{r})$  of a scalar field,  $\mathbf{G}_f(\mathbf{r}) = \nabla f(\mathbf{r})$ , the line integral is

$$\int_{\Gamma} \mathbf{dr} \cdot \mathbf{G}_f(\mathbf{r}) \equiv \int_{\Gamma} \mathbf{dr} \cdot \nabla f(\mathbf{r}) = f(\mathbf{r}_B) - f(\mathbf{r}_A), \quad (162)$$

where  $\mathbf{r}_A = \mathbf{r}(u_A)$  is the starting and  $\mathbf{r}_B = \mathbf{r}(u_B)$  the end point of  $\Gamma$ .

**Proof:** The chain rule for partial derivatives,  $\frac{d}{du} f(\mathbf{r}(u)) = \dot{\mathbf{r}}(u) \cdot \mathbf{G}_f(\mathbf{r}(u))$ , implies

$$\begin{aligned} \int_{\Gamma} \mathbf{dr} \cdot \mathbf{G}_f(\mathbf{r}) &\equiv \int_{u_A}^{u_B} du \dot{\mathbf{r}}(u) \cdot \mathbf{G}_f(\mathbf{r}(u)) \\ &= \int_{u_A}^{u_B} du \frac{d}{du} f(\mathbf{r}(u)) = f(\mathbf{r}(u_B)) - f(\mathbf{r}(u_A)). \end{aligned} \quad (163)$$

In other words: The line integral of a **gradient field** only depends on the end points of the curve, but not on its shape. Consequently, for a **closed loop**  $\Gamma$  (with  $\mathbf{r}_A = \mathbf{r}_B$ ),

$$\oint_{\Gamma} \mathbf{dr} \cdot \nabla f(\mathbf{r}) = 0. \quad (164)$$

**Remark:** When the **arc length**  $s$  is used as the curve parameter, then  $\dot{\mathbf{r}}(s) = \frac{d}{ds} \mathbf{r}(s)$  is the **unit** vector  $\mathcal{T}(\mathbf{r})$  tangential to  $\Gamma$  at the point  $\mathbf{r} = \mathbf{r}(s) \in \Gamma$ . Therefore,

$$\int_{\Gamma} \mathbf{dr} \cdot \mathbf{F}(\mathbf{r}) \equiv \int_{s_A}^{s_B} ds \dot{\mathbf{r}}(s) \cdot \mathbf{F}(\mathbf{r}(s)) = L_{\Gamma} \langle \mathcal{T}(\mathbf{r}) \cdot \mathbf{F}(\mathbf{r}) \rangle_{\mathbf{r} \in \Gamma} \quad (165)$$

or, in words: The line integral equals the average value of the tangential component  $\mathcal{T} \cdot \mathbf{F}$  of the vector field  $\mathbf{F}(\mathbf{r})$  on the curve  $\Gamma$  times the length  $L_{\Gamma} = s_B - s_A$  of  $\Gamma$ .

### 5.3 Flux (or surface) integrals

We consider **flowing air** with uniform and time-independent density  $\rho(\mathbf{r}) = \rho_0$ , e.g.,

$$\rho_0 \approx \frac{0.7 \cdot 28 \text{ kg} + 0.3 \cdot 32 \text{ kg}}{22\,400 \ell} = \frac{29.2 \text{ kg}}{22.4 \text{ m}^3} = 1.30 \frac{\text{kg}}{\text{m}^3}. \quad (166)$$

The flow velocity shall also be uniform,  $\mathbf{v}(\mathbf{r}) = \mathbf{v}_0$ , e.g. with  $|\mathbf{v}_0| = 2.00 \frac{\text{m}}{\text{s}}$ . The vector

$$\mathbf{J} = \rho_0 \mathbf{v}_0 = 2.60 \mathbf{u}_J \frac{\text{kg}}{\text{m}^2 \text{s}} \quad \left( \mathbf{u}_J := \frac{\mathbf{v}_0}{|\mathbf{v}_0|} \right) \quad (167)$$

is called **current density** or **flux**. Its magnitude  $J = |\mathbf{J}|$  gives the amount of air (in kg) flowing per unit time (1 s) through a unit area (1 m<sup>2</sup>) which has  $\mathbf{J}$  as a normal vector.

Consequently, the amount of air flowing per unit time through a planar piece of surface  $\Sigma$ , with area  $A$  and unit normal vector  $\mathcal{N}$ , is given by the scalar product

$$\dot{M} = (A\mathcal{N}) \cdot \mathbf{J} = AJ \cos \gamma, \quad (168)$$

where  $\gamma$  is the angle between the vectors  $\mathbf{J}$  and  $\mathbf{n}$ . Note that  $\dot{M}$  changes sign when the vector  $\mathcal{N}$  is replaced with  $-\mathcal{N}$ . The two possible choices  $\pm\mathcal{N}$  are called the **orientations** of  $\Sigma$ . If its orientation is specified,  $\Sigma$  is called **orientated**.

We wish to generalize this result to a flow with **non-uniform** current density, represented by a vector field  $\mathbf{J}(\mathbf{r})$ , through a **curved** (and orientated<sup>2</sup>) piece of surface

$$\Sigma = \left\{ \mathbf{r}(u, v, w_0) \mid a_1 \leq u \leq a_2, b_1 \leq v \leq b_2 \right\}, \quad (169)$$

represented by a proper set  $(u, v, w) \equiv \alpha$  of curvilinear coordinates.

Let  $\mathcal{N}(u, v)$  be the unit normal vector to  $\Sigma$  at the point  $\mathbf{r}(u, v, w_0) \in \Sigma$  and let

$$\delta A = \delta u \delta v A(u, v) \quad (170)$$

be the area of the small piece  $\delta\Sigma$  of  $\Sigma$ , consisting of the points  $\mathbf{r}(u', v', w_0)$  with

$$u \leq u' \leq u + \delta u, \quad v \leq v' \leq v + \delta v. \quad (171)$$

Then, the wanted generalization of Eq. (168) is obviously given by the double integral

$$\dot{M} = \int_{a_1}^{a_2} du \int_{b_1}^{b_2} dv A(u, v) \mathcal{N}(u, v) \cdot \mathbf{J}(\mathbf{r}(u, v, w_0)). \quad (172)$$

Generally, it can be shown (see section 4.3) that

$$A(u, v) \mathcal{N}(u, v) = \pm \mathbf{h}_u(u, v, w_0) \times \mathbf{h}_v(u, v, w_0) \quad (173)$$

(with a sign depending on the orientation of  $\Sigma$ ) and we obtain the compact formula

$$\dot{M} = \pm \int_{a_1}^{a_2} du \int_{b_1}^{b_2} dv [\mathbf{h}_u(\alpha) \times \mathbf{h}_v(\alpha)] \cdot \mathbf{J}(\mathbf{r}(\alpha)) \Big|_{w=w_0} \equiv \int_{\Sigma} d\mathbf{A} \cdot \mathbf{J}(\mathbf{r}). \quad (174)$$

- This is called the **flux (or surface) integral** of the vector field  $\mathbf{J}(\mathbf{r})$  through  $\Sigma$ .
- $d\mathbf{A}$ , the **surface vector element** of  $\Sigma$ , represents the product  $A(u, v) \mathcal{N}(u, v) du dv$ .

<sup>2</sup>There are curved surfaces  $\Sigma$  that cannot be orientated, e.g., a Moebius tape!

**Remark:** An alternative notation for flux integrals is

$$\int_{\Sigma} d\mathbf{A} \cdot \mathbf{F}(\mathbf{r}) \equiv \int_{\Sigma} d^2r \mathcal{N}(\mathbf{r}) \cdot \mathbf{F}(\mathbf{r}), \quad (175)$$

where  $\mathcal{N}(\mathbf{r})$  is the unit vector normal to  $\Sigma$  at the point  $\mathbf{r} \in \Sigma$ . In fact, one can show that

$$\int_{\Sigma} d\mathbf{A} \cdot \mathbf{F}(\mathbf{r}) = A_{\Sigma} \langle \mathcal{N}(\mathbf{r}) \cdot \mathbf{F}(\mathbf{r}) \rangle_{\mathbf{r} \in \Sigma}. \quad (176)$$

This is the average of the normal component  $\mathcal{N} \cdot \mathbf{F}$  of  $\mathbf{F}(\mathbf{r})$  on  $\Sigma$  times the area  $A_{\Sigma}$  of  $\Sigma$ .

**Example:** Choosing spherical polar coordinates,

$$\mathbf{r}(r, \theta, \phi) = \begin{pmatrix} r \sin \theta \cos \phi \\ r \sin \theta \sin \phi \\ r \cos \theta \end{pmatrix} \quad (177)$$

we compute the flux of the following vector field  $\mathbf{F}(\mathbf{r})$  **out of** the following surface  $\Sigma$ ,

$$\mathbf{F}(\mathbf{r}) = \begin{pmatrix} Ax + By + Cz \\ Lx + My + Nz \\ Ux + Vy + Wz \end{pmatrix}, \quad \Sigma = \left\{ \mathbf{r}(R, \theta, \phi) \mid 0 \leq \theta \leq \beta, \quad 0 \leq \phi \leq 2\pi \right\} \quad (178)$$

( $\Sigma$  is the region "north" from the "latitude"  $\theta = \beta$  on a "planet" with radius  $R$ ),

$$\begin{aligned} \int_{\Sigma} d\mathbf{A} \cdot \mathbf{F}(\mathbf{r}) &= \int_0^{\beta} d\theta \int_0^{2\pi} d\phi \left[ \mathbf{h}_{\theta}(\alpha) \times \mathbf{h}_{\phi}(\alpha) \right] \cdot \mathbf{F}(\mathbf{r}(r, \theta, \phi)) \Big|_{r=R} \\ &= \int_0^{\beta} d\theta \int_0^{2\pi} d\phi \left[ \begin{pmatrix} R \cos \theta \cos \phi \\ R \cos \theta \sin \phi \\ -R \sin \theta \end{pmatrix} \times \begin{pmatrix} -R \sin \theta \sin \phi \\ +R \sin \theta \cos \phi \\ 0 \end{pmatrix} \right] \cdot \mathbf{F}(\mathbf{r}(R, \theta, \phi)) \\ &= \int_0^{\beta} d\theta \int_0^{2\pi} d\phi \begin{pmatrix} R^2 \sin^2 \theta \cos \phi \\ R^2 \sin^2 \theta \sin \phi \\ R^2 \cos \theta \sin \theta \end{pmatrix} \cdot \begin{pmatrix} R \sin \theta (A \cos \phi + B \sin \phi) + RC \cos \theta \\ R \sin \theta (L \cos \phi + M \sin \phi) + RN \cos \theta \\ R \sin \theta (U \cos \phi + V \sin \phi) + RW \cos \theta \end{pmatrix}. \end{aligned}$$

It is easy to see that the integral  $\int_0^{2\pi} d\phi \dots$  yields zero for the  $B, C, L, N, U$ , and  $V$  terms. Only the  $A, M$  and  $W$  terms do contribute, and we are left with

$$\begin{aligned} \dots &= R^3 \int_0^{\beta} d\theta \int_0^{2\pi} d\phi \left[ A \sin^3 \theta \cos^2 \phi + M \sin^3 \theta \sin^2 \phi + W \cos^2 \theta \sin \theta \right] \\ &= R^3 \left[ \pi(A + M) \int_0^{\beta} d\theta \sin^3 \theta + 2\pi W \int_0^{\beta} d\theta \cos^2 \theta \sin \theta \right] \quad (179) \end{aligned}$$

Writing  $\sin^3 \theta = \sin \theta (1 - \cos^2 \theta)$ , we obtain

$$\int_{\Sigma} d\mathbf{A} \cdot \mathbf{F}(\mathbf{r}) = \pi R^3 (A + M) \left[ -\cos \theta + \frac{1}{3} \cos^3 \theta \right]_0^{\beta} + 2\pi R^3 W \left[ -\frac{1}{3} \cos^3 \theta \right]_0^{\beta}. \quad (180)$$

In the case  $A = M = W$  (and  $B, C, L, N, U, V = 0$ ), this result can be found even without integration: Since the "current density"  $\mathbf{F}(\mathbf{r}) = W\mathbf{r}$  is **perpendicular** to  $\Sigma$  at any point  $\mathbf{r} \in \Sigma$ , with **constant magnitude**  $|\mathbf{F}| = WR$  for  $\mathbf{r} \in \Sigma$ , the flux (integral) is simply the product of the area  $A_{\Sigma}$  of  $\Sigma$  with this magnitude  $WR$ ,

$$\int_{\Sigma} d\mathbf{A} \cdot \mathbf{F}(\mathbf{r}) = A_{\Sigma} WR. \quad (181)$$

In the case  $\beta = \pi$ ,  $\Sigma$  is the full spherical surface, and Eq. (180) is simplified to

$$\oint_{\Sigma} d\mathbf{A} \cdot \mathbf{F}(\mathbf{r}) = \frac{4\pi}{3} R^3 (A + M + W). \quad (182)$$

## 6 Integral theorems

### 6.1 Integral types in vector calculus, revisited

Consider a set of **orthogonal curvilinear coordinates**  $\alpha = (u, v, w)$ , given by

$$\mathbf{r}(u, v, w) \equiv \mathbf{r}(\alpha) = \begin{pmatrix} x(u, v, w) \\ y(u, v, w) \\ z(u, v, w) \end{pmatrix}. \quad (183)$$

**Tangent vectors** to the coordinate curves through  $\mathbf{r}(\alpha)$  are given by

$$\mathbf{h}_u(\alpha) = \frac{\partial \mathbf{r}(\alpha)}{\partial u}, \quad \mathbf{h}_v(\alpha) = \frac{\partial \mathbf{r}(\alpha)}{\partial v}, \quad \mathbf{h}_w(\alpha) = \frac{\partial \mathbf{r}(\alpha)}{\partial w}. \quad (184)$$

Their magnitudes  $h_u(\alpha) = |\mathbf{h}_u(\alpha)|$ , etc. are called **scale factors**.

In terms of the coordinates  $\alpha = (u, v, w)$ , a **volume region**  $\Omega$ , a piece of **surface**  $\Sigma$ , and a **curve**  $\Gamma$  can be defined by

$$\begin{aligned} \Omega &= \left\{ \mathbf{r}(\alpha) \mid u_1 \leq u \leq u_2, v_1 \leq v \leq v_2, w_1 \leq w \leq w_2 \right\}, \\ \Sigma &= \left\{ \mathbf{r}(\alpha) \mid u_1 \leq u \leq u_2, v_1 \leq v \leq v_2, w = w_0 \right\}, \\ \Gamma &= \left\{ \mathbf{r}(\alpha) \mid u_1 \leq u \leq u_2, v = v_0, w = w_0 \right\}. \end{aligned} \quad (185)$$

We now consider a **scalar field**  $f(\mathbf{r})$  and a **vector field**  $\mathbf{F}(\mathbf{r})$ ,

$$f(\mathbf{r}) = f(x, y, z), \quad \mathbf{F}(\mathbf{r}) = \begin{pmatrix} F_1(x, y, z) \\ F_2(x, y, z) \\ F_3(x, y, z) \end{pmatrix}. \quad (186)$$

Replacing  $\mathbf{r}$  with  $\mathbf{r} = \mathbf{r}(\alpha)$ , we obtain functions of  $\alpha = (u, v, w)$ ,

$$f(\mathbf{r}(\alpha)) = \tilde{f}(u, v, w), \quad \mathbf{F}(\mathbf{r}(\alpha)) = \begin{pmatrix} \tilde{F}_1(u, v, w) \\ \tilde{F}_2(u, v, w) \\ \tilde{F}_3(u, v, w) \end{pmatrix}. \quad (187)$$

The **volume integral** of the scalar field  $f(\mathbf{r})$  over  $\Omega$  is evaluated as

$$\int_{\Omega} d^3r f(\mathbf{r}) = \int_{u_1}^{u_2} du \int_{v_1}^{v_2} dv \int_{w_1}^{w_2} dw [h_u(\alpha)h_v(\alpha)h_w(\alpha)] f(\mathbf{r}(\alpha)), \quad (188)$$

where  $[h_u(\alpha)h_v(\alpha)h_w(\alpha)] \equiv J(u, v, w)$  is the **Jacobian** of the coordinates  $\alpha = (u, v, w)$ .

The **flux integral** of the vector field  $\mathbf{F}(\mathbf{r})$  through  $\Sigma$  is evaluated as

$$\int_{\Sigma} d\mathbf{A} \cdot \mathbf{F}(\mathbf{r}) = \int_{u_1}^{u_2} du \int_{v_1}^{v_2} dv [\mathbf{h}_u(\alpha) \times \mathbf{h}_v(\alpha)] \cdot \mathbf{F}(\mathbf{r}(\alpha)) \Big|_{w=w_0}, \quad (189)$$

where the **orientation** of  $\Sigma$  is fixed by the direction of  $[\mathbf{h}_u(\alpha) \times \mathbf{h}_v(\alpha)]$ .

The **line integral** of the vector field  $\mathbf{F}(\mathbf{r})$  along  $\Gamma$  is evaluated as

$$\int_{\Gamma} d\mathbf{r} \cdot \mathbf{F}(\mathbf{r}) = \int_{u_1}^{u_2} du \mathbf{h}_u(\alpha) \cdot \mathbf{F}(\mathbf{r}(\alpha)) \Big|_{v=v_0, w=w_0}. \quad (190)$$

It is instructive to notice that, in terms of the volume  $V_\Omega$  of  $\Omega$ , the surface area  $A_\Sigma$  of  $\Sigma$ , and the length  $L_\Gamma$  of  $\Gamma$ , we may write

$$\int_\Omega d^3r f(\mathbf{r}) = V_\Omega \langle f(\mathbf{r}) \rangle_{\mathbf{r} \in \Omega}, \quad (191)$$

$$\int_\Sigma d\mathbf{A} \cdot \mathbf{F}(\mathbf{r}) = A_\Sigma \langle \mathcal{N}(\mathbf{r}) \cdot \mathbf{F}(\mathbf{r}) \rangle_{\mathbf{r} \in \Sigma}, \quad (192)$$

$$\int_\Gamma d\mathbf{r} \cdot \mathbf{F}(\mathbf{r}) = L_\Gamma \langle \mathcal{T}(\mathbf{r}) \cdot \mathbf{F}(\mathbf{r}) \rangle_{\mathbf{r} \in \Gamma}, \quad (193)$$

where  $\mathcal{N}(\mathbf{r})$  is the unit normal vector at  $\mathbf{r} \in \Sigma$  and  $\mathcal{T}(\mathbf{r})$  the unit tangent vector at  $\mathbf{r} \in \Gamma$ . In practice, the quantities within  $\langle \dots \rangle$  are often **constant** (independent of  $\mathbf{r}$ ).

## 6.2 Divergence of a vector field and Gauss' theorem

**Def.:** The **divergence** of a vector field  $\mathbf{F}(\mathbf{r})$  is the scalar field  $D_{\mathbf{F}}(\mathbf{r})$ , given by

$$D_{\mathbf{F}}(\mathbf{r}_0) = \lim_{a \rightarrow 0} \frac{1}{a^3} \oint_{\partial\Omega_a} d\mathbf{A} \cdot \mathbf{F}(\mathbf{r}), \quad (194)$$

where  $\Omega_a$  is a cube, centered at  $\mathbf{r} = \mathbf{r}_0$ , with edges of length  $a$  in  $x$ -,  $y$ -, and  $z$ -directions. Remark 1: The limit  $D_{\mathbf{F}}(\mathbf{r}_0)$  does not change when the cube is rotated.

Remark 2: In usual notation,  $D_{\mathbf{F}}(\mathbf{r}) \equiv \text{div } \mathbf{F}(\mathbf{r})$ .

**Corollary (Gauss' theorem):** For any finite volume region  $\Omega$ , we have

$$\int_\Omega d^3r D_{\mathbf{F}}(\mathbf{r}) = \oint_{\partial\Omega} d\mathbf{A} \cdot \mathbf{F}(\mathbf{r}). \quad (195)$$

**Lemma:** In cartesian coordinates,  $\mathbf{F}(\mathbf{r}) = \begin{pmatrix} F_1(x, y, z) \\ F_2(x, y, z) \\ F_3(x, y, z) \end{pmatrix}$ , the divergence is given by

$$D_{\mathbf{F}}(\mathbf{r}) = \frac{\partial F_1(\mathbf{r})}{\partial x} + \frac{\partial F_2(\mathbf{r})}{\partial y} + \frac{\partial F_3(\mathbf{r})}{\partial z} \equiv \nabla \cdot \mathbf{F}(\mathbf{r}). \quad (196)$$

**Proof:** Exercises!

**Example:** Consider flowing  $\text{CO}_2$  gas in air. At any position  $\mathbf{r}$  and at any time  $t$ , let be:

- $\rho(\mathbf{r}, t)$  the mass density (in  $\frac{\text{g}}{\text{m}^3}$ ) of  $\text{CO}_2$ ,
- $\mathbf{J}(\mathbf{r}, t)$  the current density (in  $\frac{\text{g}}{\text{m}^2\text{s}}$ ) of the  $\text{CO}_2$  flow,
- $s(\mathbf{r}, t)$  the source density (in  $\frac{\text{g}}{\text{m}^3\text{s}}$ ) of chemical  $\text{CO}_2$  production.

Then, for any volume region  $\Omega$  (with surface  $\partial\Omega$ ), we have

$$\frac{d}{dt} \int_\Omega d^3r \rho(\mathbf{r}, t) = \int_\Omega d^3r s(\mathbf{r}, t) - \oint_{\partial\Omega} d\mathbf{A} \cdot \mathbf{J}(\mathbf{r}, t). \quad (197)$$

For the special case of **stationary flow**,  $f(\mathbf{r}, t) = f(\mathbf{r})$  where  $f \in \{\rho, \mathbf{J}, s\}$ , we conclude

$$\int_\Omega d^3r s(\mathbf{r}) = \oint_{\partial\Omega} d\mathbf{A} \cdot \mathbf{J}(\mathbf{r}). \quad (198)$$

Since this equation holds for any volume region  $\Omega \subseteq \mathbb{R}^3$ , Gauss' theorem implies

$$s(\mathbf{r}) = D_{\mathbf{J}}(\mathbf{r}) \equiv \nabla \cdot \mathbf{J}(\mathbf{r}). \quad (199)$$

Therefore, the divergence  $D_{\mathbf{F}}(\mathbf{r})$  of a vector field  $\mathbf{F}(\mathbf{r})$  is often called its **source density**.

## Alternative version of subsection 6.2

Consider flowing CO<sub>2</sub> gas in air. At any position  $\mathbf{r}$  and at any time  $t$ , let be:

- $\rho(\mathbf{r}, t)$  the mass density (in  $\frac{\text{g}}{\text{m}^3}$ ) of CO<sub>2</sub>,
- $\mathbf{J}(\mathbf{r}, t)$  the current density (in  $\frac{\text{g}}{\text{m}^2\text{s}}$ ) of the CO<sub>2</sub> flow,
- $s(\mathbf{r}, t)$  the source density (in  $\frac{\text{g}}{\text{m}^3\text{s}}$ ) of chemical CO<sub>2</sub> production.

Then, for any volume region  $\Omega$  with (closed) surface  $\partial\Omega$ , we have

$$\frac{d}{dt} \int_{\Omega} d^3r \rho(\mathbf{r}, t) = \int_{\Omega} d^3r s(\mathbf{r}, t) - \oint_{\partial\Omega} d\mathbf{A} \cdot \mathbf{J}(\mathbf{r}, t). \quad (200)$$

For the special case of **stationary flow**,  $f(\mathbf{r}, t) = f(\mathbf{r})$  where  $f \in \{\rho, \mathbf{J}, s\}$ , we obtain

$$\int_{\Omega} d^3r s(\mathbf{r}) = \oint_{\partial\Omega} d\mathbf{A} \cdot \mathbf{J}(\mathbf{r}). \quad (201)$$

Consequently, the scalar field  $s(\mathbf{r})$  is, in terms of the vector field  $\mathbf{J}(\mathbf{r})$ , given by

$$\begin{aligned} s(\mathbf{r}_0) &\equiv \lim_{V_{\Omega} \rightarrow 0} \langle s(\mathbf{r}) \rangle_{\mathbf{r} \in \Omega} \equiv \lim_{V_{\Omega} \rightarrow 0} \frac{1}{V_{\Omega}} \int_{\Omega} d^3r s(\mathbf{r}) \\ &= \lim_{V_{\Omega} \rightarrow 0} \frac{1}{V_{\Omega}} \oint_{\partial\Omega_a} d\mathbf{A} \cdot \mathbf{J}(\mathbf{r}), \end{aligned} \quad (202)$$

where  $V_{\Omega}$  is the volume of  $\Omega$  (with  $\mathbf{r}_0 \in \Omega$ ; SKETCH!). More precisely, we define

**Def.:** Let  $\Omega_a$  be the cube with center  $\mathbf{r}_0$  and edges of length  $a$  in  $x$ -,  $y$ -, and  $z$ -directions. The **divergence** of a vector field  $\mathbf{F}(\mathbf{r})$  is the scalar field  $D_{\mathbf{F}}(\mathbf{r})$ , given by

$$D_{\mathbf{F}}(\mathbf{r}_0) = \lim_{a \rightarrow 0} \frac{1}{a^3} \oint_{\partial\Omega_a} d\mathbf{A} \cdot \mathbf{F}(\mathbf{r}). \quad (203)$$

Remark 1: The number  $D_{\mathbf{F}}(\mathbf{r}_0)$  is not affected by a rotation of the  $x$ -,  $y$ -, and  $z$ -axes.

Remark 2: In usual notation,  $D_{\mathbf{F}}(\mathbf{r}) \equiv \text{div } \mathbf{F}(\mathbf{r})$ .

**Lemma:** For a differentiable vector field  $\mathbf{F}(\mathbf{r})$ , the divergence is given by

$$D_{\mathbf{F}}(\mathbf{r}) = \frac{\partial F_1(\mathbf{r})}{\partial x} + \frac{\partial F_2(\mathbf{r})}{\partial y} + \frac{\partial F_3(\mathbf{r})}{\partial z} \equiv \nabla \cdot \mathbf{F}(\mathbf{r}). \quad (204)$$

**Proof:** Exercises!

Consequently, during stationary flow, source and current densities are related by

$$s(\mathbf{r}) = D_{\mathbf{J}}(\mathbf{r}) \equiv \nabla \cdot \mathbf{J}(\mathbf{r}). \quad (205)$$

Therefore, the divergence  $D_{\mathbf{F}}(\mathbf{r})$  of a vector field  $\mathbf{F}(\mathbf{r})$  is often called its **source density**. Generalizing Eq. (201), one can show:

**Theorem (Gauss):** For a continuously differentiable vector field  $\mathbf{F}(\mathbf{r})$ , we have

$$\int_{\Omega} d^3r [\nabla \cdot \mathbf{F}(\mathbf{r})] = \oint_{\partial\Omega} d\mathbf{A} \cdot \mathbf{F}(\mathbf{r}), \quad (206)$$

for any compact volume region  $\Omega \subset \mathbb{R}^3$  with piecewise smooth boundary  $\partial\Omega$ . (Of course, the field  $\mathbf{F}$  must be defined on a neighborhood of  $\Omega$ .)

### 6.3 Curl of a vector field and Stokes' theorem

Let  $\Sigma$  be a piece of surface, e.g. a planar **disk** or a curved **saddle**. By  $\partial\Sigma$ , we denote the **closed curve** bordering  $\Sigma$  (rim or edge of  $\Sigma$ ). The line integral

$$\oint_{\partial\Sigma} d\mathbf{r} \cdot \mathbf{F}(\mathbf{r}) \quad (207)$$

is called the **circulation** of the vector field  $\mathbf{F}(\mathbf{r})$  around  $\Sigma$ . (To fix the sign of this integral, the orientation of  $\Sigma$  must be specified.)

While the **flux** of  $\mathbf{F}(\mathbf{r})$  out of some volume region  $\Omega$  is a measure for the tendency of the vectors  $\mathbf{F}(\mathbf{r})$  to “diverge away” from a point  $\mathbf{r}_0 \in \Omega$ , the **circulation** of  $\mathbf{F}(\mathbf{r})$  around a piece  $\Sigma$  of surface addresses the tendency of the vectors  $\mathbf{F}(\mathbf{r})$  to “rotate around”  $\mathbf{r}_0 \in \Sigma$ .

For illustration, let  $\mathbf{F}(\mathbf{r})$  be the velocity field  $\mathbf{v}(\mathbf{r})$  of a solid, rotating about the origin,

$$\mathbf{v}(\mathbf{r}) = \boldsymbol{\omega} \times \mathbf{r}. \quad (208)$$

$\mathbf{v}(\mathbf{r})$  is the velocity vector of a point with position vector  $\mathbf{r}$  on the solid. The vector  $\boldsymbol{\omega}$  has the direction of the axis of rotation (right-hand rule).  $|\boldsymbol{\omega}|$  is the angular velocity. Furthermore, let  $\Sigma$  be a disk centered at  $\mathbf{r}_0$  with radius  $R$  and unit normal vector  $\mathcal{N}_\Sigma$ .

- In the simplest case, when  $\mathcal{N}_\Sigma \parallel \boldsymbol{\omega}$  and  $\mathbf{r}_0$  lies on the axis, we obviously have

$$\oint_{\partial\Sigma} d\mathbf{r} \cdot \mathbf{v}(\mathbf{r}) = 2\pi R |\mathbf{v}(\mathbf{r})|_{\mathbf{r} \in \partial\Sigma} = 2\pi R |\boldsymbol{\omega}| R = \pi R^2 2|\boldsymbol{\omega}|. \quad (209)$$

- In the general case, we find (Exercise on Worksheet 6)

$$\oint_{\partial\Sigma} d\mathbf{r} \cdot \mathbf{v}(\mathbf{r}) = \pi R^2 (2\boldsymbol{\omega} \cdot \mathcal{N}_\Sigma). \quad (210)$$

In words: The circulation of the velocity field  $\mathbf{v}(\mathbf{r}) = \boldsymbol{\omega} \times \mathbf{r}$  around a disk  $\Sigma$  with unit normal vector  $\mathcal{N}_\Sigma$  equals the area  $A_\Sigma = \pi R^2$  of the disk times the scalar product  $2\boldsymbol{\omega} \cdot \mathcal{N}_\Sigma$ .

**Theorem:** For a vector field  $\mathbf{F}(\mathbf{r})$ , differentiable at  $\mathbf{r} = \mathbf{r}_0$ , there is a vector  $\boldsymbol{\omega}_\mathbf{F}(\mathbf{r}_0)$  with

$$2\boldsymbol{\omega}_\mathbf{F}(\mathbf{r}_0) \cdot \mathcal{N}_\Sigma = \lim_{R \rightarrow 0} \frac{1}{A_R} \oint_{\partial\Sigma_R} d\mathbf{r} \cdot \mathbf{F}(\mathbf{r}), \quad (211)$$

where  $\Sigma_R$  is the disk centered at  $\mathbf{r}_0$  with unit normal vector  $\mathcal{N}_\Sigma$ , radius  $R$  (and area  $A_\Sigma = \pi R^2$ ). In cartesian coordinates, the vector  $\boldsymbol{\omega}_\mathbf{F}(\mathbf{r})$  is given by

$$2\boldsymbol{\omega}_\mathbf{F}(\mathbf{r}) = \nabla \times \mathbf{F}(\mathbf{r}) \equiv \begin{pmatrix} \partial_2 F_3 - \partial_3 F_2 \\ \partial_3 F_1 - \partial_1 F_3 \\ \partial_1 F_2 - \partial_2 F_1 \end{pmatrix} \equiv \begin{pmatrix} \frac{\partial F_3(x,y,z)}{\partial y} - \frac{\partial F_2(x,y,z)}{\partial z} \\ \frac{\partial F_1(x,y,z)}{\partial z} - \frac{\partial F_3(x,y,z)}{\partial x} \\ \frac{\partial F_2(x,y,z)}{\partial x} - \frac{\partial F_1(x,y,z)}{\partial y} \end{pmatrix}. \quad (212)$$

[The vector field  $\nabla \times \mathbf{F}(\mathbf{r})$  is called the **curl**,  $\text{rot } \mathbf{F}(\mathbf{r})$ , of the vector field  $\mathbf{F}(\mathbf{r})$ .]

**Proof:** See section ??.

The “curl-analogy” to Gauss' divergence theorem reads:

**Stokes' theorem:** Let  $\Sigma$  be some piece of oriented surface with boundary (“rim”)  $\partial\Sigma$  and  $\mathbf{F}(\mathbf{r})$  a vector field, defined and differentiable for all  $\mathbf{r} \in \Sigma$ . Then,

$$\oint_{\partial\Sigma} d\mathbf{r} \cdot \mathbf{F}(\mathbf{r}) = \int_\Sigma d\mathbf{A} \cdot [\nabla \times \mathbf{F}(\mathbf{r})]. \quad (213)$$

### Alternative version of subsection 6.3

In section 6.2 we have considered the **flux**  $\Phi$  of a vector field  $\mathbf{F}(\mathbf{r})$ ,

$$\Phi = \oint_{\partial\Omega_a} d\mathbf{A} \cdot \mathbf{F}(\mathbf{r}), \quad (214)$$

out of the surface of a small **cube**  $\Omega_a$  with side length  $a$  and center  $\mathbf{r}_0$ .

- The number  $\Phi$  measures the tendency of  $\mathbf{F}(\mathbf{r})$  to "diverge" away from the point  $\mathbf{r}_0$ . In contrast, we now consider the **circulation**  $\Theta$  of  $\mathbf{F}(\mathbf{r})$  (a curve integral),

$$\Theta = \oint_{\partial\Sigma_a} d\mathbf{r} \cdot \mathbf{F}(\mathbf{r}), \quad (215)$$

along the rim of a small **planar square**  $\Sigma_a$  with side length  $a$  and center  $\mathbf{r}_0$ .

- The number  $\Theta$  measures the tendency of  $\mathbf{F}(\mathbf{r})$  to "rotate" around  $\mathbf{r}_0$  in the plane of  $\Sigma_a$ .

**Example:** For a given vector  $\boldsymbol{\omega}$ , the vector field

$$\mathbf{v}(\mathbf{r}) = \boldsymbol{\omega} \times \mathbf{r} \quad (216)$$

gives the velocity of a point (with instantaneous position  $\mathbf{r}$ ) on a solid, rotating with angular velocity  $\boldsymbol{\omega}$  around an axis through the origin.

Let  $\mathbf{a}_1$  and  $\mathbf{a}_2$  be two vectors with  $\mathbf{a}_1 \cdot \mathbf{a}_2 = 0$  and  $|\mathbf{a}_1| = |\mathbf{a}_2| = a$ . Then,

$$\mathbf{r}_{1,2,3,4} = \mathbf{r}_0 \pm \frac{\mathbf{a}_1}{2} \pm \frac{\mathbf{a}_2}{2}. \quad (217)$$

are the corners of a square  $\Sigma_a$  with center  $\mathbf{r}_0$ , side length  $a$ , unit normal vector  $\mathcal{N}_\Sigma = \frac{\mathbf{a}_1 \times \mathbf{a}_2}{a^2}$ . In the exercises, we shall show

$$\frac{\Theta}{a^2} = \frac{1}{a^2} \oint_{\partial\Sigma_a} d\mathbf{r} \cdot \mathbf{v}(\mathbf{r}) = \mathcal{N}_\Sigma \cdot 2\boldsymbol{\omega}. \quad (218)$$

**Def.:** Let  $\Sigma_a$  be a square with center  $\mathbf{r}_0$ , side length  $a$ , and unit normal vector  $\mathcal{N}_\Sigma$ . The **curl** of a vector field  $\mathbf{F}(\mathbf{r})$  is the vector field  $\mathbf{C}_\mathbf{F}(\mathbf{r})$  with the property

$$\mathcal{N}_\Sigma \cdot \mathbf{C}_\mathbf{F}(\mathbf{r}_0) = \lim_{a \rightarrow 0} \frac{1}{a^2} \oint_{\partial\Sigma_a} d\mathbf{r} \cdot \mathbf{F}(\mathbf{r}). \quad (219)$$

Remark 1: The number  $\mathcal{N}_\Sigma \cdot \mathbf{C}_\mathbf{F}(\mathbf{r}_0)$  is not affected by a rotation of the  $x$ -,  $y$ -, and  $z$ -axes.

Remark 2: In usual notation,  $\mathbf{C}_\mathbf{F}(\mathbf{r}) \equiv \text{rot } \mathbf{F}(\mathbf{r})$ .

**Lemma:** For a differentiable vector field  $\mathbf{F}(\mathbf{r})$ , the curl is given by

$$\mathbf{C}_\mathbf{F}(\mathbf{r}) = \begin{pmatrix} \frac{\partial F_3(\mathbf{r})}{\partial y} - \frac{\partial F_2(\mathbf{r})}{\partial z} \\ \frac{\partial F_1(\mathbf{r})}{\partial z} - \frac{\partial F_3(\mathbf{r})}{\partial x} \\ \frac{\partial F_2(\mathbf{r})}{\partial x} - \frac{\partial F_1(\mathbf{r})}{\partial y} \end{pmatrix} \equiv \begin{pmatrix} \partial_2 F_3 - \partial_3 F_2 \\ \partial_3 F_1 - \partial_1 F_3 \\ \partial_1 F_2 - \partial_2 F_1 \end{pmatrix} \equiv \nabla \times \mathbf{F}(\mathbf{r}). \quad (220)$$

**Proof:** Exercises!

**Example:** For the velocity field of the rotating solid, we find in fact

$$\nabla \times \mathbf{v}(\mathbf{r}) \equiv \nabla \times (\boldsymbol{\omega} \times \mathbf{r}) = \begin{pmatrix} \partial_x \\ \partial_y \\ \partial_z \end{pmatrix} \times \begin{pmatrix} \omega_2 z - \omega_3 y \\ \omega_3 x - \omega_1 z \\ \omega_1 y - \omega_2 x \end{pmatrix} = 2\boldsymbol{\omega}. \quad (221)$$

## 6.4 Curl and independence of path

**Def.:** A volume region  $\Omega$  is called **simply connected**, when every closed loop  $\Gamma$  in  $\Omega$  can be shrunk continuously within  $\Omega$  to a point  $\mathbf{r}_0 \in \Omega$ .

(Then, there is a piece of surface  $\Sigma$  with:  $\Gamma = \partial\Sigma$  and  $\Sigma \subset \Omega$ .)

**Example:**  $\Omega = \{(x, y, z) \mid x^2 + y^2 \geq R^2\} \subset \mathbb{R}^3$  is **not** simply connected when  $R > 0$ .

**Theorem:** Let  $\mathbf{F}(\mathbf{r})$  be a differentiable vector field, defined on a simply connected region  $\Omega \subseteq \mathbb{R}^3$ . Then, the following three statements are equivalent:

- (a) There is a scalar field  $f(\mathbf{r})$  with the property  $\mathbf{F}(\mathbf{r}) = \nabla f(\mathbf{r})$  for all  $\mathbf{r} \in \Omega$ .
- (b)  $\nabla \times \mathbf{F}(\mathbf{r}) = \mathbf{0}$  for all  $\mathbf{r} \in \Omega$ .
- (c) For a curve  $\Gamma \in \Omega$ , the value of the line integral  $\int_{\Gamma} d\mathbf{r} \cdot \mathbf{F}(\mathbf{r})$  only depends on the initial and final points  $\mathbf{r}_A$  and  $\mathbf{r}_B$ , respectively, of  $\Gamma$ , but **not on its path**.

**Proof:** • From section 5.1, we already know that (a)  $\Rightarrow$  (b).

[We also know that (a)  $\Rightarrow$  (c), since  $\int_{\Gamma} d\mathbf{r} \cdot \nabla f(\mathbf{r}) \equiv f(\mathbf{r}_B) - f(\mathbf{r}_A)$ , see section 5.2.]

• The conclusion (b)  $\Rightarrow$  (c) follows immediately from Stokes' Theorem: Given two curves  $\Gamma_1$  and  $\Gamma_2$  which have the same initial point  $\mathbf{r}_A$  and the same final point  $\mathbf{r}_B$ , let  $\bar{\Gamma}_2$  follow the same path as  $\Gamma_2$ , but in opposite direction (from  $\mathbf{r}_B$  to  $\mathbf{r}_A$ ). Then, the composition  $\bar{\Gamma}_2 \circ \Gamma_1$  makes a closed loop  $\Gamma = \partial\Sigma$  in  $\Omega$  (starting at  $\mathbf{r}_A$ , following  $\Gamma_1$  up to  $\mathbf{r}_B$  and returning along the path of  $\Gamma_2$  to  $\mathbf{r}_A$ ). Therefore, since  $\nabla \times \mathbf{F}(\mathbf{r}) = \mathbf{0}$  for  $\mathbf{r} \in \Sigma \subset \Omega$ ,

$$\int_{\Gamma_1} d\mathbf{r} \cdot \mathbf{F}(\mathbf{r}) - \int_{\Gamma_2} d\mathbf{r} \cdot \mathbf{F}(\mathbf{r}) \equiv \oint_{\partial\Sigma} d\mathbf{r} \cdot \mathbf{F}(\mathbf{r}) = \int_{\Sigma} d\mathbf{A} \cdot [\nabla \times \mathbf{F}(\mathbf{r})] = 0. \quad (222)$$

• To prove the conclusion (c)  $\Rightarrow$  (a) we define for any  $\mathbf{r} \in \Omega$  the scalar field

$$f(\mathbf{r}) := \int_{\Gamma(\mathbf{r}_0, \mathbf{r})} d\mathbf{r}' \cdot \mathbf{F}(\mathbf{r}'), \quad (223)$$

where  $\Gamma(\mathbf{r}_0, \mathbf{r})$  is **any** curve in  $\Omega$  running from some fixed point  $\mathbf{r}_0 \in \Omega$  to  $\mathbf{r}$ . For  $h \geq 0$ , let  $\Gamma(\mathbf{r}_0, \mathbf{r}, h)$  be the extension of  $\Gamma(\mathbf{r}_0, \mathbf{r})$  by a straight line  $\Gamma_h$ , running in  $x$ -direction from  $\mathbf{r}$  to the new endpoint  $\mathbf{r} + h\mathbf{u}_x$  (where  $\mathbf{u}_x$  is the unit vector in  $x$ -direction). Then,

$$\begin{aligned} \frac{\partial}{\partial x} f(x, y, z) &\equiv \lim_{h \rightarrow 0} \frac{1}{h} \left[ f(x+h, y, z) - f(x, y, z) \right] \\ &= \lim_{h \rightarrow 0} \frac{1}{h} \left[ \int_{\Gamma(\mathbf{r}_0, \mathbf{r}, h)} d\mathbf{r}' \cdot \mathbf{F}(\mathbf{r}') - \int_{\Gamma(\mathbf{r}_0, \mathbf{r})} d\mathbf{r}' \cdot \mathbf{F}(\mathbf{r}') \right] = \lim_{h \rightarrow 0} \frac{1}{h} \int_{\Gamma_h} d\mathbf{r}' \cdot \mathbf{F}(\mathbf{r}'). \end{aligned} \quad (224)$$

Using the parametrisation

$$\Gamma_h = \left\{ \mathbf{r}'(s) \mid 0 \leq s \leq h \right\}, \quad \mathbf{r}'(s) = \mathbf{r} + s\mathbf{u}_x, \quad \frac{d\mathbf{r}'(s)}{ds} = \mathbf{u}_x, \quad (225)$$

we have  $\int_{\Gamma_h} d\mathbf{r}' \cdot \mathbf{F}(\mathbf{r}') \equiv \int_0^h ds \frac{d\mathbf{r}'(s)}{ds} \cdot \mathbf{F}(\mathbf{r}'(s)) = \int_0^h ds F_1(x+s, y, z)$  and, eventually,

$$\frac{\partial}{\partial x} f(x, y, z) = \lim_{h \rightarrow 0} \frac{1}{h} \int_0^h ds F_1(x+s, y, z) = F_1(x, y, z), \quad \text{q.e.d.} \quad (226)$$

## 7 Application of vector calculus: Electrodynamics

### 7.1 Electric charge $q$

We consider electric charge as a continuum with density  $\rho(\mathbf{r}, t)$  (in  $\frac{\text{C}}{\text{m}^3}$ ) and current density  $\mathbf{J}(\mathbf{r}, t)$  (in  $\frac{\text{C}}{\text{m}^2\text{s}}$ ). In such a description, the case of a **point charge** can be included, when we introduce Dirac's  $\delta$ -function  $\delta(\mathbf{r} - \mathbf{r}_0) = \delta(x - x_0)\delta(y - y_0)\delta(z - z_0)$ : A moving point charge  $q$ , with time-dependent position vector  $\mathbf{r}_0(t)$ , has the charge density

$$\rho(\mathbf{r}, t) = q \delta(\mathbf{r} - \mathbf{r}_0(t)). \quad (227)$$

Similarly, a thin wire (covering the  $z$ -axis) with electric current  $I$  has the current density

$$\mathbf{J}(\mathbf{r}, t) = I \delta(x)\delta(y) \mathbf{e}_z. \quad (228)$$

Unlike a chemical compound (such as  $\text{CO}_2$  in our former examples), electric charge is a **conserved quantity**: It can (chemically) neither be created nor destroyed. Consequently, the source density of charge is  $s(\mathbf{r}, t) = 0$  and the amount  $\int_{\Omega} d^3r \rho(\mathbf{r}, t)$  of charge within any volume region  $\Omega$  can change only by a flow of charge across the surface  $\partial\Omega$ ,

$$\frac{d}{dt} \int_{\Omega} d^3r \rho(\mathbf{r}, t) = - \oint_{\partial\Omega} d\mathbf{S} \cdot \mathbf{J}(\mathbf{r}, t). \quad (229)$$

Since this equation holds for any  $\Omega \subset \mathbb{R}^3$ , Gauss' theorem  $\oint_{\partial\Omega} d\mathbf{S} \cdot \mathbf{J} = \int_{\Omega} d^3r \nabla \cdot \mathbf{J}$  implies

$$\frac{\partial}{\partial t} \rho(\mathbf{r}, t) + \nabla \cdot \mathbf{J}(\mathbf{r}, t) = 0. \quad (230)$$

This differential relation between the scalar field  $\rho(\mathbf{r}, t)$  and the vector field  $\mathbf{J}(\mathbf{r}, t)$  is called **continuity equation**. It is the mathematical expression for a **conservation law**.

### 7.2 Electromagnetic fields $\mathbf{E}$ and $\mathbf{B}$

In space, let  $\mathbf{F}(\mathbf{r}, \mathbf{v}, t)$  be the total force acting on a test (point) charge  $q$  at the time  $t$  when it has position  $\mathbf{r}$  and is moving with velocity  $\mathbf{v}$ . If  $\mathbf{F}$  is purely an electromagnetic force, then there exist (time dependent) vector fields  $\mathbf{E}(\mathbf{r}, t)$  and  $\mathbf{B}(\mathbf{r}, t)$ , called **electric** and **magnetic field**, respectively, such that

$$\mathbf{F}(\mathbf{r}, \mathbf{v}, t) = q \left[ \mathbf{E}(\mathbf{r}, t) + \mathbf{v} \times \mathbf{B}(\mathbf{r}, t) \right]. \quad (231)$$

Consequently, the fields have the physical dimensions<sup>3</sup>

$$[\mathbf{E}] = 1 \frac{\text{N}}{\text{C}}, \quad [\mathbf{B}] = 1 \frac{\text{Ns}}{\text{Cm}} \equiv 1\text{T} \quad (1 \text{ Tesla}). \quad (232)$$

Due to Eq. (231), the vectors  $\mathbf{E}(\mathbf{r}, t)$  and  $\mathbf{B}(\mathbf{r}, t)$ , at any position-time pair  $(\mathbf{r}, t)$ , can be measured by observing the forces acting on various moving test charges  $q$ .

<sup>3</sup>Giving  $\mathbf{B}$  the same dimension  $\text{N/C}$  as  $\mathbf{E}$ , one may define alternatively

$$\mathbf{F} = q \mathbf{E}(\mathbf{r}, t) + q \frac{\mathbf{v}}{c} \times \mathbf{B}(\mathbf{r}, t),$$

where  $c$  is the speed of light.

### 7.3 Stationary fields $\mathbf{E}(\mathbf{r})$ , $\mathbf{B}(\mathbf{r})$

In situations when the fields  $\mathbf{E}$  and  $\mathbf{B}$  are time-independent ("stationary"), they seem to be "generated" by (stationary) charges  $\rho(\mathbf{r})$  and, respectively, (stationary) currents  $\mathbf{J}(\mathbf{r})$ . This relationship between the fields and their "sources" is governed by the constants

$$\epsilon_0 = 8.854 \cdot 10^{-12} \frac{\text{As}}{\text{Vm}}, \quad \mu_0 = 4\pi \cdot 10^{-7} \frac{\text{Vs}}{\text{Am}}. \quad (233)$$

#### 7.3.1 Coulomb's law

A time-independent charge distribution  $\rho(\mathbf{r})$  gives rise to a stationary electric field,

$$\mathbf{E}(\mathbf{r}) = \frac{1}{4\pi\epsilon_0} \int d^3r' \rho(\mathbf{r}') \frac{\mathbf{r} - \mathbf{r}'}{|\mathbf{r} - \mathbf{r}'|^3}, \quad (234)$$

seen by a (test) charge at position  $\mathbf{r}$ . In this context, the "generating" charge  $\rho(\mathbf{r})$  is also called "source" charge. In the special case of a point source,  $\rho(\mathbf{r}) = Q\delta(\mathbf{r} - \mathbf{r}_0)$ , we find

$$\mathbf{E}(\mathbf{r}) = \frac{Q}{4\pi\epsilon_0} \frac{\mathbf{r} - \mathbf{r}_0}{|\mathbf{r} - \mathbf{r}_0|^3} \quad \Rightarrow \quad |\mathbf{E}(\mathbf{r})| = \frac{|Q|}{4\pi\epsilon_0} \frac{1}{|\mathbf{r} - \mathbf{r}_0|^2}. \quad (235)$$

Usually, Eq. (235) for a point source is called **Coulomb's law**. Then, Eq. (234) results from the **superposition principle**.

#### 7.3.2 Biot-Savart law

A time-independent current distribution  $\mathbf{J}(\mathbf{r})$  gives rise to a stationary magnetic field,

$$\mathbf{B}(\mathbf{r}) = \frac{\mu_0}{4\pi} \int d^3r' \mathbf{J}(\mathbf{r}') \times \frac{\mathbf{r} - \mathbf{r}'}{|\mathbf{r} - \mathbf{r}'|^3}, \quad (236)$$

seen by a (moving) test charge at position  $\mathbf{r}$ . In the special case, when the source current is an infinitely long straight wire, carrying the current  $I$ ,  $\mathbf{J}(\mathbf{r}) = I\delta(x)\delta(y)\mathbf{e}_z$ , we find

$$\begin{aligned} \mathbf{B}(\mathbf{r}) &= \frac{\mu_0}{4\pi} I \int_{-\infty}^{\infty} dz' \frac{\begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix} \times \begin{pmatrix} x \\ y \\ z - z' \end{pmatrix}}{[x^2 + y^2 + (z - z')^2]^{3/2}} \\ &= \frac{\mu_0}{4\pi} I \begin{pmatrix} -y \\ x \\ 0 \end{pmatrix} \int_{-\infty}^{\infty} dz' [x^2 + y^2 + (z - z')^2]^{-3/2} \\ &= \frac{\mu_0}{2\pi} I \begin{pmatrix} -y \\ x \\ 0 \end{pmatrix} \frac{1}{x^2 + y^2}, \end{aligned} \quad (237)$$

where we have used  $\int_{-\infty}^{\infty} du [a^2 + u^2]^{-3/2} = \frac{u}{a^2} [a^2 + u^2]^{-1/2}$ .

$$|\mathbf{B}(\mathbf{r})| = \frac{\mu_0 I}{2\pi} \frac{1}{|\mathbf{r}_\perp|}. \quad (238)$$

### 7.3.3 Some mathematical identities (problems from Worksheet 7/2019)

**Problem 4: The scalar field**  $\phi(\mathbf{r}) = \frac{1}{|\mathbf{r}-\mathbf{r}_0|}$ .

Use the explicit representation of  $|\mathbf{r}-\mathbf{r}_0|$  in cartesian coordinates,

$$|\mathbf{r}-\mathbf{r}_0| = \left[ (x-x_0)^2 + (y-y_0)^2 + (z-z_0)^2 \right]^{1/2} \equiv f(x, y, z),$$

to prove the identities

$$\nabla \frac{1}{|\mathbf{r}-\mathbf{r}_0|} = -\frac{\mathbf{r}-\mathbf{r}_0}{|\mathbf{r}-\mathbf{r}_0|^3}, \quad \nabla \times \frac{\mathbf{a}}{|\mathbf{r}-\mathbf{r}_0|} = \frac{\mathbf{a} \times (\mathbf{r}-\mathbf{r}_0)}{|\mathbf{r}-\mathbf{r}_0|^3}, \quad (239)$$

where the differential operator  $\nabla$  is understood to act on the variables  $\mathbf{r} = (x, y, z)$ , while  $\mathbf{r}_0 = (x_0, y_0, z_0)$  and  $\mathbf{a} = (a_1, a_2, a_3)$  are constant vectors.

**Problem 5: Laplacian (operator)  $\nabla^2$**

(a) For a differentiable scalar field  $g(\mathbf{r}) = g(x, y, z)$ , show that

$$\nabla \cdot [\nabla g(\mathbf{r})] = \nabla^2 g(\mathbf{r}),$$

where (in cartesian coordinates) the **Laplacian**  $\nabla^2$  reads  $\nabla^2 = \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2}$ .

(b) Evaluate the Laplacian  $\nabla^2 g(\mathbf{r})$  for the following scalar fields (where  $A, B, C, D, E, F, R, \alpha, \beta$  are constants).

$$\begin{aligned} g_1(\mathbf{r}) &= Ax^2 + By^2 - (A+B)z^2 + Dx + Ey + Fz + C, \\ g_2(\mathbf{r}) &= A \sin(\alpha x) \sin(\beta y) e^{-\gamma z}, \quad \gamma = \sqrt{\alpha^2 + \beta^2}, \\ g_3(\mathbf{r}) &= \frac{A}{\sqrt{x^2 + y^2 + z^2}} \equiv A(x^2 + y^2 + z^2)^{-1/2} \quad (\mathbf{r} \neq \mathbf{0}), \\ g_4(\mathbf{r}) &= (Ax + By + Cz)(x^2 + y^2 + z^2)^{-3/2} \quad (\mathbf{r} \neq \mathbf{0}), \\ g_5(\mathbf{r}) &= \frac{Q}{4\pi\epsilon_0} \begin{cases} \frac{3R^2 - r^2}{2R^3} & (r \leq R), \\ \frac{1}{r} & (r > R). \end{cases} \end{aligned}$$

(c) Use the result of part (b) for  $g_5(\mathbf{r})$ , to verify the identity

$$\nabla^2 \frac{1}{|\mathbf{r}-\mathbf{r}_0|} = -4\pi \delta(\mathbf{r}-\mathbf{r}_0). \quad (240)$$

**Problem 6: Second derivatives and integration by parts**

(a) For a differentiable scalar field  $g(\mathbf{r})$  and a differentiable vector field  $\mathbf{F}(\mathbf{r})$ , show that

$$\nabla \cdot [\nabla g(\mathbf{r})] = \nabla^2 g(\mathbf{r}) \quad (\text{already shown in problem 5(a)}), \quad (241)$$

$$\nabla \times [\nabla g(\mathbf{r})] = \mathbf{0}, \quad (242)$$

$$\nabla \cdot [\nabla \times \mathbf{F}(\mathbf{r})] = 0, \quad (243)$$

$$\nabla \times [\nabla \times \mathbf{F}(\mathbf{r})] = \nabla[\nabla \cdot \mathbf{F}(\mathbf{r})] - \nabla^2 \mathbf{F}(\mathbf{r}). \quad (244)$$

(b) Integrating by parts, show that

$$\int d^3r \mathbf{F}(\mathbf{r}) \cdot [\nabla g(\mathbf{r})] = - \int d^3r [\nabla \cdot \mathbf{F}(\mathbf{r})] g(\mathbf{r}), \quad (245)$$

provided that the product  $\mathbf{P}(\mathbf{r}) = g(\mathbf{r}) \mathbf{F}(\mathbf{r})$  goes to zero as  $|\mathbf{r}| \rightarrow \infty$ .

### 7.3.4 Scalar and vector potentials $\phi$ and $\mathbf{A}$

Defining the so-called (electrostatic) **scalar potential**

$$\phi(\mathbf{r}) = \frac{1}{4\pi\epsilon_0} \int d^3r' \frac{\rho(\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|}, \quad (246)$$

and the so-called (electromagnetic) **vector potential**

$$\mathbf{A}(\mathbf{r}) = \frac{\mu_0}{4\pi} \int d^3r' \frac{\mathbf{J}(\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|}, \quad (247)$$

and applying the identities (239), the integrals (234) and (236) show that

$$\mathbf{E}(\mathbf{r}) = -\nabla\phi(\mathbf{r}), \quad \mathbf{B}(\mathbf{r}) = \nabla \times \mathbf{A}(\mathbf{r}). \quad (248)$$

**Maxwell's equations** for stationary fields,

$$(I) \quad \nabla \cdot \mathbf{E}(\mathbf{r}) = \frac{1}{\epsilon_0} \rho(\mathbf{r}), \quad (II) \quad \nabla \cdot \mathbf{B}(\mathbf{r}) = 0, \quad (249)$$

$$(III) \quad \nabla \times \mathbf{B}(\mathbf{r}) = \mu_0 \mathbf{J}(\mathbf{r}), \quad (IV) \quad \nabla \times \mathbf{E}(\mathbf{r}) = \mathbf{0}, \quad (250)$$

can be derived from the representations (248) as follows:

- (II) and (IV) are immediate consequences of (248) due to Eqs. (242) and (243).
- (I) is obtained from (248) as follows:

$$\nabla \cdot \mathbf{E}(\mathbf{r}) \equiv -\nabla^2\phi(\mathbf{r}) = -\frac{1}{4\pi\epsilon_0} \int d^3r' \rho(\mathbf{r}') \nabla^2 \frac{1}{|\mathbf{r} - \mathbf{r}'|} = \frac{1}{\epsilon_0} \rho(\mathbf{r}), \quad (251)$$

where we have used the famous identity Eq. (240) in the last step.

- To derive Equation (III) from (248), we first notice that

$$\nabla \times \mathbf{B}(\mathbf{r}) \equiv \nabla \times [\nabla \times \mathbf{A}(\mathbf{r})] = \nabla[\nabla \cdot \mathbf{A}(\mathbf{r})] - \nabla^2 \mathbf{A}(\mathbf{r}), \quad (252)$$

due to the identity (244). Here, we have

$$\nabla \cdot \mathbf{A}(\mathbf{r}) = \frac{\mu_0}{4\pi} \int d^3r' \mathbf{J}(\mathbf{r}') \cdot \nabla_{\mathbf{r}} \frac{1}{|\mathbf{r} - \mathbf{r}'|} \equiv -\frac{\mu_0}{4\pi} \int d^3r' \mathbf{J}(\mathbf{r}') \cdot \nabla_{\mathbf{r}'} \frac{1}{|\mathbf{r} - \mathbf{r}'|}, \quad (253)$$

and integration by parts, see Eq. (245), yields

$$\nabla \cdot \mathbf{A}(\mathbf{r}) = +\frac{\mu_0}{4\pi} \int d^3r' [\nabla_{\mathbf{r}'} \cdot \mathbf{J}(\mathbf{r}')] \frac{1}{|\mathbf{r} - \mathbf{r}'|} = 0, \quad (254)$$

where we have used the continuity equation  $\nabla \cdot \mathbf{J}(\mathbf{r}) = -\frac{\partial\rho(\mathbf{r})}{\partial t} = 0$ . Consequently,

$$\nabla \times \mathbf{B}(\mathbf{r}) = -\nabla^2 \mathbf{A}(\mathbf{r}) = -\frac{\mu_0}{4\pi} \int d^3r' \mathbf{J}(\mathbf{r}') \nabla^2 \frac{1}{|\mathbf{r} - \mathbf{r}'|} = \mu_0 \mathbf{J}(\mathbf{r}), \quad (255)$$

where we have again used the identity Eq. (240) in the last step.

It seems that the divergence of  $\mathbf{E}$  is the charge density  $\rho$  and the curls of  $\mathbf{B}$  are the currents  $\mathbf{J}$ , while  $\mathbf{E}$  has no curls and  $\mathbf{B}$  has zero divergence. In the following section 7.4, however, we shall switch to **time-dependent** fields  $\mathbf{E}(\mathbf{r}, t)$  and  $\mathbf{B}(\mathbf{r}, t)$ , revealing that  $\frac{\partial\mathbf{B}}{\partial t}$  produces curls of  $\mathbf{E}$  and  $\frac{\partial\mathbf{E}}{\partial t}$  produces (additional) curls of  $\mathbf{B}$ .

## 7.4 Maxwell's equations

### 7.4.1 Differential form

Generally, along with the electric densities  $\rho(\mathbf{r}, t)$  and  $\mathbf{J}(\mathbf{r}, t)$  of charge and current, the time-dependent fields  $\mathbf{E}(\mathbf{r}, t)$  and  $\mathbf{B}(\mathbf{r}, t)$  satisfy **Maxwell's equations**,

$$\nabla \cdot \mathbf{E}(\mathbf{r}, t) = \frac{1}{\epsilon_0} \rho(\mathbf{r}, t), \quad \nabla \cdot \mathbf{B}(\mathbf{r}, t) = 0, \quad (256)$$

$$\nabla \times \mathbf{B}(\mathbf{r}, t) = \mu_0 \mathbf{J}(\mathbf{r}, t) + \epsilon_0 \mu_0 \frac{\partial \mathbf{E}(\mathbf{r}, t)}{\partial t}, \quad \nabla \times \mathbf{E}(\mathbf{r}, t) = -\frac{\partial \mathbf{B}(\mathbf{r}, t)}{\partial t}. \quad (257)$$

**Remark:** The extra term  $\epsilon_0 \mu_0 \frac{\partial \mathbf{E}(\mathbf{r}, t)}{\partial t}$  makes  $\rho$  and  $\mathbf{J}$  satisfy the continuity equation,

$$\frac{1}{\mu_0} \nabla \cdot \left[ \nabla \times \mathbf{B}(\mathbf{r}, t) \right] \equiv \nabla \cdot \mathbf{J}(\mathbf{r}, t) + \frac{\partial}{\partial t} \rho(\mathbf{r}, t) = 0. \quad (258)$$

The first one of Eqs. (256) is called **Gauss' Law**.

Eqs. (257) are called **Ampere's law** and **Faraday's law**, respectively.

### 7.4.2 Integral form

Applying Gauss' theorem  $\oint_{\partial\Omega} d\mathbf{A} \cdot \mathbf{F}(\mathbf{r}) = \int_{\Omega} d^3r [\nabla \cdot \mathbf{F}(\mathbf{r})]$  to Maxwell's equations (256), we obtain the integral relations

$$\begin{aligned} \text{(I)} \quad \oint_{\partial\Omega} d\mathbf{A} \cdot \mathbf{E}(\mathbf{r}, t) &= \frac{1}{\epsilon_0} \int_{\Omega} d^3r \rho(\mathbf{r}, t) \equiv \frac{Q_{\Omega}(t)}{\epsilon_0}, \\ \text{(II)} \quad \oint_{\partial\Omega} d\mathbf{A} \cdot \mathbf{B}(\mathbf{r}, t) &= 0. \end{aligned} \quad (259)$$

In words, (I): *The flux of the electric field out of the surface  $\partial\Omega$  of a volume region  $\Omega$  always equals ( $\frac{1}{\epsilon_0}$  times) the amount  $Q_{\Omega}(t)$  of charge contained within  $\Omega$  at the same time.*  
 (II): *There is no magnetic flux out of closed surfaces. The field  $\mathbf{B}(\mathbf{r}, t)$  has no sources.*

Similarly, applying Stokes' theorem  $\oint_{\partial\Sigma} d\mathbf{r} \cdot \mathbf{F}(\mathbf{r}) = \int_{\Sigma} d\mathbf{A} \cdot [\nabla \times \mathbf{F}(\mathbf{r})]$  to Maxwell's equations (257), we obtain the integral relations

$$\text{(III)} \quad \oint_{\partial\Sigma} d\mathbf{r} \cdot \mathbf{B}(\mathbf{r}, t) = \mu_0 \int_{\Sigma} d\mathbf{A} \cdot \left[ \mathbf{J}(\mathbf{r}, t) + \epsilon_0 \frac{\partial \mathbf{E}(\mathbf{r}, t)}{\partial t} \right], \quad (260)$$

$$\text{(IV)} \quad \oint_{\partial\Sigma} d\mathbf{r} \cdot \mathbf{E}(\mathbf{r}, t) = -\frac{d}{dt} \int_{\Sigma} d\mathbf{A} \cdot \mathbf{B}(\mathbf{r}, t). \quad (261)$$

(IV) is **Faraday's induction law**: *The circulation of the electric field around the rim  $\partial\Sigma$  of any piece  $\Sigma$  of surface, called the **induction voltage**  $U_{\text{ind}}(t)$  at the time  $t$ , always equals minus the rate of change of the magnetic flux  $\Phi(t)$  through  $\Sigma$  at that time,*

$$U_{\text{ind}}(t) = -\frac{d}{dt} \Phi(t). \quad (262)$$

In the absence of a time-dependent electric field, Eq. (III) becomes

$$\oint_{\partial\Sigma} d\mathbf{r} \cdot \mathbf{B}(\mathbf{r}, t) = \mu_0 \int_{\Sigma} d\mathbf{A} \cdot \mathbf{J}(\mathbf{r}, t) \equiv \mu_0 I_{\Sigma}(t). \quad (263)$$

In words: *The circulation of the magnetic field around the rim  $\partial\Sigma$  at the time  $t$  equals ( $\mu_0$  times) the total electric current  $I_{\Sigma}(t)$  flowing through  $\Sigma$  at that time (**Ampere's law**).*

### 7.4.3 Derivation of the laws of Coulomb and Biot-Savart

A uniformly charged sphere with charge  $Q$  and finite radius  $R_0$  has the charge density

$$\rho(r, \theta, \phi) = \begin{cases} \rho_0 & (r \leq R_0) \\ 0 & (r > R_0) \end{cases}, \quad \rho_0 = \frac{Q}{\frac{4\pi}{3}R_0^3}. \quad (264)$$

Due to the spherical symmetry of  $\rho$ , the electric field has the form

$$\mathbf{E}(\mathbf{r}) = E(r) \frac{\mathbf{r}}{r} \equiv E(r) \mathbf{u}_r, \quad (265)$$

with a yet unknown  $r$ -dependent magnitude  $E(r)$ . Let  $\Omega$  be a concentric sphere with radius  $R$ . Then, Maxwell's equation  $\oint_{\partial\Omega} d\mathbf{A} \cdot \mathbf{E}(\mathbf{r}) = \frac{1}{\epsilon_0} \int_{\Omega} d^3r \rho(\mathbf{r})$  has the LHS

$$\oint_{\partial\Omega} d\mathbf{A} \cdot \mathbf{E}(\mathbf{r}) \equiv A_{\partial\Omega} \left\langle \mathcal{N}(\mathbf{r}) \cdot \mathbf{E}(\mathbf{r}) \right\rangle_{\mathbf{r} \in \partial\Omega} = 4\pi R^2 E(R), \quad (266)$$

where we have utilised that  $A_{\partial\Omega} = 4\pi R^2$  and  $\mathcal{N}(\mathbf{r}) = \mathbf{u}_r$ . On the RHS, we have

$$\frac{1}{\epsilon_0} \int_{\Omega} d^3r \rho(\mathbf{r}) = \frac{1}{\epsilon_0} \begin{cases} V_{\Omega} \rho_0 \equiv Q \frac{R^3}{R_0^3} & (R \leq R_0), \\ Q & (R \geq R_0), \end{cases} \quad (267)$$

since  $V_{\Omega} = \frac{4\pi}{3}R^3$ . Comparing the last two results (and setting  $R = r$ ), we obtain

$$E(r) = \frac{1}{4\pi\epsilon_0} Q \begin{cases} \frac{r}{R_0^3} & (r \leq R_0), \\ \frac{1}{r^2} & (r \geq R_0). \end{cases} \quad (268)$$

Coulomb's law (235) for a point charge  $Q$  is recovered in the limit  $R_0 \rightarrow 0$ .

A uniform current  $I$  inside an infinitely long cylinder with radius  $R_0$  has the density

$$\mathbf{J}(s, \phi, z) = \begin{cases} J_0 \mathbf{u}_z & (s \leq R_0) \\ 0 & (s > R_0) \end{cases}, \quad J_0 = \frac{I}{\pi R_0^2}. \quad (269)$$

Due to the cylindrical symmetry of  $\mathbf{J}$ , the magnetic field has the form

$$\mathbf{B}(\mathbf{r}) = B(s) \frac{\mathbf{u}_z \times \mathbf{r}}{|\mathbf{u}_z \times \mathbf{r}|} \equiv B(s) \mathbf{u}_{\phi}, \quad (270)$$

with a yet unknown  $s$ -dependent magnitude  $B(s)$ . Let  $\Sigma$  be a disk with radius  $R$ , centered on the  $z$ -axis and normal to it. Then, Maxwell's equation  $\oint_{\partial\Sigma} d\mathbf{r} \cdot \mathbf{B}(\mathbf{r}) = \mu_0 \int_{\Sigma} d\mathbf{A} \cdot \mathbf{J}(\mathbf{r})$  has the LHS

$$\oint_{\partial\Sigma} d\mathbf{r} \cdot \mathbf{B}(\mathbf{r}) \equiv L_{\partial\Sigma} \left\langle \mathcal{T}(\mathbf{r}) \cdot \mathbf{B}(\mathbf{r}) \right\rangle_{\mathbf{r} \in \partial\Sigma} = 2\pi R B(R), \quad (271)$$

where we have utilised that  $L_{\partial\Sigma} = 2\pi R$  and  $\mathcal{T}(\mathbf{r}) = \mathbf{u}_{\phi}$ . On the RHS, we have

$$\mu_0 \int_{\Sigma} d\mathbf{A} \cdot \mathbf{J}(\mathbf{r}) = \mu_0 \begin{cases} A_{\Sigma} J_0 \equiv I \frac{R^2}{R_0^2} & (R \leq R_0), \\ I & (R \geq R_0), \end{cases} \quad (272)$$

since  $A_{\Sigma} = \pi R^2$ . Comparing the last two results (and setting  $R = s$ ), we obtain

$$B(s) = \frac{\mu_0}{2\pi} I \begin{cases} \frac{s}{R_0^2} & (s \leq R_0), \\ \frac{1}{s} & (s \geq R_0). \end{cases} \quad (273)$$

Eq. (237) is recovered in the limit  $R_0 \rightarrow 0$ .

#### 7.4.4 Electromagnetic waves

The laws of Coulomb/Biot-Savart are suggesting that the fields  $\mathbf{E}$  and  $\mathbf{B}$  are “generated” by source charges  $Q$  and currents  $I$ , respectively. To demonstrate that this is not the full truth, we now consider Maxwell’s equations in ”vacuum”,  $\rho(\mathbf{r}, t) = 0$  and  $\mathbf{J}(\mathbf{r}, t) = \mathbf{0}$ ,

$$(I) \quad \nabla \cdot \mathbf{E}(\mathbf{r}, t) = 0, \quad (II) \quad \nabla \cdot \mathbf{B}(\mathbf{r}, t) = 0, \quad (274)$$

$$(III) \quad \nabla \times \mathbf{B}(\mathbf{r}, t) = \epsilon_0 \mu_0 \frac{\partial \mathbf{E}(\mathbf{r}, t)}{\partial t}, \quad (IV) \quad \nabla \times \mathbf{E}(\mathbf{r}, t) = -\frac{\partial \mathbf{B}(\mathbf{r}, t)}{\partial t}. \quad (275)$$

Taking the curl on both sides of (III), we obtain on the LHS

$$\nabla \times [\nabla \times \mathbf{B}(\mathbf{r}, t)] \equiv \nabla \left[ \underbrace{\nabla \cdot \mathbf{B}(\mathbf{r}, t)}_0 \right] - \nabla^2 \mathbf{B}(\mathbf{r}, t), \quad (276)$$

where we have used an identity (exercises!) and (II). Consequently, (III) becomes

$$-\nabla^2 \mathbf{B}(\mathbf{r}, t) = \epsilon_0 \mu_0 \frac{\partial}{\partial t} [\nabla \times \mathbf{E}(\mathbf{r}, t)], \quad (277)$$

Using now (IV) for  $\nabla \times \mathbf{E}$ , we eventually find

$$\nabla^2 \mathbf{B}(\mathbf{r}, t) = \frac{1}{c^2} \frac{\partial^2 \mathbf{B}(\mathbf{r}, t)}{\partial t^2}, \quad c = \frac{1}{\sqrt{\epsilon_0 \mu_0}} \equiv 2.998 \cdot 10^8 \frac{\text{m}}{\text{s}}. \quad (278)$$

This is the **wave equation** for the components  $B_1, B_2, B_3$  of  $\mathbf{B}$ ,

$$\left( \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right) B_i(x, y, z, t) = \frac{1}{c^2} \frac{\partial^2}{\partial t^2} B_i(x, y, z, t) \quad (i = 1, 2, 3), \quad (279)$$

a **partial differential equation (PDE)** for the unknown function  $B_i(\mathbf{r}, t)$ .

Starting with (IV), a similar reasoning leads to the equation

$$\nabla^2 \mathbf{E}(\mathbf{r}, t) = \frac{1}{c^2} \frac{\partial^2 \mathbf{E}(\mathbf{r}, t)}{\partial t^2}. \quad (280)$$

We see: There are solutions  $\mathbf{E}(\mathbf{r}, t) \neq \mathbf{0}$  and  $\mathbf{B}(\mathbf{r}, t) \neq \mathbf{0}$ , even when no charges  $Q$  and no currents  $I$  are around.

## 7.5 Electrostatic potential and Poisson's equation

When the magnetic field is stationary,  $\frac{\partial}{\partial t}\mathbf{B}(\mathbf{r}, t) = \mathbf{0}$ , Faraday's law implies

$$\nabla \times \mathbf{E}(\mathbf{r}, t) = \mathbf{0}. \quad (281)$$

Consequently, due to the theorem of section 6.4, there is a scalar field  $\phi(\mathbf{r}, t)$  such that

$$\mathbf{E}(\mathbf{r}, t) = -\nabla\phi(\mathbf{r}, t). \quad (282)$$

In such a situation, the vector field  $\mathbf{E}(\mathbf{r}, t)$  is fixed by a (simpler) scalar field  $\phi(\mathbf{r}, t)$ , called the **electrostatic potential**.<sup>4</sup>

Combining Eq. (282) with Gauss' law, we see that  $\phi(\mathbf{r}, t)$  satisfies **Poisson's equation**,

$$\nabla^2\phi(\mathbf{r}, t) = -\frac{1}{\epsilon_0}\rho(\mathbf{r}, t). \quad (283)$$

For any given  $\rho(\mathbf{r}, t)$ , this is a **partial differential equation (PDE)** for the unknown function  $\phi(\mathbf{r}, t)$ . Once Eq. (283) is solved for  $\phi(\mathbf{r}, t)$ ,  $\mathbf{E}(\mathbf{r}, t)$  is obtained from Eq. (282).

**Remark:** For time-independent  $\rho = \rho(\mathbf{r})$ , a solution to Poisson's equation is given by

$$\phi(\mathbf{r}) = \frac{1}{4\pi\epsilon_0} \int d^3r' \frac{\rho(\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|}, \quad (284)$$

since (as we shall prove in the exercises)<sup>5</sup>

$$\begin{aligned} \nabla^2\phi(\mathbf{r}) &\equiv \frac{1}{4\pi\epsilon_0} \int d^3r' \rho(\mathbf{r}') \nabla^2 \frac{1}{|\mathbf{r} - \mathbf{r}'|} \\ &= \frac{1}{4\pi\epsilon_0} \int d^3r' \rho(\mathbf{r}') [-4\pi\delta(\mathbf{r} - \mathbf{r}')] \equiv -\frac{1}{\epsilon_0}\rho(\mathbf{r}). \end{aligned} \quad (285)$$

However, Eq. (284) is not the only solution. In particular, it is useless in cases when the charge distribution  $\rho(\mathbf{r}, t)$  is not known from the beginning. As an example, take a point charge  $Q$  at a distance  $a$  in front of a grounded metal wall: The potential  $\phi(\mathbf{r})$ , and thus also the field  $\mathbf{E}(\mathbf{r}) = -\nabla\phi(\mathbf{r})$ , in front of the wall is not only due to  $Q$  itself, but also due to an unknown charge distribution on the metal surface, **influenced** by  $Q$ .

<sup>4</sup>For a time-independent field  $\mathbf{E}(\mathbf{r})$ , the potential  $\phi(\mathbf{r})$  can be chosen to be time-independent as well. Then, the work done by  $\mathbf{E}(\mathbf{r})$  on a charge  $q$  with mass  $m$ , moving on some curve  $\Gamma \subset \Omega$ , is given by

$$W_\Gamma \equiv q \int_\Gamma d\mathbf{r} \cdot \mathbf{E}(\mathbf{r}) = -q[\phi(\mathbf{r}_B) - \phi(\mathbf{r}_A)].$$

This integral only depends on the starting point  $\mathbf{r}_A$  and the endpoint  $\mathbf{r}_B$  of  $\Gamma$ . With the corresponding speeds  $v_A$  and  $v_B$  of the charge, energy conservation then implies

$$\frac{m}{2}v_B^2 = \frac{m}{2}v_A^2 + W_\Gamma \quad \Leftrightarrow \quad \frac{m}{2}v_A^2 + q\phi(\mathbf{r}_A) = \frac{m}{2}v_B^2 + q\phi(\mathbf{r}_B).$$

Consequently, as a function of its position  $\mathbf{r}$ , the value  $q\phi(\mathbf{r})$  is the electrostatic **potential energy** of the charge  $q$  due to the electric field  $\mathbf{E}(\mathbf{r})$ .

<sup>5</sup>Dirac's delta function  $\delta(\mathbf{r})$  in 3D is the limit  $R \rightarrow 0$  of the finite function

$$\Delta_R(\mathbf{r}) = \begin{cases} \frac{1}{\frac{4\pi}{3}R^3} & (|\mathbf{r}| \leq R), \\ 0 & (|\mathbf{r}| > R). \end{cases}$$

## PART II

### Partial Differential Equations (PDEs)

## 8 Differential equations (DEs)

### 8.1 Partial versus ordinary

With a given charge density  $\rho(\mathbf{r})$ , e.g.:  $\rho(\mathbf{r}) = Q \frac{2a^{5/2}}{\pi^{3/2}} z^2 e^{-a(x^2+y^2+z^2)}$ , Poisson's equation

$$\nabla^2 \phi(\mathbf{r}) \equiv \left( \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right) \phi(x, y, z) = -\frac{1}{\epsilon_0} \rho(x, y, z) \quad (286)$$

for the unknown wanted function  $\phi(\mathbf{r})$  is a differential equation (DE), since derivatives of  $\phi$  enter. More precisely, it is a partial DE (**PDE**), since (partial) derivatives with respect to **more variables than one** are involved.

In contrast, an ordinary DE (**ODE**) is for a function  $f(x)$  with only one variable, e.g.,

$$f''(x) + f'(x) = 3x(2+x). \quad (287)$$

This ODE is of **second** order, since the highest derivative involved is  $f''(x)$ . Consequently (see below), its infinitely many solutions form a **two-parameter** set of functions,

$$f(x) = x^3 + Ae^{-x} + B \quad (A, B \in \mathbb{R}). \quad (288)$$

The set of solutions to a PDE, in contrast, is more involved. E.g., Laplace's equation,

$$\nabla^2 \phi(\mathbf{r}) \equiv \left( \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right) \phi(x, y, z) = 0, \quad (289)$$

which is a special case of Poisson's (with  $\rho \equiv 0$ ), has very different types of solutions. We list a few of them (with free parameters  $A, B, C, D, E, F, \alpha, \beta \in \mathbb{R}$ ):

$$\begin{aligned} \phi_1(\mathbf{r}) &= Ax^2 + By^2 - (A+B)z^2 + Dx + Ey + Fz + C, \\ \phi_2(\mathbf{r}) &= A \sin(\alpha x) \sin(\beta y) e^{-\gamma z}, \quad \gamma = \sqrt{\alpha^2 + \beta^2}, \\ \phi_3(\mathbf{r}) &= \frac{A}{\sqrt{x^2 + y^2 + z^2}} \equiv A(x^2 + y^2 + z^2)^{-1/2} \quad (\mathbf{r} \neq \mathbf{0}), \\ \phi_4(\mathbf{r}) &= (Ax + By + Cz)(x^2 + y^2 + z^2)^{-3/2} \quad (\mathbf{r} \neq \mathbf{0}). \end{aligned} \quad (290)$$

### 8.2 Some facts about ODEs

The ODE Eq. (287) is an **inhomogeneous linear ODE** with **constant coefficients**. In general, this important type of ODE, when of  $n$ -th order, has the form

$$f^{(n)}(x) + a_{n-1}f^{(n-1)}(x) + \dots + a_1f'(x) + a_0f(x) = b(x). \quad (291)$$

The  $a_k$  are given numbers and the **inhomogeneity**  $b(x)$  is a given function.

In Eq. (287), e.g.,  $a_2 = a_1 = 1$ ,  $a_0 = 0$ , and  $b(x) = 3x(2+x)$ .

In the **general solution** of Eq. (291),

$$f(x) = f_{\text{spec}}^{\text{inh}}(x) + f_{\text{gen}}^{\text{hom}}(x), \quad (292)$$

- $f_{\text{spec}}^{\text{inh}}(x)$  is **one particular** solution (found by an educated guess) and
  - $f_{\text{gen}}^{\text{hom}}(x)$  is the **general solution** of the related **homogeneous** ODE with  $b(x) = 0$ .
- Generally,  $f_{\text{gen}}^{\text{hom}}(x)$  is a linear combination of  $n$  linearly independent functions. The ansatz

$$f^{\text{hom}}(x) = e^{\lambda x} \quad (293)$$

transforms the homogeneous ODE into an **algebraic equation** for  $\lambda$ ,

$$\lambda^n + a_{n-1}\lambda^{n-1} + \dots + a_1\lambda + a_0 = 0. \quad (294)$$

In typical cases, its roots  $\lambda_1, \dots, \lambda_k$  yield the wanted  $n$  independent functions.

**Example 1:** A third-order inh. lin. ODE with constant coefficients is

$$f'''(x) - 4f''(x) + 13f'(x) + 50f(x) = 50x^2 + 26x - 8. \quad (295)$$

By educated guess, we find

$$f_{\text{spec}}^{\text{inh}}(x) = x^2. \quad (296)$$

The ansatz  $f^{\text{hom}}(x) = e^{\lambda x}$  for the homogeneous equation yields the algebraic equation

$$\lambda^3 - 4\lambda^2 + 13\lambda + 50 \equiv (\lambda + 2)(\lambda^2 - 6\lambda + 25) = 0, \quad (297)$$

which has three different roots,

$$\lambda_1 = -2, \quad \lambda_{2,3} = 3 \pm 4i. \quad (298)$$

The corresponding complex-valued solutions of the homogeneous ODE are

$$f_1(x) = e^{-2x}, \quad f_{2,3}(x) = e^{(3 \pm 4i)x} \equiv e^{3x} [\cos 4x \pm i \sin 4x]. \quad (299)$$

In each of them, both the real and the imaginary parts provide a real-valued solution. Since three linearly independent solutions of that kind are provided, we conclude

$$f_{\text{gen}}^{\text{hom}}(x) = A e^{-2x} + B e^{3x} \cos 4x + C e^{3x} \sin 4x \quad (A, B, C \in \mathbb{R}). \quad (300)$$

Consequently, any solution of the original inhomogeneous ODE has the form

$$f(x) = x^2 + A e^{-2x} + B e^{3x} \cos 4x + C e^{3x} \sin 4x. \quad (301)$$

$$= x^2 + A e^{-2x} + (B \cos 4x + C \sin 4x) e^{3x}. \quad (302)$$

**Example 2:** In section 10 we shall face the following ODEs,

$$\begin{aligned} X''(x) &= -\alpha^2 X(x), \\ Y''(y) &= -\beta^2 Y(y), \\ Z''(z) &= (\alpha^2 + \beta^2) Z(z). \end{aligned} \quad (303)$$

These are homogeneous 2nd-order ( $n = 2$ ) ODEs of the type of Eq. (291), with  $b(x) = 0$ . In the case  $\alpha, \beta > 0$ , the exponential ansatz yields the following general solutions

$$\begin{aligned} X(x) &= A_1 \cos(\alpha x) + A_2 \sin(\alpha x), \\ Y(y) &= B_1 \cos(\beta y) + B_2 \sin(\beta y), \\ Z(z) &= C_1 e^{\gamma z} + C_2 e^{-\gamma z}, \quad \gamma = \sqrt{\alpha^2 + \beta^2}. \end{aligned} \quad (304)$$

**Remark:** A completely different method ("separable variables"), applicable to many **linear or non-linear first-order ODEs**, will be treated in section 11.4.

## 9 Arising of PDEs in physics

**Poisson's equation** and the **wave equation** were our first examples of PDEs. We shall now derive several further important PDEs in physics.

### 9.1 Diffusion

#### 9.1.1 Densities of mass and current

We consider a gas (or liquid, etc.) of  $N$  identical particles (“molecules”) with masses  $m$ . For  $i \in \{1, \dots, N\}$ ,  $\mathbf{r}_i(t)$  is the position and  $\mathbf{v}_i(t)$  the velocity of particle  $i$  at time  $t$ .

Let  $\Omega_a(\mathbf{r})$  be a small sphere with radius  $a$  (and volume  $\frac{4\pi}{3}a^3$ ), centered at  $\mathbf{r}$ . The **mass density**  $\rho(\mathbf{r}, t)$  of the gas at position  $\mathbf{r}$  and time  $t$  is defined as

$$\rho(\mathbf{r}, t) = \frac{N_a(\mathbf{r}, t) m}{\frac{4\pi}{3}a^3}, \quad (305)$$

where  $N_a(\mathbf{r}, t)$  is the number of particles with positions  $\mathbf{r}_i(t)$  inside  $\Omega_a(\mathbf{r})$  at the time  $t$ ,

$$N_a(\mathbf{r}, t) = \sum_{\mathbf{r}_i(t) \in \Omega_a(\mathbf{r})} 1. \quad (306)$$

$\rho(\mathbf{r}, t) \equiv \rho(x, y, z, t)$  is a **smooth function** of  $x, y, z, t$ , when the radius  $a$ , smaller than the resolution  $L$  of position measurement, is considerably larger than the typical distance  $\ell$  between molecules,

$$L \gg a \gg \ell. \quad (307)$$

The **average velocity** of a molecule with position vector  $\mathbf{r}$  at time  $t$  is defined as

$$\mathbf{v}(\mathbf{r}, t) = \frac{1}{N_a(\mathbf{r}, t)} \sum_{\mathbf{r}_i(t) \in \Omega_a(\mathbf{r})} \mathbf{v}_i(t). \quad (308)$$

Defining the vector field of the **current density** of the gas as the product

$$\mathbf{J}(\mathbf{r}, t) = \rho(\mathbf{r}, t) \mathbf{v}(\mathbf{r}, t), \quad (309)$$

the amount  $I_\Sigma(t)$  of gas (in kg/s) passing at the time  $t$  per unit time through a piece  $\Sigma$  of surface is given by the flux integral

$$I_\Sigma(t) = \int_\Sigma d\mathbf{A} \cdot \mathbf{J}(\mathbf{r}, t) \equiv \int_\Sigma d^2r \mathbf{n}(\mathbf{r}) \cdot \mathbf{J}(\mathbf{r}, t). \quad (310)$$

### 9.1.2 Continuity equation

We now consider some arbitrary finite volume region  $\Omega$  (with surface  $\partial\Omega$ ). If the mass of  $\text{CO}_2$  created chemically per unit time within a small volume  $d^3r$  around the position  $\mathbf{r}$  at time  $t$  is given by  $s(\mathbf{r}, t) d^3r$ , the rate of change of its mass within  $\Omega$  is

$$\frac{d}{dt} \int_{\Omega} d^3r \rho(\mathbf{r}, t) = \int_{\Omega} d^3r s(\mathbf{r}, t) - \oint_{\partial\Omega} d\mathbf{A} \cdot \mathbf{J}(\mathbf{r}, t). \quad (311)$$

Here, the RHS represents the amount of  $\text{CO}_2$  created chemically within  $\Omega$  minus the amount leaving  $\Omega$  by flow through the surface  $\partial\Omega$ . [Note that  $\oint_{\partial\Omega} d\mathbf{A} \cdot \mathbf{J}(\mathbf{r}, t)$  is defined as the flux of  $\mathbf{J}$  *out* of  $\Omega$ .] By Gauß's law, we have

$$\oint_{\partial\Omega} d\mathbf{A} \cdot \mathbf{J}(\mathbf{r}, t) = \int_{\Omega} d^3r \nabla \cdot \mathbf{J}(\mathbf{r}, t). \quad (312)$$

Therefore, since Eq. (311) must hold for any region  $\Omega$ , we conclude

$$\frac{\partial}{\partial t} \rho(\mathbf{r}, t) + \nabla \cdot \mathbf{J}(\mathbf{r}, t) = s(\mathbf{r}, t). \quad (313)$$

When the gas is a **conserved quantity**,  $s(\mathbf{r}, t) = 0$ , we have the **continuity equation**

$$\frac{\partial}{\partial t} \rho(\mathbf{r}, t) + \nabla \cdot \mathbf{J}(\mathbf{r}, t) = 0. \quad (314)$$

### 9.1.3 Diffusion and convection

The velocity field is typically due to both **diffusion** and **convection**,

$$\mathbf{v}(\mathbf{r}, t) = \mathbf{v}_{\text{diff}}(\mathbf{r}, t) + \mathbf{v}_{\text{conv}}(\mathbf{r}, t). \quad (315)$$

For example, if the “gas” is  $\text{CO}_2$  in air,  $\mathbf{v}_{\text{conv}}(\mathbf{r}, t)$  is the **wind velocity** of the carrier medium (air). In contrast,  $\mathbf{v}_{\text{diff}}(\mathbf{r}, t)$  is the **average velocity** of the  $\text{CO}_2$  molecules in the absence of wind, due to **diffusion** between the air molecules ( $\text{N}_2$ ,  $\text{O}_2$ , Ar).

**Diffusion law:** The origin of diffusion is the density gradient,

$$\mathbf{J}_{\text{diff}}(\mathbf{r}, t) \equiv \rho(\mathbf{r}, t) \mathbf{v}_{\text{diff}}(\mathbf{r}, t) = -D \nabla \rho(\mathbf{r}, t). \quad (316)$$

The **diffusion constant**  $D$  has the physical dimension  $1 \frac{\text{m}^2}{\text{s}}$ .<sup>6</sup>

The total flux of the gas is

$$\mathbf{J}(\mathbf{r}, t) = -D \nabla \rho(\mathbf{r}, t) + \rho(\mathbf{r}, t) \mathbf{v}_{\text{conv}}(\mathbf{r}, t). \quad (317)$$

---

<sup>6</sup>Boltzmann's transport theory yields the expression

$$D = \frac{t_c k_B T}{m},$$

where the relaxation time  $t_c$  is close to the average collision time between molecules. Consequently, this “constant” can actually be a function  $D(\mathbf{r}, t)$ .

### 9.1.4 Diffusion equation

For simplicity, we assume  $\mathbf{v}_{\text{conv}}(\mathbf{r}, t) = 0$ ,

$$\mathbf{J}(\mathbf{r}, t) = -D \nabla \rho(\mathbf{r}, t). \quad (318)$$

Taking here the divergence  $\nabla \cdot \dots$  and using Eq. (313), we obtain the **diffusion equation**,

$$\frac{\partial}{\partial t} \rho(\mathbf{r}, t) = D \nabla^2 \rho(\mathbf{r}, t) + s(\mathbf{r}, t). \quad (319)$$

For a given value of the constant  $D$  and a given function  $s(\mathbf{r}, t)$ , this is a PDE for the unknown wanted function  $\rho(\mathbf{r}, t) \equiv \rho(x, y, z, t)$  of four variables  $x, y, z$ , and  $t$ ,

$$\frac{\partial}{\partial t} \rho(x, y, z, t) = D \left( \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right) \rho(x, y, z, t) + s(x, y, z, t). \quad (320)$$

### 9.1.5 Example for a solution

$\rho(\mathbf{r}, t)$  shall be the time-dependent density distribution of **ink in water**. In this case, the diffusion constant has the empirical value

$$D \approx 2.5 \cdot 10^{-4} \frac{\text{mm}^2}{\text{s}} \quad [\text{Eur. J. Phys. } \mathbf{25}, 331 (2004)]. \quad (321)$$

At the time  $t = 0$ , a small spherical droplet of ink is placed at the position  $\mathbf{r} = \mathbf{0}$  deep inside the water. No more ink is added to this system at any time  $t > 0$ ,

$$s(\mathbf{r}, t) = 0 \quad (\text{for } t > 0). \quad (322)$$

Due to diffusion, the ink, originally concentrated around the point  $\mathbf{r} = \mathbf{0}$ , will slowly spread out into the surrounding water. This effect is qualitatively described by the function

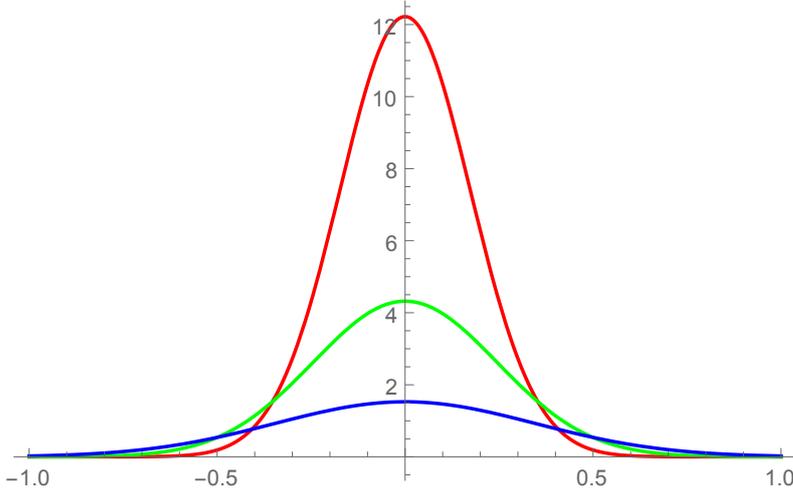
$$\begin{aligned} \rho(\mathbf{r}, t) = \rho_G(\mathbf{r}, t) &\equiv \frac{M}{(4\pi Dt)^{3/2}} e^{-\mathbf{r}^2/4Dt} \\ &\equiv \frac{M}{(4\pi Dt)^{3/2}} e^{-(x^2+y^2+z^2)/4Dt} = \rho_G(x, y, z, t) \quad (t > 0), \end{aligned} \quad (323)$$

where  $M$  is a constant with the dimension of a mass (kg). We shall see below that  $\rho_G(x, y, z, t)$  is in fact an exact solution of the diffusion equation (320). Note that the exponent  $\frac{\mathbf{r}^2}{4Dt}$  is dimensionless, while the prefactor has the correct dimension  $1 \frac{\text{kg}}{\text{m}^3}$  of a mass density. The function  $\rho_G(x, 0, 0, t)$  is plotted versus  $x$  in the following figure for the three different time values  $t = 60$  s (red),  $t = 120$  s (green), and  $t = 240$  s (blue).

Since no ink is added after  $t = 0$ , the total mass in the density distribution  $\rho_G(\mathbf{r}, t)$  should be a constant, not depending on  $t$ . In fact, elementary triple integration yields

$$\int d^3r \rho_G(\mathbf{r}, t) \equiv \frac{M}{(4\pi Dt)^{3/2}} \int_{-\infty}^{\infty} dx \int_{-\infty}^{\infty} dy \int_{-\infty}^{\infty} dz e^{-(x^2+y^2+z^2)/4Dt} = M, \quad (324)$$

where we have used  $\int_0^{\infty} du u^2 e^{-u^2} = \frac{1}{4} \sqrt{\pi}$ . After all,  $\rho_G(\mathbf{r}, t)$  is a reasonable candidate for the unknown true time-dependent density distribution of the ink.



To see that  $\rho_G(\mathbf{r}, t)$  is in fact a solution of Eq. (320), we apply the product rule to find

$$\begin{aligned} \frac{\partial}{\partial t} \rho_G(x, y, z, t) &= -\frac{3}{2} \frac{\rho_G(x, y, z, t)}{t} - \left[ -\frac{(x^2 + y^2 + z^2)}{4Dt^2} \rho_G(x, y, z, t) \right] \\ &= \left( -\frac{3}{2t} + \frac{(x^2 + y^2 + z^2)}{4Dt^2} \right) \rho_G(x, y, z, t). \end{aligned} \quad (325)$$

On the other hand, we have

$$\begin{aligned} D \frac{\partial^2}{\partial x^2} \rho_G(x, y, z, t) &= D \frac{\partial}{\partial x} \left( \frac{-2x}{4Dt} \rho_G(x, y, z, t) \right) \\ &= D \left( -\frac{2}{4Dt} + \frac{(-2x)^2}{(4Dt)^2} \right) \rho_G(x, y, z, t) \\ &= \left( -\frac{1}{2t} + \frac{x^2}{4Dt^2} \right) \rho_G(x, y, z, t). \end{aligned} \quad (326)$$

Consequently, adding the corresponding expressions for  $D \frac{\partial^2 \rho_G}{\partial y^2}$  and  $D \frac{\partial^2 \rho_G}{\partial z^2}$ , we obtain

$$D \left( \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right) \rho_G(x, y, z, t) = \left( -\frac{3}{2t} + \frac{(x^2 + y^2 + z^2)}{4Dt^2} \right) \rho_G(x, y, z, t), \quad (327)$$

which is the same as the result (325). Therefore,  $\rho_G(\mathbf{r}, t)$  is an exact solution of the PDE (320) in the special case  $s(\mathbf{r}, t) = 0$ .

**Remark 1:** The integral (324) is more appropriately evaluated in spherical coordinates,

$$\int d^3r \rho_G(\mathbf{r}, t) = \frac{M}{(4\pi Dt)^{3/2}} \int_0^\infty dr (4\pi r^2) e^{-r^2/4Dt} = \frac{M}{\pi^{3/2}} 4\pi \int_0^\infty du u^2 e^{-u^2} = M. \quad (328)$$

**Remark 2:** The result (327) can be obtained in vector notation: With the gradient

$$\nabla \rho_G(\mathbf{r}, t) = -\frac{\mathbf{r}}{2Dt} \rho_G(\mathbf{r}, t), \quad (329)$$

we have

$$\begin{aligned} D \nabla^2 \rho_G(\mathbf{r}, t) &\equiv D \nabla \cdot [\nabla \rho_G(\mathbf{r}, t)] = D \nabla \cdot \left[ -\frac{\mathbf{r}}{2Dt} \rho_G(\mathbf{r}, t) \right] \\ &= -\frac{1}{2t} \left[ (\nabla \cdot \mathbf{r}) \rho_G(\mathbf{r}, t) + \mathbf{r} \cdot \nabla \rho_G(\mathbf{r}, t) \right] \\ &= -\frac{1}{2t} \left[ 3 \rho_G(\mathbf{r}, t) - \frac{\mathbf{r}^2}{2Dt} \rho_G(\mathbf{r}, t) \right] \\ &= \left( -\frac{3}{2t} + \frac{\mathbf{r}^2}{4Dt^2} \right) \rho_G(\mathbf{r}, t). \end{aligned} \quad (330)$$

## 9.2 Heat flow

### 9.2.1 Specific heat

In a **uniform medium** (liquid water, solid silicon, etc.), we treat **heat (energy)** as a continuous quantity with generally non-uniform density  $\rho(\mathbf{r}, t)$  (in kJ/m<sup>3</sup>), satisfying

$$\rho(\mathbf{r}, t) = c\mu [T(\mathbf{r}, t) - T_0]. \quad (331)$$

Here,  $T(\mathbf{r}, t)$  is the **temperature distribution** in the medium,  $\mu = \frac{\Delta m}{\Delta V}$  is its uniform mass density (in kg/m<sup>3</sup>), and  $c$  is its **specific heat** (in kJ/kgK).  $T_0 \approx 300$  K is a constant reference temperature.

**Remark:** Multiplying with  $\Delta V$  yields the familiar relation  $Q = cm \Delta T$ .

If there are no heat sources or sinks in the medium, density  $\rho(\mathbf{r}, t)$  and current density  $\mathbf{J}(\mathbf{r}, t)$  of heat satisfy the continuity equation (314),

$$\frac{\partial}{\partial t} \rho(\mathbf{r}, t) + \nabla \cdot \mathbf{J}(\mathbf{r}, t) = 0. \quad (332)$$

On the other hand, Eq. (331) implies

$$\frac{\partial}{\partial t} \rho(\mathbf{r}, t) = c\mu \frac{\partial}{\partial t} T(\mathbf{r}, t). \quad (333)$$

### 9.2.2 Heat conduction

**Law of heat conduction:**  $\mathbf{J}(\mathbf{r}, t)$  is proportional to the gradient of  $T(\mathbf{r}, t)$ ,

$$\mathbf{J}(\mathbf{r}, t) = -\lambda \nabla T(\mathbf{r}, t). \quad (334)$$

This law is quite reminiscent of the diffusion law (316).

Combining Eqs. (332) and (333), we may use Eq. (334) to eliminate  $\mathbf{J}$ ,

$$c\mu \frac{\partial}{\partial t} T(\mathbf{r}, t) = -\nabla \cdot \mathbf{J}(\mathbf{r}, t) = \lambda \nabla \cdot [\nabla T(\mathbf{r}, t)] \equiv \lambda \nabla^2 T(\mathbf{r}, t). \quad (335)$$

This is a PDE for the temperature distribution  $T(\mathbf{r}, t)$  in the medium,

$$\frac{\partial}{\partial t} T(\mathbf{r}, t) = \frac{\lambda}{c\mu} \nabla^2 T(\mathbf{r}, t). \quad (336)$$

### 9.2.3 Laplace's equation

A **stationary** temperature distribution, with  $\frac{\partial}{\partial t} T(\mathbf{r}, t) = 0$ , obeys **Laplace's equation**,

$$\nabla^2 T(\mathbf{r}) = 0. \quad (337)$$

## 9.3 Waves

### 9.3.1 Young's modulus of a spring

By **Hooke's law**, the restoring force of a spring with spring constant  $k > 0$  is

$$F(u) = -ku, \quad (338)$$

where  $u$  is the displacement ( $u > 0$ : stretched spring,  $u < 0$ : compressed spring). In any case,  $F$  has the opposite sign (direction) of  $u$ .

As the spring is cut into two equal pieces, the resulting two smaller springs have spring constant  $k' = 2k$  each. Generally, a spring with equilibrium length  $L$  has the constant

$$k = \frac{Y}{L}, \quad (339)$$

where **Young's modulus**  $Y$  (in N) is a material constant, independent of  $L$ .

### 9.3.2 Continuum limit of a linear chain

A linear chain is an array of (infinitely) many identical point masses  $m$  on the  $x$ -axis which at equilibrium occupy the equidistant sites

$$x_\alpha = \alpha \cdot \Delta x \quad (\alpha \in \mathbb{Z}). \quad (340)$$

Joined by identical springs (each with spring constant  $k$  and equilibrium length  $L = \Delta x$ ), the point masses may perform longitudinal oscillations in  $x$ -direction.

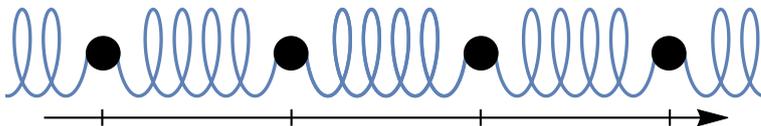


Figure 8: A section of a linear chain, with four masses (black dots) at their equilibrium positions  $x_{\alpha-1}$ ,  $x_\alpha$ ,  $x_{\alpha+1}$ , and  $x_{\alpha+2}$ , respectively, on the  $x$ -axis.

Let  $u_\alpha(t)$  be the displacement of mass  $\alpha$  from its equilibrium position  $x_\alpha$  at time  $t$ . Then, its equation of motion reads

$$m \ddot{u}_\alpha(t) = k[u_{\alpha+1}(t) - u_\alpha(t)] - k[u_\alpha(t) - u_{\alpha-1}(t)]. \quad (341)$$

Writing  $u_\alpha(t) = u(x_\alpha, t)$ , with a function  $u(x, t)$  of  $x$  and  $t$ , this equation reads

$$m \frac{\partial^2 u(x_\alpha, t)}{\partial t^2} = k [u(x_\alpha + \Delta x, t) - u(x_\alpha, t)] - k [u(x_\alpha, t) - u(x_\alpha - \Delta x, t)]. \quad (342)$$

We divide by  $\Delta x$ , and write

$$\frac{m}{\Delta x} \frac{\partial^2 u(x_\alpha, t)}{\partial t^2} = k \Delta x \cdot \frac{\frac{u(x_\alpha + \Delta x, t) - u(x_\alpha, t)}{\Delta x} - \frac{u(x_\alpha, t) - u(x_\alpha - \Delta x, t)}{\Delta x}}{\Delta x}. \quad (343)$$

Here,  $\frac{m}{\Delta x} = \mu$  is the mass per unit length (in  $\frac{\text{kg}}{\text{m}}$ ) of the chain and  $k\Delta x = Y$  is Young's modulus of the springs.

In the limit  $\Delta x \rightarrow 0$ , when the chain becomes a **continuous rod**, we obtain

$$\begin{aligned} \mu \frac{\partial^2 u(x, t)}{\partial t^2} &= Y \lim_{\Delta x \rightarrow 0} \frac{\frac{\partial u(x, t)}{\partial x} - \frac{\partial u(x - \Delta x, t)}{\partial x}}{\Delta x} \\ &\equiv Y \frac{\partial^2 u(x, t)}{\partial x^2}. \end{aligned} \quad (344)$$

This is a PDE for the function  $u(x, t)$ . Describing **longitudinal vibrations** of the rod,  $u(x, t)$  is the displacement  $u$  at time  $t$  of a point with equilibrium position  $x$  on the rod.

### 9.3.3 The wave equation

The latter PDE is the one-dimensional **wave equation**,

$$\frac{1}{c^2} \frac{\partial^2 u(x, t)}{\partial t^2} = \frac{\partial^2 u(x, t)}{\partial x^2}. \quad (345)$$

The constant  $c$ , here  $c = \sqrt{Y/\mu}$ , in general turns out to be the **velocity** in  $x$ -direction of **waves** that are described by the solution  $u(x, t)$  (see section 10.3.1).

The same equation holds for the **transversal vibration** in  $y$ -direction of a **violin string** on the  $x$ -axis, when  $Y$  is replaced with the **string tension**  $F$ . When a 1D **sound wave** is propagating in a thin rigid straight pipe (along the  $x$ -axis), the pressure  $P(x, t)$  of the medium in the pipe satisfies the wave equation

$$\frac{\partial^2 P(x, t)}{\partial x^2} - \frac{1}{c^2} \frac{\partial^2 P(x, t)}{\partial t^2} = 0, \quad c = \frac{1}{\sqrt{\kappa\rho}}, \quad (346)$$

where  $\rho$  is the mass density and  $\kappa$  the **adiabatic compressibility** of the medium,

$$\kappa = -\frac{1}{V} \left( \frac{\partial V}{\partial P} \right)_S. \quad (347)$$

For perfect gases, with  $dQ = \frac{f+2}{2} P dV + \frac{f}{2} V dP$ , we find:  $\kappa = \frac{f}{f+2} \frac{1}{P}$ .

System	$u(x, t)$	$c$
Vibrating <b>rod</b>	<b>Longitudinal displacement</b> at time $t$ in $x$ -direction of a point on the rod which at rest would be at position $x$ .	$\sqrt{\frac{Y}{\mu}}$
Vibrating <b>string</b>	<b>Transversal displacement</b> at time $t$ in $y$ -direction of a point on the string with $x$ -coordinate $x$ .	$\sqrt{\frac{F}{\mu}}$
<b>Sound</b> in a thin straight pipe	<b>Air pressure</b> $u(x, t) = P(x, t)$ at time $t$ at position $x$ in the pipe	$\frac{1}{\sqrt{\kappa\rho}}$

Eq. (345) is a special case of the wave equation in three dimensions,

$$\frac{1}{c^2} \frac{\partial^2 u(x, y, z, t)}{\partial t^2} = \frac{\partial^2 u(x, y, z, t)}{\partial x^2} + \frac{\partial^2 u(x, y, z, t)}{\partial y^2} + \frac{\partial^2 u(x, y, z, t)}{\partial z^2} \quad (348)$$

or, in short-hand notation,

$$\frac{1}{c^2} \frac{\partial^2 u(\mathbf{r}, t)}{\partial t^2} = \nabla^2 u(\mathbf{r}, t). \quad (349)$$

## 9.4 Classification of PDEs

A second-order linear PDE for a function  $u(x, y)$  with two variables has the general form

$$A u_{xx} + B u_{xy} + C u_{yy} + D u_x + E u_y + F u + G = 0, \quad (350)$$

in the short-hand notation  $u_{xx} = \frac{\partial^2}{\partial x^2} u(x, y)$ ,  $u_{xy} = \frac{\partial^2}{\partial x \partial y} u(x, y)$ ,  $u_x = \frac{\partial}{\partial x} u(x, y)$ , etc.

The coefficients  $A, B, \dots, G$  are given functions of  $(x, y)$  or, typically, constant numbers.

### Examples:

Name of the PDE	Form of the PDE	non-zero coeffs	PDE type
1D wave equation	$\frac{1}{c^2} \frac{\partial^2}{\partial t^2} u(x, t) = \frac{\partial^2}{\partial x^2} u(x, t)$	$A = -1, C = \frac{1}{c^2}$	hyperbolic
1D heat conduction	$\frac{\partial}{\partial t} u(x, t) = \frac{\lambda}{c\mu} \frac{\partial^2}{\partial x^2} u(x, t)$	$A = -\frac{\lambda}{c\mu}, E = 1$	parabolic
2D Laplace equation	$\frac{\partial^2}{\partial x^2} u(x, y) + \frac{\partial^2}{\partial y^2} u(x, y) = 0$	$A = 1, C = 1$	elliptic

**Def.:** The PDE Eq. (350) is called<sup>7</sup>

$$\left. \begin{array}{l} \text{hyperbolic} \\ \text{parabolic} \\ \text{elliptic} \end{array} \right\} \text{ when } \left\{ \begin{array}{l} B^2 - 4AC > 0, \\ B^2 - 4AC = 0, \\ B^2 - 4AC < 0. \end{array} \right.$$

**Theorem:** The PDE Eq. (350) has  $n$  **real characteristics** (RCs), where  $n = 2$  in the hyperbolic case,  $n = 1$  in the parabolic case, and  $n = 0$  in the elliptic case.

<sup>7</sup>Geometrical reason: The solutions  $(x, y) \in \mathbb{R}^2$  to the two-variable quadratic equation

$$ax^2 + bxy + cy^2 = 1$$

form a hyperbola ( $b^2 - 4ac > 0$ ), a parabola ( $b^2 - 4ac = 0$ ) or an ellipse ( $b^2 - 4ac < 0$ ). To see this, we rotate in the case  $b \neq 0$  the axes in the  $xy$ -plane by a proper angle  $\phi$ ,

$$\begin{aligned} x &= \xi \cos \phi - \eta \sin \phi, \\ y &= \xi \sin \phi + \eta \cos \phi, \end{aligned}$$

to obtain the simpler equation  $\frac{\xi^2}{\alpha} + \frac{\eta^2}{\beta} = 1$ . The required angle  $\phi$  only needs to satisfy

$$\cos^2 \phi = \frac{1}{2} \left[ 1 \pm \sqrt{\frac{(c-a)^2}{b^2 + (c-a)^2}} \right].$$

## 10 Analytical methods for solving PDEs

### 10.1 Diffusion equation: Green's function

**Theorem:** Let  $\rho(\mathbf{r}, t = 0) = \rho_0(\mathbf{r})$  be the **initial density distribution** (at the time  $t = 0$ ) of a diffusing substance (e.g., ink in water). Then, provided that there are no sources or sinks, the future distribution (at times  $t > 0$ ) is given by the volume integral

$$\rho(\mathbf{r}, t) = \int d^3r' \rho_0(\mathbf{r}') G(\mathbf{r}' - \mathbf{r}, t) \quad (\text{a CONVOLUTION integral !}), \quad (351)$$

with the so-called **Green's function** of the diffusion equation (cf. section 9.1.5)

$$G(\mathbf{r}' - \mathbf{r}, t) = \frac{e^{-(\mathbf{r}' - \mathbf{r})^2/4Dt}}{(4\pi Dt)^{3/2}} \equiv \frac{1}{M} \rho_G(\mathbf{r}' - \mathbf{r}, t). \quad (352)$$

**Proof:** • Since  $G(\mathbf{r}, t) = \frac{\rho_G(\mathbf{r}, t)}{M}$  satisfies the diffusion equation, see section 9.1.5, we have

$$\begin{aligned} \frac{\partial}{\partial t} G(\mathbf{r}' - \mathbf{r}, t) &\equiv \left[ \frac{\partial}{\partial t} G(\mathbf{s}, t) \right] \Big|_{\mathbf{s}=\mathbf{r}'-\mathbf{r}} \\ &= \left[ D \nabla_{\mathbf{s}}^2 G(\mathbf{s}, t) \right] \Big|_{\mathbf{s}=\mathbf{r}'-\mathbf{r}} = D \nabla_{\mathbf{r}}^2 G(\mathbf{r}' - \mathbf{r}, t). \end{aligned} \quad (353)$$

Consequently, we have in Eq. (351)

$$\begin{aligned} \frac{\partial}{\partial t} \rho(\mathbf{r}, t) &\equiv \int d^3r' \rho_0(\mathbf{r}') \frac{\partial}{\partial t} G(\mathbf{r}' - \mathbf{r}, t) \\ &= \int d^3r' \rho_0(\mathbf{r}') D \nabla_{\mathbf{r}}^2 G(\mathbf{r}' - \mathbf{r}, t) \\ &= D \nabla_{\mathbf{r}}^2 \int d^3r' \rho_0(\mathbf{r}') G(\mathbf{r}' - \mathbf{r}, t) \equiv D \nabla^2 \rho(\mathbf{r}, t), \end{aligned} \quad (354)$$

and  $\rho(\mathbf{r}, t)$  is in fact a solution of the diffusion equation.

• We still have to show that  $\rho(\mathbf{r}, t) \rightarrow \rho_0(\mathbf{r})$  as  $t \rightarrow 0$ :

Since  $\int d^3r \rho_G(\mathbf{r}, t) = M$ , see Eq. (324) in section 9.1.5, we have at any time  $t > 0$

$$\int d^3r' G(\mathbf{r}' - \mathbf{r}, t) = 1. \quad (355)$$

On the other hand, for any  $\mathbf{r}' \neq \mathbf{r}$ , we have  $G(\mathbf{r}' - \mathbf{r}, t) \rightarrow 0$  as  $t \rightarrow 0$ , see Eq. (352). Consequently, as a function of  $\mathbf{r}'$ ,  $G(\mathbf{r}' - \mathbf{r}, t)$  becomes infinitely **peaked** around  $\mathbf{r}' = \mathbf{r}$ ,

$$\lim_{t \rightarrow 0^+} \rho(\mathbf{r}, t) \equiv \int d^3r' \rho_0(\mathbf{r}') \underbrace{\left[ \lim_{t \rightarrow 0^+} G(\mathbf{r}' - \mathbf{r}, t) \right]}_{\delta(\mathbf{r}' - \mathbf{r})} = \rho_0(\mathbf{r}). \quad (356)$$

**Example:** A large water container is separated by a wall ( $yz$ -plane) into two halves. The left half ( $x < 0$ ) contains ink with uniform density  $\rho_1$ , the right half ( $x > 0$ ) is ink-free,

$$\rho_0(\mathbf{r}) \equiv \rho_0(x, y, z) = \begin{cases} \rho_1 & (x < 0), \\ 0 & (x > 0). \end{cases} \quad (357)$$

As the wall is removed at  $t = 0$ , the ink distribution  $\rho(\mathbf{r}, t)$  for any time  $t > 0$  is

$$\rho(\mathbf{r}, t) \equiv \int d^3r' \rho_0(\mathbf{r}') G(\mathbf{r}' - \mathbf{r}, t) = \rho_1 \int_{-\infty}^0 dx' \int_{-\infty}^{\infty} dy' \int_{-\infty}^{\infty} dz' G(\mathbf{r}' - \mathbf{r}, t). \quad (358)$$

With Eq. (352), we explicitly have

$$\rho(\mathbf{r}, t) = \frac{\rho_1}{(4\pi Dt)^{3/2}} \int_{-\infty}^0 dx' e^{-(x'-x)^2/4Dt} \int_{-\infty}^{\infty} dy' e^{-(y'-y)^2/4Dt} \int_{-\infty}^{\infty} dz' e^{-(z'-z)^2/4Dt}. \quad (359)$$

Substituting  $\frac{x'-x}{\sqrt{4Dt}} = u$ , with  $dx' = (4Dt)^{1/2} du$ , as well as  $\frac{y'-y}{\sqrt{4Dt}} = v$ ,  $\frac{z'-z}{\sqrt{4Dt}} = w$ , we get

$$\begin{aligned} \rho(\mathbf{r}, t) &= \frac{\rho_1}{(4\pi Dt)^{3/2}} (4Dt)^{3/2} \int_{-\infty}^{-x/\sqrt{4Dt}} du e^{-u^2} \underbrace{\int_{-\infty}^{\infty} dv e^{-v^2}}_{\sqrt{\pi}} \underbrace{\int_{-\infty}^{\infty} dw e^{-w^2}}_{\sqrt{\pi}} \\ &= \frac{\rho_1}{\sqrt{\pi}} \int_{-\infty}^{-x/\sqrt{4Dt}} du e^{-u^2} \equiv \rho(x, t). \end{aligned} \quad (360)$$

In terms of the **error function**  $\operatorname{erf}(s) = \frac{2}{\sqrt{\pi}} \int_0^s du e^{-u^2}$ , we obtain

$$\begin{aligned} \rho(x, t) &= \frac{\rho_1}{\sqrt{\pi}} \left\{ \int_{-\infty}^0 + \int_0^{-x/\sqrt{4Dt}} \right\} du e^{-u^2} \\ &= \frac{\rho_1}{\sqrt{\pi}} \left\{ \frac{\sqrt{\pi}}{2} + \frac{\sqrt{\pi}}{2} \operatorname{erf}\left(-\frac{x}{\sqrt{4Dt}}\right) \right\} \\ &= \frac{\rho_1}{2} \left\{ 1 + \operatorname{erf}\left(-\frac{x}{\sqrt{4Dt}}\right) \right\}. \end{aligned} \quad (361)$$

- The function  $f(x) = \frac{1}{2}\{1 + \operatorname{erf}(-\frac{x}{a})\}$  is monotonically decreasing from  $f(-\infty) = 1$  to  $f(+\infty) = 0$ , with  $f(0) = \frac{1}{2}$  and  $f'(0) = -\frac{1}{a\sqrt{\pi}}$ , see Fig. 9.

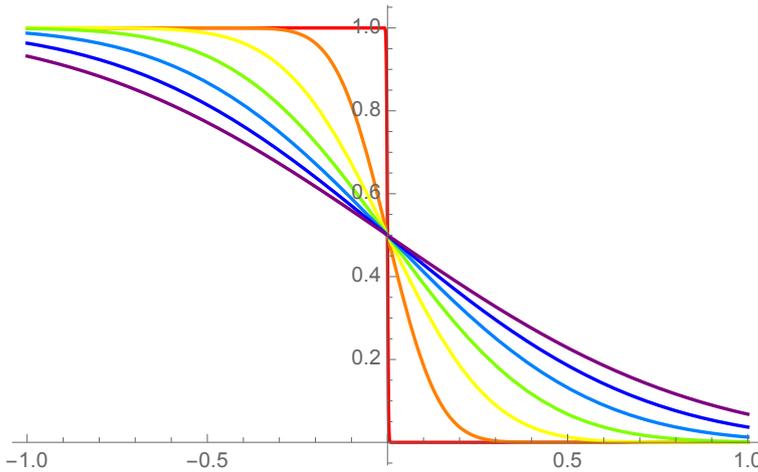


Figure 9: The function  $\rho(x, t)$  of Eq. (361) in units of  $\rho_1$ , plotted versus  $x$  (in mm), for different times  $t$  (in s):  $t = 0.01$  (red), 25 (orange) 100, 225, 400, 625 and 900 (violet), using the diffusion constant  $D = 2.5 \cdot 10^{-4} \frac{\text{mm}^2}{\text{s}}$  for **ink in water**, Eq. (321).

- Notice that  $\rho(x, t)$  of Eq. (361) is a solution of the **1D diffusion equation**  $\frac{\partial \rho}{\partial t} = D \frac{\partial^2 \rho}{\partial x^2}$ .

## 10.2 Heat flow: Separation ansatz

A rectangular metal plate in the  $xy$ -plane, with corners  $(x, y) = (0, 0), (a, 0), (0, b), (a, b)$ , has its top edge  $y = b$  held at some temperature  $T = T_0$  (red in the figure), while the other three edges (blue in the figure) are held at  $T = 0$  (in °C).

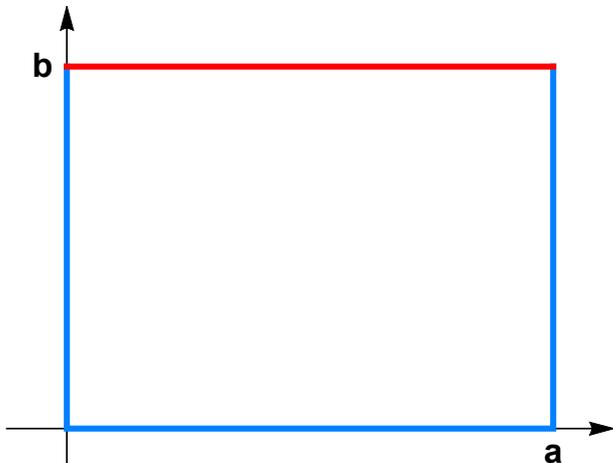


Figure 10: The boundary conditions of Eqs. (363) and (364).

When the plate, away from its edges, has no heat sources or sinks, its temperature distribution  $T(x, y)$  at steady-state equilibrium is that solution to **Laplace's equation**,

$$\nabla^2 T(x, y) \equiv \left( \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} \right) T(x, y) = 0, \quad (362)$$

which satisfies the corresponding **boundary conditions** (BCs)

$$T(0, y) = T(a, y) = T(x, 0) = 0 \quad (0 \leq x \leq a, 0 \leq y \leq b), \quad (363)$$

$$T(x, b) = t(x) \quad (0 \leq x \leq a), \quad (364)$$

where  $t(x)$  may be any **given function**. In our present example,  $t(x) = T_0$  (a constant).

### 10.2.1 Factorizing solutions

As a starting point, we look for solutions of the factorized form (**separation ansatz**)

$$T_{\text{sep}}(x, y) = X(x)Y(y). \quad (365)$$

Typically, such a solution can satisfy only the homogeneous BCs, Eq. (363), but not the inhomogeneous one, Eq. (364), with  $t(x) \neq 0$ , see section 10.2.2.

With the ansatz (365), our PDE Eq. (362) reads

$$\nabla^2 T_{\text{sep}}(x, y) \equiv X''(x)Y(y) + X(x)Y''(y) = 0.$$

Dividing both sides by  $X(x)Y(y)$ , we obtain

$$\frac{X''(x)}{X(x)} + \frac{Y''(y)}{Y(y)} = 0. \quad (366)$$

Since the variables  $x \in [0, a]$  and  $y \in [0, b]$  may vary independently of each other, Eq. (366) can be true only when each summand on the LHS is a **constant**,  $\mp k^2$  say (with  $k \in \mathbb{R}$ ),

$$X''(x) = -k^2 X(x), \quad Y''(y) = k^2 Y(y). \quad (367)$$

(Choosing opposite signs, we could not satisfy the first two of the four conditions below!) The general solutions of these ODEs (for any value of the yet unknown  $k$ ) are

$$X(x) = A \cos(kx) + B \sin(kx), \quad Y(y) = C e^{ky} + D e^{-ky} \quad (A, B, C, D \in \mathbb{R}).$$

- The condition  $T_{\text{sep}}(0, y) = 0$  requires that  $X(0) = 0$  or, equivalently,  $A = 0$ .
- Then, the condition  $T_{\text{sep}}(a, y) = 0$  requires that  $X(a) = 0$  or  $ka = n\pi$ , with  $n \in \mathbb{Z}$ ,

$$k = n \frac{\pi}{a} \equiv k_n \quad (n \in \mathbb{Z}).$$

- The condition  $T_{\text{sep}}(x, 0) = 0$  requires that  $Y(0) = 0$  or  $D = -C$ , and we have

$$\begin{aligned} T_{\text{sep}}(x, y) &= B \sin(k_n x) C [e^{k_n y} - e^{-k_n y}] \\ &\equiv B' \sin(k_n x) \sinh(k_n y) \quad (n \in \mathbb{N}), \end{aligned} \quad (368)$$

where  $\sinh(u) = \frac{1}{2}(e^u - e^{-u})$  and thus  $B' = 2BC$ ; the integer  $n \in \mathbb{N}$  is not fixed. For  $n = 5$  and  $B' = 1$ , this function is plotted in Fig. 11.

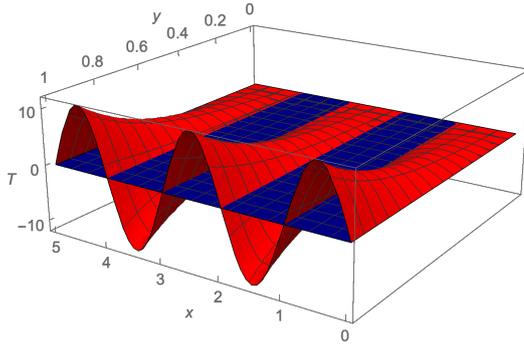


Figure 11: The function  $T_{\text{sep}}(x, y) = \sin(k_5 x) \sinh(k_5 y)$  for  $a = 5$  and  $0 \leq y \leq 1$ .

### 10.2.2 Superposition

$T_{\text{sep}}(x, y)$  cannot satisfy the inh BC  $T(x, b) = t(x)$ , with a given function  $t(x)$ , since

$$T_{\text{sep}}(x, b) \equiv B' \sin(k_n x) \sinh(k_n b) \quad (369)$$

is either zero (for  $n = 0$ ) or (for  $n \neq 0$ ) is a **sine function** of  $x \in [0, a]$ , see Fig. 11. However, any **superposition** of these factorising functions,

$$T(x, y) = \sum_{n=1}^{\infty} B'_n \sin(k_n x) \sinh(k_n y), \quad (370)$$

with arbitrary values of  $B'_1, B'_2, B'_3, \dots$ , still satisfies the homogeneous conditions (363). Choosing, e.g.,  $(B'_1, B'_2, B'_3) = (15, 3, 2)$  and  $B'_n = 0$  for  $n \geq 4$ , we obtain the plot:

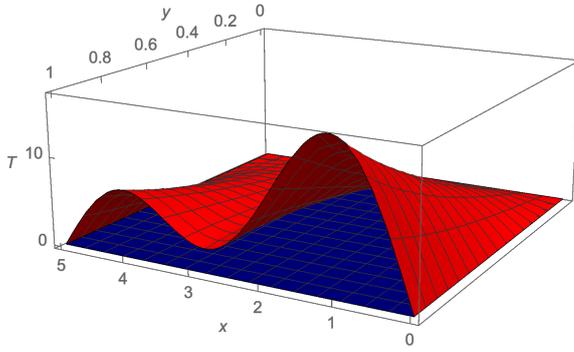


Figure 12:  $T(x, y) = \sum_{n=1}^3 B'_n \sin(k_n x) \sinh(k_n y)$  for  $(B'_1, B'_2, B'_3) = (15, 3, 2)$ .

Fig. 12 is indicating that, by a proper choice of the  $B'_n$ , we may succeed to satisfy also the inhomogeneous BC  $T(x, b) = t(x)$ , for any given function  $t(x)$ ,

$$T(x, b) \equiv \sum_{n=1}^{\infty} \underbrace{B'_n \sinh(k_n b)}_{\beta_n} \sin(k_n x) = t(x). \quad (371)$$

Writing  $x = \frac{a}{\pi} z$  (so that  $k_n x = nz$ ) and  $t(\frac{a}{\pi} z) = f(z)$ , this reads

$$\underbrace{\sum_{n=1}^{\infty} \beta_n \sin(nz)}_{f_S(z)} = f(z) \quad (0 < z < \pi). \quad (372)$$

The LHS is the **Fourier series**  $f_S(z)$  of a function  $f(z)$  for  $z \in [0, \pi]$ .

In the case  $f(z) = T_0 = \text{const}$ , the coefficients  $\beta_n$  are given by Eq. (386) of section 10.2.4,

$$(\beta_1, \beta_2, \beta_3, \beta_4, \beta_5, \beta_6, \beta_7, \dots) = \frac{4T_0}{\pi} \times (1, 0, \frac{1}{3}, 0, \frac{1}{5}, 0, \frac{1}{7}, \dots), \quad (373)$$

and the exact solution of our original problem, Eqs. (362) with the BCs (363) and (364), is given by Eq. (370) with  $B'_n = \frac{\beta_n}{\sinh(k_n b)}$ ,

$$T(x, y) = \frac{4T_0}{\pi} \sum_{n=1,3,5,\dots} \frac{1}{n \sinh(k_n b)} \sin(k_n x) \sinh(k_n y) \quad \left(k_n = n \frac{\pi}{a}\right). \quad (374)$$

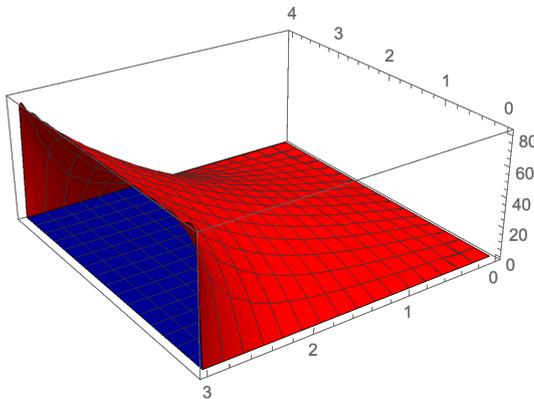


Figure 13:  $T(x, y)$  of Eq. (374), truncated at  $n = 199$ , using  $(T_0, a, b) = (80, 4, 3)$ .

### 10.2.3 More general boundary conditions

Now, each one of the four edges of the rectangle shall have a different temperature,

$$T(a, y) = T_1, \quad T(x, b) = T_2, \quad T(0, y) = T_3, \quad T(x, 0) = T_4. \quad (375)$$

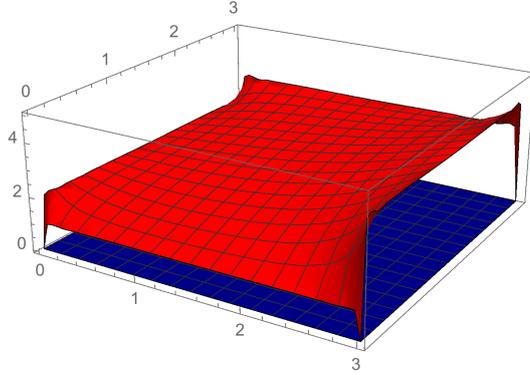


Figure 14:  $T(x, y)$  for  $a = b = 3$  and  $(T_1, T_2, T_3, T_4) = (1, 2, 3, 4)$ .  
In this case,  $T(1, 1) = 3.02287$ ,  $T(1, 2) = T(2, 1) = \frac{5}{2}$  and  $T(2, 2) = 1.97713$ .

Then, in terms of the functions

$$t_U(x, y) = \frac{4}{\pi} \sum_{n=1,3,5,\dots} \frac{1}{n \sinh(n\pi \frac{b}{a})} \sin\left(n\pi \frac{x}{a}\right) \sinh\left(n\pi \frac{y}{a}\right), \quad (376)$$

$$t_R(x, y) = \frac{4}{\pi} \sum_{n=1,3,5,\dots} \frac{1}{n \sinh(n\pi \frac{a}{b})} \sinh\left(n\pi \frac{x}{b}\right) \sin\left(n\pi \frac{y}{b}\right), \quad (377)$$

the solution is obviously given by the linear combination

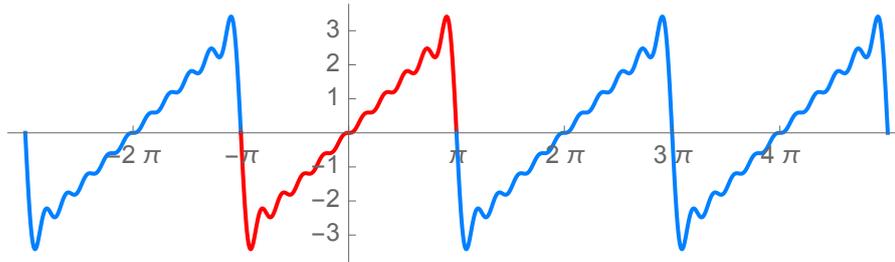
$$T(x, y) = T_1 \cdot t_R(x, y) + T_2 \cdot t_U(x, y) + T_3 \cdot t_R(a - x, y) + T_4 \cdot t_U(x, b - y). \quad (378)$$

### 10.2.4 Fourier analysis

Consider (for  $z \in \mathbb{R}$ ) the function

$$\begin{aligned} f_S^{(N)}(z) &= 2 \left[ \sin z - \frac{1}{2} \sin 2z + \frac{1}{3} \sin 3z - + \dots + (-1)^{N+1} \frac{1}{N} \sin(Nz) \right] \\ &\equiv 2 \sum_{n=1}^N \frac{(-1)^{n+1}}{n} \sin(nz), \end{aligned} \quad (379)$$

plotted here versus  $z \in \mathbb{R}$  for the case  $N = 10$ :



Clearly, this function

- has the **periodicity** interval  $[-\pi, \pi]$ .
- is an **odd function**,  $f_S^{(N)}(-z) = -f_S^{(N)}(z)$ .
- for  $z \in [-\pi, \pi]$  seems to approach the function  $f(z) = z$ , as  $N \rightarrow \infty$ .

**Theorem:** For  $z \in [-\pi, \pi]$ , let  $f(z)$  be **any given continuous** function.

Then, there are numbers  $\alpha_0, \alpha_1, \beta_1, \alpha_2, \beta_2, \dots$  (called the Fourier coefficients of  $f$ ) with the property that for each  $z \in ]-\pi, \pi[$ , the value  $f(z)$  is given by the limit

$$f(z) = \frac{\alpha_0}{2} + \lim_{N \rightarrow \infty} \sum_{n=1}^N [\alpha_n \cos(nz) + \beta_n \sin(nz)] \quad (-\pi < z < \pi). \quad (380)$$

The right-hand side of Eq. (380) is called a **Fourier series**, in short-hand notation

$$f_S(z) = \frac{\alpha_0}{2} + \sum_{n=1}^{\infty} [\alpha_n \cos(nz) + \beta_n \sin(nz)]. \quad (381)$$

**Example 1:** In the case  $f(z) = z$ , the coefficients are

$$\begin{aligned} \alpha_n &= 0 & (n = 0, 1, 2, 3, \dots), \\ \beta_n &= (-1)^{n+1} \frac{2}{n} & (n = 1, 2, 3, \dots). \end{aligned} \quad (382)$$

**Remark 1:** To prove this theorem, the set  $\{1, \cos z, \sin z, \cos 2z, \sin 2z, \dots\}$  must be shown to be a "**complete set**" of functions on the interval  $-\pi < z < \pi$ .

**Corollary:** Given  $f(z)$ , the Fourier coefficients are easily obtained as integrals,

$$\begin{aligned} \alpha_n &= \frac{1}{\pi} \int_{-\pi}^{\pi} dz f(z) \cos(nz) & (n = 0, 1, 2, 3, \dots), \\ \beta_n &= \frac{1}{\pi} \int_{-\pi}^{\pi} dz f(z) \sin(nz) & (n = 1, 2, 3, \dots). \end{aligned} \quad (383)$$

**Proof:** Due to the theorem, we have  $f(z) = f_S(z)$  for  $-\pi < z < \pi$ . Consequently, e.g.,

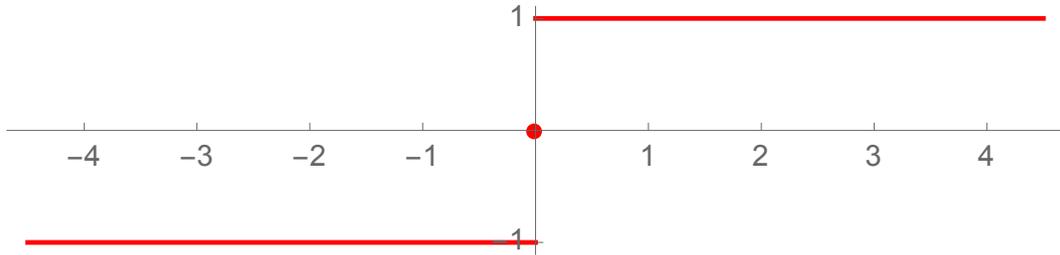
$$\begin{aligned}
 \int_{-\pi}^{\pi} dz f(z) \cos(mz) &= \int_{-\pi}^{\pi} dz f_S(z) \cos(mz) \\
 &= \int_{-\pi}^{\pi} dz \left\{ \frac{\alpha_0}{2} + \sum_{n=1}^{\infty} [\alpha_n \cos(nz) + \beta_n \sin(nz)] \right\} \cos(mz) \\
 &= \frac{\alpha_0}{2} \underbrace{\int_{-\pi}^{\pi} dz \cos(mz)}_0 + \sum_{n=1}^{\infty} \alpha_n \underbrace{\int_{-\pi}^{\pi} dz \cos(nz) \cos(mz)}_{\pi \delta_{n,m}} + \sum_{n=1}^{\infty} \beta_n \underbrace{\int_{-\pi}^{\pi} dz \sin(nz) \cos(mz)}_0 \\
 &= \alpha_m \pi.
 \end{aligned}$$

**Remark 2:** The theorem can be generalized to **discontinuous functions** with a finite number  $\ell$  of discontinuities  $z_1, \dots, z_\ell \in [-\pi, \pi]$ :

**Example 2:** The **sign function**  $f(z) = \text{sgn}(z)$ ,

$$\text{sgn}(z) = \begin{cases} -1 & (z < 0), \\ 0 & (z = 0), \\ 1 & (z > 0), \end{cases} \quad (384)$$

has one single ( $\ell = 1$ ) discontinuity  $z_1 = 0$  within  $[-\pi, \pi]$ .



It has the Fourier series

$$\begin{aligned}
 f_S(z) &= \frac{4}{\pi} \cdot \left\{ \sin z + \frac{1}{3} \sin 3z + \frac{1}{5} \sin 5z + \dots \right\} \\
 &= \frac{4}{\pi} \cdot \sum_{n=1,3,5,\dots}^{\infty} \frac{1}{n} \sin(nz). \quad (385)
 \end{aligned}$$

In this case,  $\alpha_n = 0$  (for all  $n = 0, 1, 2, 3, \dots$ ) and

$$\beta_n = \begin{cases} \frac{4}{\pi} \cdot \frac{1}{n} & (n = 1, 3, 5, \dots), \\ 0 & (n = 2, 4, 6, \dots). \end{cases} \quad (386)$$

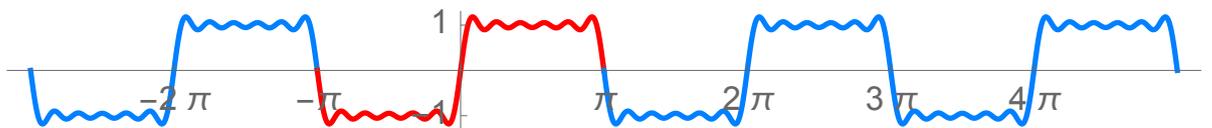


Figure 15: The truncated Fourier series  $f_S(z) = \frac{4}{\pi} \cdot \sum_{n=1,3,5,\dots}^N \frac{1}{n} \sin(nz)$  with  $N = 11$ .

## 10.3 Wave equation in 1D

The transversal displacement  $y = u(x, t)$  (in  $y$ -direction) of a violin string (which at rest covers the  $x$ -axis) satisfies the 1D wave equation (345),

$$\frac{1}{c^2} \frac{\partial^2}{\partial t^2} u(x, t) = \frac{\partial^2}{\partial x^2} u(x, t), \quad c^2 = \frac{F}{\mu}. \quad (387)$$

$F$  (in N) is the string tension and  $\mu$  (in  $\frac{\text{g}}{\text{m}}$ ) is the mass density of the string.

### 10.3.1 D'Alembert solution

The general solution of Eq. (387) is given by

$$u(x, t) = f(x - ct) + g(x + ct), \quad (388)$$

where  $f(x)$  and  $g(x)$  are arbitrary functions, with derivatives  $f'(x)$ ,  $f''(x)$  and  $g'(x)$ ,  $g''(x)$ , respectively. To see this, consider

$$\begin{aligned} \frac{\partial u(x, t)}{\partial t} &= (-c) f'(x - ct) + c g'(x + ct), \\ \Rightarrow \frac{\partial^2 u(x, t)}{\partial t^2} &= (-c)^2 f''(x - ct) + c^2 g''(x + ct) \equiv c^2 \frac{\partial^2 u(x, t)}{\partial x^2}. \end{aligned} \quad (389)$$

Plotted as a function of  $x$  for different (fixed) values of  $t$ , the term  $f(x - ct)$  describes a **rigid profile**, moving at constant **velocity**  $c$  in positive  $x$ -direction along the  $x$ -axis. Similarly,  $g(x + ct)$  describes another profile, moving in negative  $x$ -direction, cf. Fig. 16. Generally, such a motion of a rigid profile is called a **wave**.

In particular, the constant  $c = \sqrt{F/\mu}$  is the **speed of waves** on the string. E.g.,

$$c = \sqrt{\frac{100 \text{ N}}{0.010 \frac{\text{kg}}{\text{m}}}} = \sqrt{10\,000 \frac{\text{m}^2}{\text{s}^2}} = 100 \frac{\text{m}}{\text{s}}. \quad (390)$$

### 10.3.2 Initial-value conditions

A solution  $u(x, t)$  of Eq. (387) is fixed uniquely by **initial value conditions** (ICs),

$$u(x, 0) = y(x), \quad u_t(x, 0) = v(x), \quad (391)$$

where  $u_t(x, t) \equiv \frac{\partial u(x, t)}{\partial t}$ . Here,  $y(x)$  and  $v(x)$  are **two given functions**.

Using Eq. (388) for  $u(x, t)$ , these ICs read

$$f(x) + g(x) = y(x), \quad c[g'(x) - f'(x)] = v(x). \quad (392)$$

Defining  $V(x) = \int_0^x v(x') dx' + C$ , we obtain  $f(x)$  and  $g(x)$  in terms of  $y(x)$  and  $v(x)$ ,

$$f(x) = \frac{1}{2} \left[ y(x) - \frac{V(x)}{c} \right] + \frac{C}{2c}, \quad g(x) = \frac{1}{2} \left[ y(x) + \frac{V(x)}{c} \right] - \frac{C}{2c}. \quad (393)$$

The resulting solution (388) does not depend on  $C$ ,

$$\begin{aligned} u(x, t) &= \frac{1}{2} [y(x - ct) + y(x + ct)] + \frac{1}{2c} [V(x + ct) - V(x - ct)] \\ &\equiv \frac{1}{2} [y(x - ct) + y(x + ct)] + \frac{1}{2c} \int_{x-ct}^{x+ct} dx' v(x'). \end{aligned} \quad (394)$$

An example with  $v(x) = 0$ :  $u(x, t) = \frac{1}{2}[y(x - ct) + y(x + ct)]$ , is illustrated in Fig. 16.

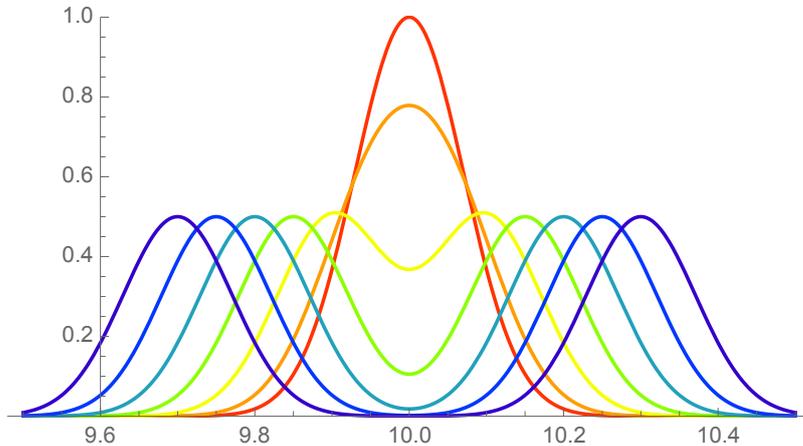


Figure 16: The solution  $u(x, t)$  to the ICs  $y(x) = e^{-(x-10)^2/0.1^2}$  and  $v(x) = 0$  for an infinitely long string (choosing the value  $c = 1$ ), plotted versus  $x$  for a few **small** times:  $t = 0$  (red),  $t = 0.05$  (orange),  $t = 0.1$  (yellow),  $t = 0.15$  (green), ...,  $t = 0.3$  (dark-blue).

### 10.3.3 Boundary conditions

A violin string, fixed at  $x = 0$  and at  $x = \ell$ , satisfies the **boundary conditions** (BCs)

$$u(0, t) = 0, \quad u(\ell, t) = 0 \quad (\text{for all } t). \quad (395)$$

More generally, the BCs can have the form

$$u(0, t) = a(t), \quad u(\ell, t) = b(t) \quad (\text{for all } t), \quad (396)$$

with two given functions  $a(t)$  and  $b(t)$ , imposing on the ICs the constraints

$$y(0) = a(0), \quad y(\ell) = b(0), \quad v(0) = \dot{a}(0), \quad v(\ell) = \dot{b}(0). \quad (397)$$

In the presence of BCs, the application of d'Alembert's solution is **not quite clear**, since then the functions  $y(x)$  and  $v(x)$  in Eq. (394) are known for  $0 \leq x \leq \ell$  only, while the argument  $|x \pm ct|$  can grow beyond any limit as  $t$  increases. E.g., in the case  $\ell = 20$ , the behavior of the wave in Fig. 16 becomes wrong as soon as  $t \approx 9.7$  when both partial waves will reach the boundaries at  $x = 0$  and at  $x = 20$ , respectively.

To find the proper continuation of  $y(x)$  onto all of  $\mathbb{R}$  [we choose  $v(x) = 0$  for simplicity], we shall now solve the wave equation directly.

### 10.3.4 Factorizing solutions

As in section 10.2.1, we first attempt to satisfy all **homogeneous** conditions [which are: the BCs (395) plus the second IC of Eq. (391), here  $v(x) = 0$ ] by a separation ansatz

$$u(x, t) = X(x)T(t). \quad (398)$$

In this case, the homogeneous conditions read

$$X(0) = 0, \quad X(\ell) = 0, \quad \dot{T}(0) = 0. \quad (399)$$

Similar steps as in section 10.2.1 (see problem 9.7) yield

$$X(x)T(t) = U \sin(k_n x) \cos(k_n ct), \quad k_n = n \frac{\pi}{\ell} \quad (n = 1, 2, 3, \dots). \quad (400)$$

All homogeneous conditions are still satisfied by any arbitrary linear combination

$$u(x, t) = \sum_{n=1}^{\infty} U_n \sin(k_n x) \cos(k_n ct). \quad (401)$$

The addition theorem  $\sin \phi \cos \chi = \frac{1}{2}[\sin(\phi - \chi) + \sin(\phi + \chi)]$  yields

$$\begin{aligned} u(x, t) &= \sum_{n=1}^{\infty} \frac{U_n}{2} \left[ \sin(k_n(x - ct)) + \sin(k_n(x + ct)) \right] \\ &\equiv \frac{1}{2} \left[ F(x - ct) + F(x + ct) \right], \end{aligned} \quad (402)$$

where  $F(x)$  is an odd function with period  $P = 2\ell$ , given for all  $x \in \mathbb{R}$  by the series

$$F(x) = \sum_{n=1}^{\infty} U_n \sin(k_n x) \quad (x \in \mathbb{R}). \quad (403)$$

Now, the first (**inhomogeneous**) IC of Eq. (391) implies that  $F(x) \equiv y(x)$  for  $0 < x < \ell$ . Consequently,  $F(x)$  is the wanted continuation of  $y(x)$ ,

$$F(x) = \begin{cases} -y(|x|) & (-\ell \leq x \leq 0) \\ y(x) & (0 \leq x \leq \ell) \end{cases}, \quad F(x \pm 2\ell) = F(x) \quad (x \in \mathbb{R}). \quad (404)$$

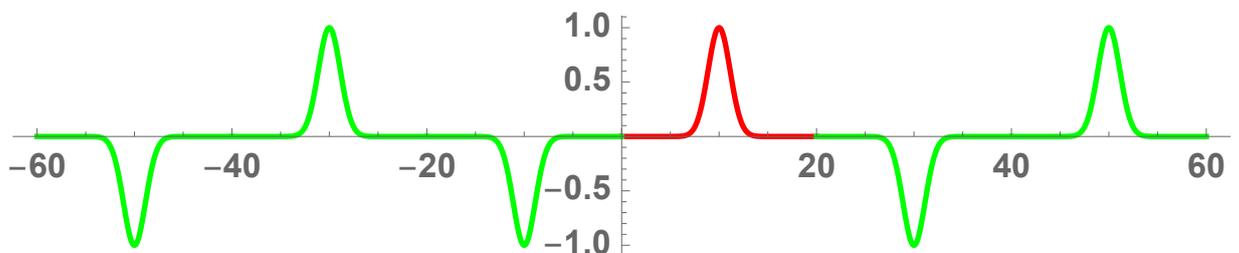


Figure 17: The periodic continuation  $F(x)$  to  $x \in \mathbb{R}$  (green) of the IC  $y(x) = be^{-(x-10)^2/a^2}$  for  $0 < x < 20$  (red). The parameters are  $a = 1.6$  and  $b = 1.0$ .

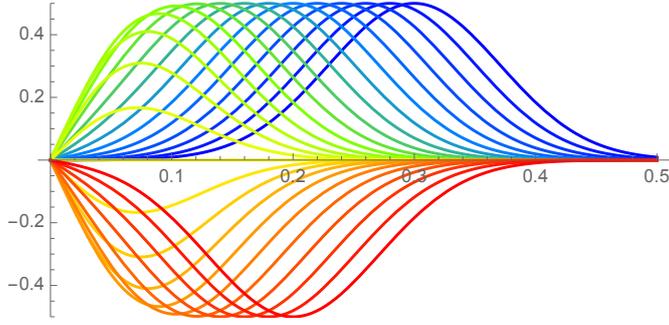


Figure 18: The solution  $u(x, t)$  to the ICs  $y(x) = e^{-(x-10)^2/0.1^2}$  and  $v(x) = 0$  for the string fixed at  $x = 0$  and at  $x = 20$  (choosing the value  $c = 1$ ), plotted versus  $x$  for the times:  $t = 9.7, 9.72, 9.74$  (blue), ...,  $t = 10.16, 10.18, 10.2$  (red).

### 10.3.5 Fourier analysis

Fixed by Eq. (403), the coefficients  $U_n$  are not needed for constructing the function  $F(x)$ . However, they yield the **spectrum** of the sound produced by the vibrating string.

For a Fourier analysis of  $F(x)$ , we consider one period,  $-\ell < x < \ell$ . Due to Eqs. (383), with  $f(z) = F(\frac{\ell}{\pi}z)$ ,  $x_0 = 0$ ,  $a = \ell$ ,  $-\pi < z < \pi$ , the coefficients are

$$\begin{aligned} U_n \equiv \beta_n &= \frac{1}{\pi} \int_{-\pi}^{\pi} dz F\left(\frac{\ell}{\pi}z\right) \sin(nz) \\ &= \frac{2}{\pi} \int_0^{\pi} dz y\left(\frac{\ell}{\pi}z\right) \sin(nz) \equiv \frac{2}{\ell} \int_0^{\ell} dx y(x) \sin(k_n x). \end{aligned} \quad (405)$$

and  $\alpha_n = 0$  for all  $n$ .

As an example, with the BCs (395), we consider a string that was picked at the time  $t = 0$ ,

$$y(x) = \begin{cases} \frac{b}{a} \cdot x & (0 \leq x \leq a), \\ b \cdot \frac{\ell-x}{\ell-a} & (a \leq x \leq \ell). \end{cases}, \quad v(x) = 0. \quad (406)$$

In this case, we have (writing  $k_n = k$ ),

$$\begin{aligned} U_n &= \frac{2}{\ell} \left\{ \frac{b}{a} \int_0^a dx \sin(kx) x + \frac{b}{\ell-a} \int_a^{\ell} dx \sin(kx) (\ell-x) \right\} \\ &= \frac{2}{\ell} \left\{ \frac{b}{a} \left[ \frac{\sin(kx)}{k^2} - \frac{x \cos(kx)}{k} \right]_0^a + \right. \\ &\quad \left. + \frac{b\ell}{\ell-a} \left[ -\frac{\cos(kx)}{k} \right]_a^{\ell} - \frac{b}{\ell-a} \left[ \frac{\sin(kx)}{k^2} - \frac{x \cos(kx)}{k} \right]_a^{\ell} \right\} \\ &= \frac{2b}{a(\ell-a)} \frac{\sin(ka)}{k^2} \\ &= \frac{2b}{a(\ell-a)} \frac{\ell^2}{\pi^2} \frac{\sin\left(\frac{n\pi}{\ell}a\right)}{n^2}. \end{aligned} \quad (407)$$

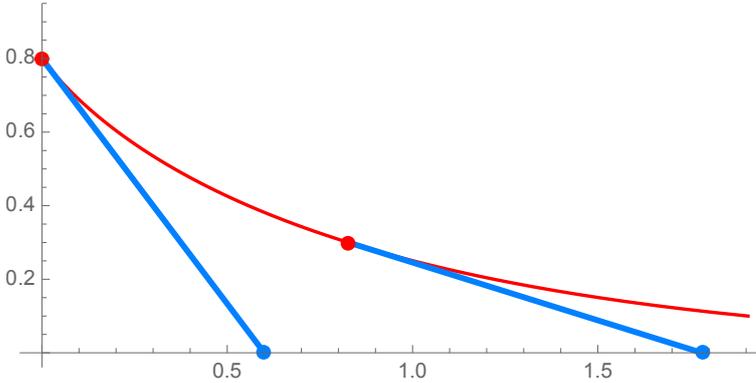
## PART III

### Numerical Solution of Differential Equations

# 11 Ordinary equations (ODE's)

## 11.1 The trailing curve as an example

On a horizontal table, one end  $A$  of a stretched thread (length  $a$ , **blue**) is being pulled along the  $x$ -axis. What is the curve  $y = y(x) > 0$  drawn by a piece of **red chalk** dragged at the other end  $B$  of the thread, when  $B$  was initially at  $(x|y) = (0|b)$  (with  $b < a$ )?



For any point  $B(x_B|y(x_B))$  on that curve, the pulling end  $A$  of the thread on the  $x$ -axis has an  $x$ -coordinate  $x_A > x_B$ , fixed by Pythagoras' theorem,

$$y(x_B)^2 + (x_A - x_B)^2 = a^2 \quad \Rightarrow \quad x_A - x_B = \sqrt{a^2 - y(x_B)^2}. \quad (408)$$

In particular, since the pulling thread lies **tangential** to the curve in  $B$ , it has the slope

$$y'(x_B) = -\frac{y(x_B)}{x_A - x_B}, \quad (409)$$

where  $y'(x)$  is the derivative of the wanted function  $y(x)$ . Using Eq. (408) for  $x_A - x_B$  and replacing  $x_B$  with  $x$ , we obtain a **non-linear ODE** for  $y(x)$ ,

$$y'(x) = -\frac{y(x)}{\sqrt{a^2 - y(x)^2}}. \quad (410)$$

Being non-linear, this ODE cannot be solved by the strategy discussed in section 8.2. Before we shall solve it exactly in section 11.4 (using a different strategy), we here want to use it for demonstrating the performance of **approximate numerical** techniques.

## 11.2 Finite difference methods (FDMs)

### 11.2.1 Euler's method

Instead of treating  $x$  as a **continuous** variable, we choose a finite **step size**  $h > 0$  and introduce  $N + 1$  equidistant **discrete mesh points**

$$x_n = nh \quad (n = 0, 1, 2, \dots, N). \quad (411)$$

The definition of the derivative,  $y'(x) \equiv \lim_{h \rightarrow 0} \frac{y(x+h) - y(x)}{h}$ , for **finite** but sufficiently small  $h > 0$  yields an approximation for a **finite difference**,

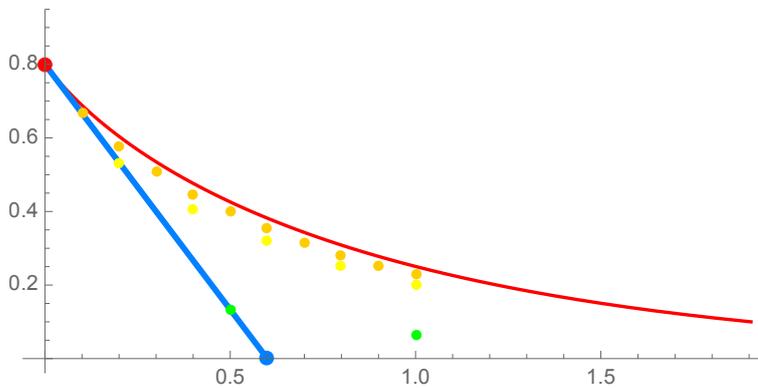
$$y(\underbrace{x_n + h}_{x_{n+1}}) - y(x_n) \approx h \cdot y'(x_n). \quad (412)$$

Using for  $y'(x_n)$  the ODE Eq. (410), we obtain a **recursion formula**,

$$y(x_{n+1}) \approx y(x_n) - h \cdot \frac{y(x_n)}{\sqrt{a^2 - y(x_n)^2}}. \quad (413)$$

Starting with the initial value  $y(x_0) = b$ , it successively yields approximate values for  $y_h(x_1)$ ,  $y_h(x_2)$ , etc. This FDM is known as **Euler's method**.

$x$	0.0	0.1	0.2	0.3	0.4	0.5	0.6	0.7	0.8	0.9	1.0
$y_{h=0.5}$	0.8					0.133					0.066
$y_{h=0.2}$	0.8		0.533		0.407		0.318		0.251		0.199
$y_{h=0.1}$	0.8	0.667	0.577	0.507	0.448	0.398	0.354	0.316	0.283	0.254	0.227
$y_{\text{exact}}$	0.8	0.689	0.604	0.535	0.476	0.426	0.382	0.343	0.309	0.278	0.251



Here is a **MATLAB code**, using the values  $a = 1$ ,  $b = 0.8$ ,  $h = 0.01$  and  $N = 100$ :

```

a=1;
b=0.8;
dx=0.01;
N=100;
X=zeros(1,N+1);           % zero matrix 'X' with 1 row and N+1 columns
Y=zeros(1,N+1);           % zero matrix 'Y' with 1 row and N+1 columns
x=0;
y=b;
Y(1)=y;
for (n=2:N+1)
    yp=-y/sqrt(a.^2-y.^2);
    y=y+yp*dx;
    x=x+dx;
    [x,y]                % print the pair (x,y) on the screen
    X(n)=x;
    Y(n)=y;
end
plot(X,Y)
axis equal
grid on

```

## 11.2.2 Runge-Kutta methods

## 11.3 The MATLAB tool “ode23”

### 11.3.1 First-order ODEs

”ode23” numerically solves **first-order ODEs** (or, more generally, systems of several first-order ODEs). Generally, such a (typically **non-linear**) ODE has the form

$$y'(x) = f(x, y(x)), \quad (414)$$

with a given function  $f(x, y)$  of two variables.

• **Example 1:** A **fantasy ODE** and the corresponding function  $f$  are

$$y'(x) = -2xy(x)^2, \quad f(x, y) = -2xy^2. \quad (415)$$

In passing, we note that the **exact** solution is (check!)

$$y(x) = \frac{1}{a + x^2} \quad (a \in \mathbb{R}). \quad (416)$$

To apply ”ode23” to such an ODE, we first define the function  $f(x, y)$  in a separate m-file. For the above example, this file (introducing the name ”fantasy” for  $f$ ) could read:

```
% fantasy.m
% Function for the fantasy ODE
function out=fantasy(x,y)
out=-2*x*y.^2
```

• **Example 2:** For the ODE (410) of the **trailing curve**,  $f(x, y)$  does not depend on  $x$ ,

$$y'(x) = -\frac{y(x)}{\sqrt{a^2 - y(x)^2}}, \quad f(x, y) = -\frac{y}{\sqrt{a^2 - y^2}}. \quad (417)$$

In this case, the separate m-file (with the name ”trailing” for  $f$ ) could read:

```
% trailing.m
% Function for the trailing-curve ODE
function out=trailing(x,y)
a=1.;
out=-y/sqrt(a.^2-y.^2)
```

After having defined the ODE this way, we can call ”ode23” by the command:

```
[X Y]=ode23('trailing', [0 3], 0.8);
```

The first argument ’trailing’ calls the function  $f(x, y)$  that defines the ODE, the second one [0 3] fixes the  $x$ -interval  $x_0 (= 0) \leq x \leq x_f (= 3)$  for which the solution is wanted. The third argument 0.8 gives the initial value  $y(x_0)$ , here  $y(0) = b (= 0.8)$ .

The output [X Y] gives two vectors X and Y, containing, respectively,  $x$ - and  $y$ -values of the solution  $y(x)$  found for the ODE. This result can be plotted by

```
plot(X,Y)
axis equal
grid on
```

### 11.3.2 Higher-order ODEs

For completeness, we apply "ode23" to a system of two first-order ODEs:

The exact equation of motion for a mathematical pendulum is a **second-order** ODE,

$$\ddot{\phi}(t) = -\frac{g}{\ell} \sin \phi(t). \quad (418)$$

Here,  $g = 9.81 \frac{\text{m}}{\text{s}^2}$  (gravitational acceleration on the surface of Earth) and  $\ell$  is the length of the pendulum.  $\phi(t)$  is, as a function of time  $t$ , the **angular displacement** from the pendulum's equilibrium position ( $\phi = 0$ ).

Writing  $\phi(t) = y_1(t)$  and  $\dot{\phi}(t) = y_2(t)$ , we obtain a **system of two first-order** ODEs,

$$\begin{aligned} \dot{y}_1(t) &= y_2(t), \\ \dot{y}_2(t) &= -\frac{g}{\ell} \sin(y_1(t)). \end{aligned} \quad (419)$$

Combining the two solutions to a **vector**,  $\begin{pmatrix} y_1(t) \\ y_2(t) \end{pmatrix} = \mathbf{y}(t)$ , we formally have

$$\dot{\mathbf{y}}(t) = \mathbf{f}(t, \mathbf{y}(t)), \quad (420)$$

where the vector function  $\mathbf{f}(t, \mathbf{y})$  here does not depend on  $t$ , but only on  $\mathbf{y} = (y_1, y_2)$ ,

$$\mathbf{f}(t, \mathbf{y}) \equiv \begin{pmatrix} f_1(t, \mathbf{y}) \\ f_2(t, \mathbf{y}) \end{pmatrix} = \begin{pmatrix} y_2 \\ -\frac{g}{\ell} \sin y_1 \end{pmatrix}. \quad (421)$$

The small-amplitude ( $\phi_{\max} \ll \pi$ ) oscillation period  $T_{\text{sa}} = 2\pi\sqrt{\ell/g}$  is exactly 1 s when we choose  $\ell = 24.85$  cm or  $g/\ell = 39.48 \text{ s}^{-2}$ . Then, the **vector function**  $\mathbf{f}(t, \mathbf{y})$  is defined by

```
% pend.m
function out=pend(t,y)
out=zeros(2,1);
out(1)=y(2);
out(2)=-39.48*sin(y(1));
```

Notice that the input  $\mathbf{y}$  is a vector, with two components  $\mathbf{y}(1)$  and  $\mathbf{y}(2)$ . So is the output.

Now, calling "ode23" using the time interval  $0 \leq t \leq 3$  s and the initial conditions  $y_1(0) \equiv \phi(0) = 0.15 \ll \pi$  (**small amplitude**) and  $y_2(0) \equiv \dot{\phi}(0) = 0$ ,

```
[T Y]=ode23('pend', [0 3], [0.15 0]);
```

we obtain an almost perfect sine curve with period  $\approx 1$  s and amplitude 0.15:

```
plot(X,Y(:,1))
grid on
```

In contrast, the initial conditions  $y_1(0) \equiv \phi(0) = 3.13 \approx \pi$  (**almost maximum possible amplitude**) and  $y_2(0) \equiv \dot{\phi}(0) = 0$ ,

```
[T Y]=ode23('pend', [0 3], [3.13 0]);
```

yield a curve totally different from a sine, and a period  $T$  considerably larger than 1 s.

**Entertaining example:** The Lorenz attractor.

## 11.4 Exact trailing curve

Our first-order ODE Eq. (410) is a **separable-variables** equation: It can be written as

$$\frac{dy}{dx} = -\frac{y}{\sqrt{a^2 - y^2}} \quad \Rightarrow \quad -\frac{\sqrt{a^2 - y^2}}{y} dy = dx. \quad (422)$$

Now, only  $y$ 's appear on the LHS while the RHS has  $x$ 's only !

A simpler example for this kind of ODE is  $y'(x) = \frac{5y(x)}{x}$ . In this case we write

$$\frac{dy}{dx} = \frac{5y(x)}{x} \quad \Rightarrow \quad \frac{1}{y} dy = \frac{5}{x} dx. \quad (423)$$

Formally integrating on the left side over  $y$  and on the right side over  $x$  yields

$$\ln y + C_1 = 5 \ln x + C_2. \quad (424)$$

There is only one independent constant,  $C_2 - C_1 = C$ .

Eventually, resolving for  $y = y(x)$  yields the general solution

$$y(x) = e^{5 \ln x + C} = ax^5, \quad a = e^C. \quad (425)$$

By checking, we confirm that this is in fact a solution to  $y'(x) = \frac{5y(x)}{x}$ , for all  $a \in \mathbb{R}$ .

Now, we apply the same integration technique on both sides of Eq. (422), to obtain<sup>8</sup>

$$a \ln \frac{a + \sqrt{a^2 - y^2}}{y} - \sqrt{a^2 - y^2} + C_1 = x + C_2. \quad (426)$$

We note in passing:  $\ln(u + \sqrt{u^2 - 1}) \equiv \operatorname{arcosh}(u)$ , where  $u = \frac{a}{y} \geq 1$ .

For plotting the (invertible!) function  $y = y(x)$ , it is sufficient to resolve for  $x$ ,

$$x(y) = a \ln \frac{a + \sqrt{a^2 - y^2}}{y} - \sqrt{a^2 - y^2} - C \quad (C = C_2 - C_1). \quad (427)$$

To satisfy the starting condition  $y(0) = b$  or, equivalently,  $x(b) = 0$ , we choose

$$C = a \ln \frac{a + \sqrt{a^2 - b^2}}{b} - \sqrt{a^2 - b^2}. \quad (428)$$

<sup>8</sup>We check Eq. (426) by differentiating it:

$$\begin{aligned} \frac{d}{dy} \left\{ a \ln \frac{a + \sqrt{a^2 - y^2}}{y} - \sqrt{a^2 - y^2} \right\} &= a \frac{y}{a + \sqrt{\dots}} \cdot \frac{y \cdot \frac{-2y}{2\sqrt{\dots}} - 1 \cdot (a + \sqrt{\dots})}{y^2} - \frac{-2y}{2\sqrt{\dots}} \\ &= a \frac{y(a - \sqrt{\dots})}{y^2} \cdot \frac{-1}{\sqrt{\dots}} - \frac{a}{y} + \frac{y}{\sqrt{\dots}} \\ &= -\frac{a^2}{y\sqrt{\dots}} + \frac{a}{y} - \frac{a}{y} + \frac{y^2}{y\sqrt{\dots}} = -\frac{\sqrt{a^2 - y^2}}{y}. \end{aligned}$$

## 12 Partial differential equations (PDE's)

We here consider, as an example, the PDE

$$-p \nabla^2 u(\mathbf{r}) = s(\mathbf{r}) \quad (429)$$

for the **wanted function**  $u(\mathbf{r})$  (a scalar field) in  $n = 2$  or  $n = 3$  dimensions,  $\mathbf{r} \in \mathbb{R}^n$ .  $p > 0$  is a given number (**parameter**) and  $s(\mathbf{r})$  is a given function (**source**).

Eq. (429) represents the time-independent (stationary) case of any one of the PDEs

$$\frac{\partial}{\partial t} \rho(\mathbf{r}, t) = D \nabla^2 \rho(\mathbf{r}, t) + s(\mathbf{r}, t) \quad (\text{diffusion equation}), \quad (430)$$

$$c\mu \frac{\partial}{\partial t} T(\mathbf{r}, t) = \lambda \nabla^2 T(\mathbf{r}, t) + s(\mathbf{r}, t) \quad (\text{heat conduction equation}), \quad (431)$$

$$\nabla^2 \Phi(\mathbf{r}, t) = -\frac{1}{\epsilon_0} \rho(\mathbf{r}, t) \quad (\text{Poisson's equation}), \quad (432)$$

when parameter  $p$ , source  $s(\mathbf{r})$ , and solution  $u(\mathbf{r})$  are interpreted properly:

PDE	$p$	$u(\mathbf{r})$	$s(\mathbf{r})$
Diffusion	$D$	$\rho(\mathbf{r})$	$s(\mathbf{r})$
Heat conduction	$\lambda$	$T(\mathbf{r})$	$s(\mathbf{r})$
Poisson	$\epsilon_0$	$\Phi(\mathbf{r})$	$\rho(\mathbf{r})$

In the **diffusion equation** ( $p = D$ ),  $u = \rho$  is the **mass density** of the diffusing substance, in the **heat conduction eqn.** ( $p = \lambda$ ),  $u = T$  is the **temperature distribution**, while in **Poisson's eqn.** ( $p = \epsilon_0$ ), where the source is given by the **electric charge density**,  $s = \rho$ , the wanted function is the **electrostatic potential**  $u = \Phi$ .

**Time-dependent** problems (including waves) will be considered later.

### 12.1 Boundary conditions

Typically, a solution  $u(\mathbf{r})$  of Eq. (429) is wanted for  $\mathbf{r} \in \Omega$ , where  $\Omega \subset \mathbb{R}^n$  is a **given region** in 3D or 2D space. In addition, on the **surface**  $\partial\Omega$  of  $\Omega$  (in the case  $n = 2$ ,  $\partial\Omega$  is the **rim** of  $\Omega$ ),  $u(\mathbf{r})$  shall satisfy the **boundary conditions**

$$(D) \quad u(\mathbf{r}) = f(\mathbf{r}) \quad (\text{for all } \mathbf{r} \in \partial\Omega_1), \quad (433)$$

$$(N) \quad \mathcal{N}(\mathbf{r}) \cdot \nabla u(\mathbf{r}) = g(\mathbf{r}) \quad (\text{for all } \mathbf{r} \in \partial\Omega_2). \quad (434)$$

Here,  $\partial\Omega$  is divided into two pieces  $\partial\Omega_1$  and  $\partial\Omega_2$  with zero overlap,  $\partial\Omega_1 \cap \partial\Omega_2 = \emptyset$ .

$\mathcal{N}(\mathbf{r})$  is the unit normal vector at  $\mathbf{r} \in \partial\Omega$ , pointing outward of  $\Omega$ .

$f(\mathbf{r})$  and  $g(\mathbf{r})$  are given functions for  $\mathbf{r} \in \partial\Omega_1$  and  $\mathbf{r} \in \partial\Omega_2$ , respectively.

**Interpretation:** In the case of **Poisson's equation**, the **Dirichlet BC (D)** fixes the electrostatic potential  $\Phi(\mathbf{r}) = f(\mathbf{r})$  on the part  $\partial\Omega_1$  of the surface, while the **Neumann BC (N)** fixes the surface charge density  $\sigma(\mathbf{r})$  on  $\partial\Omega_2$ .

**Theorem:** A solution  $u(\mathbf{r})$  of Eq. (429) is fixed **uniquely** for all  $\mathbf{r} \in \Omega$  by the conditions (433) and (434).

• For the proof, we need **Green's first identity**:

Consider two scalar fields  $\phi(\mathbf{r})$  and  $\psi(\mathbf{r})$ .

Applying Gauß's theorem to the vectorfield  $\mathbf{F}(\mathbf{r}) = \phi(\mathbf{r})\nabla\psi(\mathbf{r})$  yields

**Green's first identity**,

$$\oint_{\partial\Omega} d\mathbf{S} \cdot [\phi\nabla\psi] = \int_{\Omega} d^3r \left[ \phi\nabla^2\psi + (\nabla\phi) \cdot (\nabla\psi) \right]. \quad (435)$$

Interchanging here  $\phi$  and  $\psi$  and subtracting the resulting equation yields

**Green's second identity**,

$$\oint_{\partial\Omega} d\mathbf{S} \cdot [\phi\nabla\psi - \psi\nabla\phi] = \int_{\Omega} d^3r \left[ \phi\nabla^2\psi - \psi\nabla^2\phi \right]. \quad (436)$$

**Proof:** Let  $u_1(\mathbf{r})$  and  $u_2(\mathbf{r})$  be two solutions satisfying (433) and (434). Their difference  $w(\mathbf{r})$ ,

$$w(\mathbf{r}) = u_1(\mathbf{r}) - u_2(\mathbf{r}), \quad (437)$$

satisfies (433) and (434) with  $f = g = 0$ . Consequently,

$$\begin{aligned} \oint_{\partial\Omega} d\mathbf{S} \cdot [w(\mathbf{r})\nabla w(\mathbf{r})] &\equiv \oint_{\partial\Omega} d^{n-1}r w(\mathbf{r}) [\mathcal{N}(\mathbf{r}) \cdot \nabla w(\mathbf{r})] \\ &= \int_{\partial\Omega_1} d^{n-1}r 0 [\mathcal{N}(\mathbf{r}) \cdot \nabla w(\mathbf{r})] + \int_{\partial\Omega_2} d^{n-1}r w(\mathbf{r}) 0 \\ &= 0. \end{aligned} \quad (438)$$

Moreover, for  $\mathbf{r} \in \Omega$ ,  $w(\mathbf{r})$  satisfies Laplace's Equation,

$$\nabla^2 w(\mathbf{r}) = 0 \quad (\text{for all } \mathbf{r} \in \Omega). \quad (439)$$

Consequently, writing Green's first identity with  $\phi(\mathbf{r}) = \psi(\mathbf{r}) = w(\mathbf{r})$ , we obtain

$$0 \equiv \oint_{\partial\Omega} d\mathbf{S} \cdot [w(\mathbf{r})\nabla w(\mathbf{r})] = \int_{\Omega} d^n r |\nabla w(\mathbf{r})|^2. \quad (440)$$

Since  $w(\mathbf{r})$  is a continuous function, the RHS then implies that

$$|\nabla w(\mathbf{r})|^2 = 0 \quad \Leftrightarrow \quad u_1(\mathbf{r}) - u_2(\mathbf{r}) = C \quad (\text{for all } \mathbf{r} \in \Omega), \quad (441)$$

with a constant  $C$ . Since  $u_1(\mathbf{r})$  and  $u_2(\mathbf{r})$  both satisfy the first BC (433), we have  $C = 0$ , provided that  $\partial\Omega_1 \neq \emptyset$ .

## 12.2 Finite difference methods (FDM)

We here consider the PDE, Eq. (429), in **two dimensions** ( $n = 2$ ) and set  $p = 1$ ,

$$\nabla^2 u(\mathbf{r}) \equiv \frac{\partial^2 u(x, y)}{\partial x^2} + \frac{\partial^2 u(x, y)}{\partial y^2} = -s(x, y). \quad (442)$$

Given some region  $\Omega$  in the  $xy$ -plane and a function  $f(\mathbf{r})$  defined on the rim  $\partial\Omega$  of  $\Omega$ , we want the solution  $u(x, y)$  that is fixed by the Dirichlet BC

$$u(\mathbf{r}) = f(\mathbf{r}), \quad \mathbf{r} \in \partial\Omega. \quad (443)$$

### 12.2.1 Transformation into a FD equation

Using a small **step size**  $h > 0$ , we introduce discrete **mesh points** in the  $xy$ -plane,

$$\mathbf{r}_{nm} \equiv (x_n, y_m) = (nh, mh) \quad (n, m \in \mathbb{Z}). \quad (444)$$

Writing  $\frac{\partial u(x, y)}{\partial x} = u_x(x, y)$ , the definition of derivatives implies

$$\begin{aligned} \frac{\partial^2 u(x, y)}{\partial x^2} &= \lim_{h \rightarrow 0} \frac{u_x(x, y) - u_x(x - h, y)}{h} \\ &= \lim_{h \rightarrow 0} \frac{\frac{u(x+h, y) - u(x, y)}{h} - \frac{u(x, y) - u(x-h, y)}{h}}{h} \\ &= \lim_{h \rightarrow 0} \frac{u(x+h, y) - 2u(x, y) + u(x-h, y)}{h^2}. \end{aligned} \quad (445)$$

Writing  $\frac{\partial^2 u(x, y)}{\partial x^2} = u_{xx}(x, y)$ , we obtain for sufficiently small  $h > 0$  the approximation

$$\begin{aligned} u_{xx}(x_n, y_m) &\approx \frac{u(x_{n+1}, y_m) - 2u(x_n, y_m) + u(x_{n-1}, y_m)}{h^2} \\ &\equiv \frac{u_{n+1, m} - 2u_{n, m} + u_{n-1, m}}{h^2}, \end{aligned} \quad (446)$$

where we have introduced the notation  $u(x_n, y_m) = u_{n, m}$  in the second step. With the corresponding approximation for  $u_{yy}(x_n, y_m)$ , the **FD version of the PDE** reads

$$\frac{u_{n+1, m} - 2u_{n, m} + u_{n-1, m}}{h^2} + \frac{u_{n, m+1} - 2u_{n, m} + u_{n, m-1}}{h^2} = -s_{n, m}. \quad (447)$$

### 12.2.2 Iteration method

Resolving Eq. (447) for  $u_{n, m}$ , we obtain a set of **iteration equations**,

$$u_{n, m} = \frac{1}{4} \left[ u_{n+1, m} + u_{n-1, m} + u_{n, m+1} + u_{n, m-1} \right] + \frac{h^2}{4} s_{n, m}. \quad (448)$$

Using on the RHS approximate **trial values**  $u_{k, \ell} \equiv u(x_k, y_\ell)$  that cover the region  $\Omega$ , the LHS yields new values. Provided that these "new" values are a better approximation to the wanted solution, we may continue, using them as the new "old" ones, etc.

**Remark:** In the case  $s = 0$  (as in the following example), each "new" value  $u_{nm}$  is simply the **average of the four neighboring "old" values!**

**Example 1:** Revisiting the **heat-flow example** with  $s(x, y) = 0$  (Laplace's equation) of section 10.2, Eq. (362) there, we here consider the special case  $b = a$  of the square

$$\Omega = \left\{ (x, y) \mid 0 \leq x, y \leq a \right\}. \quad (449)$$

Writing  $T(x, y) = u(x, y)$ , the BCs (363) and (364) now read

$$u(0, y) = u(x, 0) = u(a, y) = 0, \quad u(x, a) = T_0. \quad (450)$$

Choosing the step size  $h = \frac{a}{3}$ , we have as few as  $N = 4$  mesh points  $\mathbf{r}_{n,m} = (x_n, y_m)$  within  $\Omega$  (and twelve mesh points on the rim  $\partial\Omega$ ), with the coordinates (see the figure)

$$\begin{aligned} \mathbf{r}_1 &\equiv \mathbf{r}_{12} = \left(\frac{a}{3}, \frac{2a}{3}\right), & \mathbf{r}_2 &\equiv \mathbf{r}_{22} = \left(\frac{2a}{3}, \frac{2a}{3}\right), \\ \mathbf{r}_3 &\equiv \mathbf{r}_{11} = \left(\frac{a}{3}, \frac{a}{3}\right), & \mathbf{r}_4 &\equiv \mathbf{r}_{21} = \left(\frac{2a}{3}, \frac{a}{3}\right). \end{aligned} \quad (451)$$

Here, we have re-labeled the points  $\mathbf{r}_{nm}$  within  $\Omega$  as  $\mathbf{r}_\nu$ , with a **single index**  $\nu \in \{1, \dots, N\}$ .

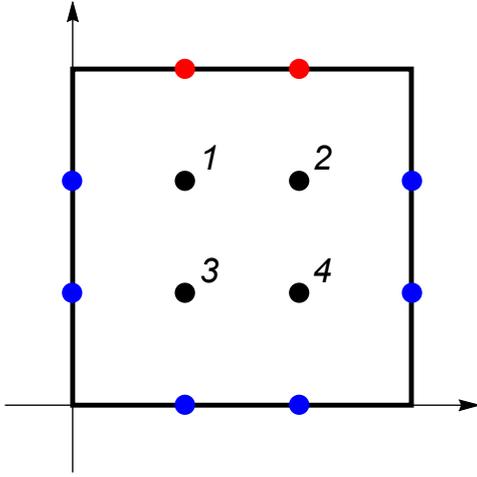


Figure 19: The region  $\Omega$ , Eq. (449), in the  $xy$ -plane, with four mesh points inside (black). On the boundary  $\partial\Omega$ , the blue/red dots correspond to the first/second BC of Eq. (450).

Writing  $u(\mathbf{r}_\nu) = u_\nu$ , the four relevant **iteration equations** (448) read

$$\begin{aligned} u_1 &= \frac{1}{4}(u_2 + u_3 + T_0), \\ u_2 &= \frac{1}{4}(u_1 + u_4 + T_0), \\ u_3 &= \frac{1}{4}(u_1 + u_4), \\ u_4 &= \frac{1}{4}(u_2 + u_3). \end{aligned} \quad (452)$$

We write them in matrix notation,  $u = Au + c$ , with

$$A = \frac{1}{4} \begin{pmatrix} 0 & 1 & 1 & 0 \\ 1 & 0 & 0 & 1 \\ 1 & 0 & 0 & 1 \\ 0 & 1 & 1 & 0 \end{pmatrix}, \quad c = \frac{T_0}{4} \begin{pmatrix} 1 \\ 1 \\ 0 \\ 0 \end{pmatrix}. \quad (453)$$

Choosing  $T_0 = 80$  and starting with the **trial values**  $u^{(0)} = (40, 40, 20, 20)^T$ , the first two iterations  $u^{(k+1)} = Au^{(k)} + c$  (with  $k = 0$  and  $k = 1$ ) successively yield

$$u^{(1)} = \begin{pmatrix} 15 \\ 15 \\ 15 \\ 15 \end{pmatrix} + \begin{pmatrix} 20 \\ 20 \\ 0 \\ 0 \end{pmatrix} = \begin{pmatrix} 35 \\ 35 \\ 15 \\ 15 \end{pmatrix}, \quad u^{(2)} = \begin{pmatrix} 12.5 \\ 12.5 \\ 12.5 \\ 12.5 \end{pmatrix} + \begin{pmatrix} 20 \\ 20 \\ 0 \\ 0 \end{pmatrix} = \begin{pmatrix} 32.5 \\ 32.5 \\ 12.5 \\ 12.5 \end{pmatrix}. \quad (454)$$

In this **simple example**, we can guess that this iteration procedure converges toward

$$u = \lim_{k \rightarrow \infty} u^{(k)} = \begin{pmatrix} 30 \\ 30 \\ 10 \\ 10 \end{pmatrix}, \quad (455)$$

which is the obvious solution of the vector equation  $u = Au + c$ . For the **general case**, the convergence behavior of this iteration strategy is studied in section 12.2.3 below.

The exact solution  $u(x, y) = T(x, y)$  is given by Eq. (374) of section 10.2, with  $b = a$ ,

$$T(x, y) = \frac{4T_0}{\pi} \sum_{n=1,3,5,\dots} \frac{1}{n \sinh(n\pi)} \sin\left(\frac{n\pi}{a} x\right) \sinh\left(\frac{n\pi}{a} y\right). \quad (456)$$

Numerical summation of the terms  $n = 1, 3, 5, \dots, 199$  yields the values

$$\begin{aligned} u_1 &\equiv T\left(\frac{a}{3}, \frac{2a}{3}\right) = 30.45747 = u_2, \\ u_3 &\equiv T\left(\frac{a}{3}, \frac{a}{3}\right) = 9.54253 = u_4. \end{aligned} \quad (457)$$

Our present result (455) is **not exact**, since we have used a finite (and very rough) step size  $h = \frac{a}{3}$ . Slightly refining the step size to  $h = \frac{a}{4}$  would more than double the number  $N$  of mesh points within  $\Omega$ , see Fig. 20.

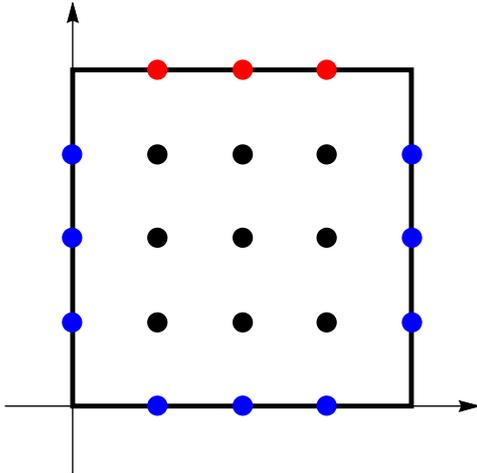


Figure 20: The region  $\Omega$ , Eq. (449), in the  $xy$ -plane, with nine mesh points inside (black), cf. Fig. 19, corresponding to the slightly refined step size  $h = \frac{a}{4}$ .

### 12.2.3 Theory

Using in Eq. (448),  $u = Au + c$ , a trial vector  $u^{(0)}$ , we obtain to 1st and 2nd orders

$$u^{(1)} = Au^{(0)} + c, \quad (458)$$

$$\begin{aligned} u^{(2)} &= A(Au^{(0)} + c) + c \\ &\equiv A^2u^{(0)} + A^1c + A^0c \end{aligned} \quad (459)$$

(where  $A^0 = 1$  denotes the unit matrix). Generally, to  $k$ th order, this expression becomes

$$u^{(k)} = A^k u^{(0)} + \sum_{j=0}^{k-1} A^j c. \quad (460)$$

As the  $N \times N$ -matrix  $A$  is **symmetric**,  $A_{\mu\nu} = A_{\nu\mu}$ ,

$$A_{\mu\nu} = \begin{cases} \frac{1}{4} & \text{(when site } \mu \text{ is neighboring site } \nu), \\ 0 & \text{(otherwise),} \end{cases} \quad (461)$$

there is a **basis set**  $\{v_1, \dots, v_N\}$  of  $N$  (pairwise orthogonal) **eigenvectors**  $v_\nu$  of  $A$ ,

$$Av_\nu = \lambda_\nu v_\nu \quad (\nu = 1, \dots, N), \quad (462)$$

with **real-valued** eigenvalues  $\lambda_\nu \in \mathbb{R}$ . Consequently, we may expand

$$u^{(0)} = \sum_{\nu=1}^N \alpha_\nu v_\nu, \quad c = \sum_{\nu=1}^N \gamma_\nu v_\nu, \quad (463)$$

to obtain in Eq. (460)

$$\begin{aligned} u^{(k)} &= A^k \left( \sum_{\nu=1}^N \alpha_\nu v_\nu \right) + \sum_{j=0}^{k-1} A^j \left( \sum_{\nu=1}^N \gamma_\nu v_\nu \right) \\ &= \sum_{\nu=1}^N \left( \alpha_\nu \lambda_\nu^k + \gamma_\nu \sum_{j=0}^{k-1} \lambda_\nu^j \right) v_\nu. \end{aligned} \quad (464)$$

Since  $|\lambda_\nu| < 1$  for  $\nu = 1, \dots, N$  (see Theorem below), this expression converges for  $k \rightarrow \infty$ , and the solution of Eq. (448) is given by the limit

$$u = \lim_{k \rightarrow \infty} u^{(k)} = \sum_{\nu=1}^N \frac{\gamma_\nu}{1 - \lambda_\nu} v_\nu. \quad (465)$$

**Theorem:** Each eigenvalue  $\lambda$  of the matrix  $A$  has the property

$$|\lambda| < 1. \quad (466)$$

**Proof:** Notice that the sum  $S_\nu$  of the elements in the  $\nu$ -th row of  $A$ ,

$$S_\nu = \sum_{\rho=1}^N A_{\nu\rho}, \quad (467)$$

has for all  $\nu$  the property  $|S_\nu| \leq 1$ . In particular,  $|S_\mu| < 1$  for at least one  $\mu \in \{1, \dots, N\}$ . (This is so, since a point  $\mathbf{r}_{nm}$  inside  $\Omega$  cannot have more than four next neighbors inside  $\Omega$ . Moreover, all the points next to the boundary  $\partial\Omega$  have less than four such neighbors.) If there is an eigenvector  $v \equiv (x_1, \dots, x_N)$  with  $|x_1| = |x_2| = \dots = |x_N| = 1$ , we have

$$|\lambda| = |\lambda x_\mu| = |(Av)_\mu| = \left| \sum_{\rho=1}^N A_{\mu\rho} x_\rho \right| \leq \sum_{\rho=1}^N A_{\mu\rho} |x_\rho| = S_\mu < 1, \quad (468)$$

where we have utilized, that  $A_{\nu\rho} \geq 0$  for all  $\nu$  and  $\rho$ . Any other eigenvector  $v \equiv (x_1, \dots, x_N)$  can be normalized in such a way that  $x_\nu = 1$  for some  $\nu$ ,  $|x_\rho| \leq 1$  for  $\rho \neq \nu$ , and  $|x_\rho| < 1$  for at least one  $\rho \neq \nu$ . Then,

$$|\lambda| = |\lambda x_\nu| = |(Av)_\nu| = \left| \sum_{\rho=1}^N A_{\nu\rho} x_\rho \right| \leq \sum_{\rho=1}^N A_{\nu\rho} |x_\rho| < S_\nu \leq 1. \quad (469)$$

**Example 2:** We now consider a problem with  $s(x, y) = 1$ ,

$$\frac{\partial^2 u(x, y)}{\partial x^2} + \frac{\partial^2 u(x, y)}{\partial y^2} = -1. \quad (470)$$

**Aluminium**, e.g., has the heat conductivity  $\lambda = 0.23 \frac{\text{kJ}}{\text{s}\cdot\text{m}\cdot\text{K}}$ . Choosing a uniform heating  $s = 0.23 \frac{\text{kJ}}{\text{s}\cdot\text{m}^3}$ , the heat flow equation  $c\mu \frac{\partial T}{\partial t} = \lambda \nabla^2 T + s$  in the stationary case reads

$$\nabla^2 T(\mathbf{r}) = -1 \frac{\text{K}}{\text{m}^2}. \quad (471)$$

For comparison with the FEM treatment in section 12.3.4 below, we choose Dirichlet BCs,  $u(\mathbf{r}) = 0$  for  $\mathbf{r} \in \partial\Omega$ , with respect to the region  $\Omega$  shown in Fig. 21. In this case, the analytical method (**separation ansatz**) of section 10.2 does **not apply!**

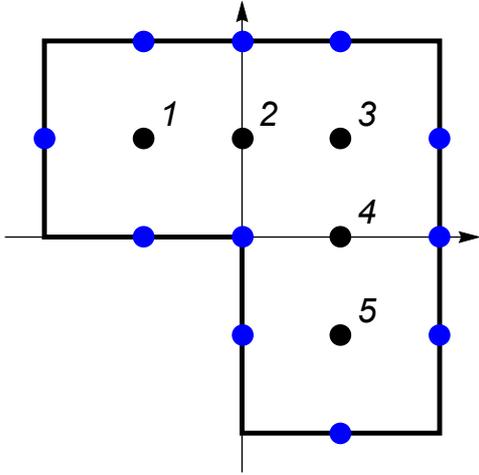


Figure 21: The region  $\Omega$  in the  $xy$ -plane, with  $\mathbf{r}_3 = (\frac{1}{2} | \frac{1}{2})$  (cf. Fig. 23 in section 12.3.4).

Now, since  $s(x, y) \neq 0$ , the step size  $h = \frac{1}{2}$  enters the relevant **iteration equations** (448),

$$\begin{aligned} u_1 &= \frac{1}{4}(u_2 + h^2), \\ u_2 &= \frac{1}{4}(u_1 + u_3 + h^2), \\ u_3 &= \frac{1}{4}(u_2 + u_4 + h^2), \\ u_4 &= \frac{1}{4}(u_3 + u_5 + h^2), \\ u_5 &= \frac{1}{4}(u_4 + h^2). \end{aligned} \quad (472)$$

In matrix notation, these equations are summarized to read  $u = Au + c$ , with

$$A = \frac{1}{4} \begin{pmatrix} 0 & 1 & 0 & 0 & 0 \\ 1 & 0 & 1 & 0 & 0 \\ 0 & 1 & 0 & 1 & 0 \\ 0 & 0 & 1 & 0 & 1 \\ 0 & 0 & 0 & 1 & 0 \end{pmatrix}, \quad c = \frac{h^2}{4} \begin{pmatrix} 1 \\ 1 \\ 1 \\ 1 \\ 1 \end{pmatrix}. \quad (473)$$

The eigenvectors  $v_1, v_2, \dots, v_5$  of the symmetric matrix  $A$  represent an ON basis of  $\mathbb{R}^5$ ,

$$\frac{1}{2\sqrt{3}} \begin{pmatrix} 1 \\ -\sqrt{3} \\ 2 \\ -\sqrt{3} \\ 1 \end{pmatrix}, \quad \frac{1}{2\sqrt{3}} \begin{pmatrix} 1 \\ \sqrt{3} \\ 2 \\ \sqrt{3} \\ 1 \end{pmatrix}, \quad \frac{1}{2} \begin{pmatrix} -1 \\ 1 \\ 0 \\ -1 \\ 1 \end{pmatrix}, \quad \frac{1}{2} \begin{pmatrix} -1 \\ -1 \\ 0 \\ 1 \\ 1 \end{pmatrix}, \quad \frac{1}{\sqrt{3}} \begin{pmatrix} 1 \\ 0 \\ -1 \\ 0 \\ 1 \end{pmatrix}. \quad (474)$$

Consequently, the representation  $c = \sum_{\nu=1}^5 \gamma_{\nu} v_{\nu}$  has the coefficients  $\gamma_{\nu} = v_{\nu} \cdot c$ ,

$$\gamma_{1,2} = \frac{h^2}{4\sqrt{3}}(2 \mp \sqrt{3}), \quad \gamma_{3,4} = 0, \quad \gamma_5 = \frac{h^2}{4\sqrt{3}}. \quad (475)$$

With the corresponding eigenvalues,

$$\lambda_{1,2} = \mp \frac{1}{4}\sqrt{3}, \quad \lambda_{3,4} = \mp \frac{1}{4}, \quad \lambda_5 = 0, \quad (476)$$

Eq. (465) eventually yields the converged iterative solution

$$u = \frac{h^2(2 - \sqrt{3})}{4\sqrt{3} + 3} v_1 + \frac{h^2(2 + \sqrt{3})}{4\sqrt{3} - 3} v_2 + \frac{h^2}{4\sqrt{3}} v_5 = \frac{h^2}{52} \begin{pmatrix} 19 \\ 24 \\ 25 \\ 24 \\ 19 \end{pmatrix} = \begin{pmatrix} 0.091346 \\ 0.115385 \\ 0.120192 \\ 0.115385 \\ 0.091346 \end{pmatrix}, \quad (477)$$

where  $h = \frac{1}{2}$  was used in the last step.

**The first 8 iterations:** Starting with the trial vector

$$u^{(0)} = \begin{pmatrix} 1.0 \\ 1.0 \\ 1.0 \\ 1.0 \\ 1.0 \end{pmatrix}, \quad (478)$$

we numerically find for  $u^{(k+1)} = Au^{(k)} + c$  the (rounded) results

$$\begin{aligned} u^{(1)} &= \begin{pmatrix} 0.31 \\ 0.56 \\ 0.56 \\ 0.56 \\ 0.31 \end{pmatrix}, & u^{(2)} &= \begin{pmatrix} 0.20 \\ 0.28 \\ 0.34 \\ 0.28 \\ 0.20 \end{pmatrix}, & u^{(3)} &= \begin{pmatrix} 0.13 \\ 0.20 \\ 0.20 \\ 0.20 \\ 0.13 \end{pmatrix}, & u^{(4)} &= \begin{pmatrix} 0.11 \\ 0.15 \\ 0.16 \\ 0.15 \\ 0.11 \end{pmatrix}, \\ u^{(5)} &= \begin{pmatrix} 0.099 \\ 0.131 \\ 0.136 \\ 0.131 \\ 0.099 \end{pmatrix}, & u^{(6)} &= \begin{pmatrix} 0.095 \\ 0.121 \\ 0.128 \\ 0.121 \\ 0.095 \end{pmatrix}, & u^{(7)} &= \begin{pmatrix} 0.093 \\ 0.118 \\ 0.123 \\ 0.118 \\ 0.093 \end{pmatrix}, & u^{(8)} &= \begin{pmatrix} 0.092 \\ 0.116 \\ 0.122 \\ 0.116 \\ 0.092 \end{pmatrix}. \end{aligned}$$

## 12.3 Finite elements method (FEM)

Again, we consider for the unknown function  $u(\mathbf{r})$  the PDE

$$\nabla^2 u(\mathbf{r}) = -s(\mathbf{r}) \quad (\mathbf{r} \in \Omega), \quad (479)$$

where  $s(\mathbf{r})$  is a given source function.

Again,  $\partial\Omega_1$  and  $\partial\Omega_2 = \partial\Omega \setminus \partial\Omega_1$  shall be two disjoint pieces of the boundary  $\partial\Omega$  of  $\Omega$ .

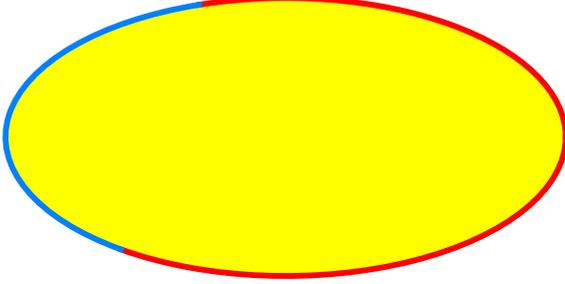


Figure 22: Two disjoint pieces  $\partial\Omega_1$  (blue) and  $\partial\Omega_2$  (red) of the boundary  $\partial\Omega$  of a region  $\Omega$  (yellow).

### 12.3.1 Weak formulation of a PDE

For a given region  $\Omega \subset \mathbb{R}^n$ , let  $\mathcal{B}$  be the set of all functions  $v : \Omega \rightarrow \mathbb{R}$ ,  $\mathbf{r} \mapsto v(\mathbf{r})$  with "reasonable" properties (continuity, two-fold partial differentiability, etc.).

**Theorem:** The following two statements are equivalent:

- (I) The function  $u \in \mathcal{B}$  is a solution of  $\nabla^2 u = -s$  that satisfies the Neumann BC
- (N) 
$$\mathcal{N}(\mathbf{r}) \cdot \nabla u(\mathbf{r}) = g(\mathbf{r}) \quad \forall \mathbf{r} \in \partial\Omega_2. \quad (480)$$

- (II) The function  $u \in \mathcal{B}$  satisfies the equation

$$\int_{\Omega} d^n r \left\{ w(\mathbf{r})s(\mathbf{r}) - [\nabla w(\mathbf{r})] \cdot [\nabla u(\mathbf{r})] \right\} - \int_{\partial\Omega_2} d^{n-1} r w(\mathbf{r})g(\mathbf{r}) = 0, \quad (481)$$

for **all** functions  $w \in \mathcal{B}_1^{\text{hom}}$ , where

$$\mathcal{B}_1^{\text{hom}} = \left\{ w \in \mathcal{B} \mid w(\mathbf{r}) = 0 \quad \forall \mathbf{r} \in \partial\Omega_1 \right\}. \quad (482)$$

**Corollary:** The **unique** solution  $u(\mathbf{r})$  of  $\nabla^2 u = -s$  satisfying both the BCs (N) and

$$(D) \quad u(\mathbf{r}) = f(\mathbf{r}) \quad \forall \mathbf{r} \in \partial\Omega_1 \quad (483)$$

is that function  $u \in \mathcal{B}$  that (i) satisfies Eq. (481) for **all**  $w \in \mathcal{B}_1^{\text{hom}}$  and (ii) satisfies (D). [This is called the **weak formulation** of the PDE  $\nabla^2 u = -s$  with (D) plus (N).]

**Proof of the theorem:** Statement (I) is equivalent to saying that  $u(\mathbf{r})$  has the property

$$0 = \int_{\Omega} d^n r w(\mathbf{r}) \left[ \underbrace{s(\mathbf{r}) + \nabla^2 u(\mathbf{r})}_0 \right] + \int_{\partial\Omega_2} d^{n-1} r w(\mathbf{r}) \left[ \underbrace{\mathcal{N}(\mathbf{r}) \cdot \nabla u(\mathbf{r}) - g(\mathbf{r})}_0 \right], \quad (484)$$

for all  $w \in \mathcal{B}_1^{\text{hom}}$ . Since  $\nabla \cdot (w \nabla u) = w \nabla^2 u + (\nabla w) \cdot (\nabla u)$ , we may write this as

$$0 = \int_{\Omega} d^n r \left[ w s + \nabla \cdot (w \nabla u) - (\nabla w) \cdot (\nabla u) \right] + \int_{\partial\Omega_2} d^{n-1} r w \left[ \mathcal{N} \cdot \nabla u - g \right]. \quad (485)$$

Now, Gauss' theorem yields

$$\begin{aligned} \int_{\Omega} d^n r \left[ \nabla \cdot (w \nabla u) \right] &= \oint_{\partial\Omega} d^{n-1} r \mathcal{N} \cdot (w \nabla u) \\ &\equiv \oint_{\partial\Omega} d^{n-1} r w \left[ \mathcal{N} \cdot \nabla u \right]. \end{aligned} \quad (486)$$

Since  $w \in \mathcal{B}_1^{\text{hom}}$ , we may replace  $\partial\Omega$  with  $\partial\Omega_2$  here, and Eq. (485) becomes Eq. (481). Consequently, statements (I) and (II) are equivalent.

### 12.3.2 System of linear equations

Confining the set  $\mathcal{B}_1^{\text{hom}}$  properly (e.g., by excluding all functions  $w$  that are not square integrable), we may select an infinite series  $\{w_j\}_{j \in \mathbb{N}}$  of linearly independent functions  $w_j \in \mathcal{B}_1^{\text{hom}}$  such that each function  $w \in \mathcal{B}_1^{\text{hom}}$  has a representation

$$w(\mathbf{r}) = \sum_{j=1}^{\infty} \xi_j w_j(\mathbf{r}), \quad (487)$$

where the  $\xi_j$  are constant ( $\mathbf{r}$ -independent) coefficients, fixed uniquely by  $w(\mathbf{r})$ .

Let  $u_{\text{D}}(\mathbf{r})$  be **some** function satisfying (D) and let  $u(\mathbf{r})$  be the **unique solution** of the PDE (479) that also satisfies (N). Since  $u - u_{\text{D}} \in \mathcal{B}_1^{\text{hom}}$ , we have

$$u(\mathbf{r}) = u_{\text{D}}(\mathbf{r}) + \sum_{j=1}^{\infty} w_j(\mathbf{r}) \xi_j. \quad (488)$$

Given the functions  $u_{\text{D}}(\mathbf{r})$  and  $w_1(\mathbf{r}), w_2(\mathbf{r}), \dots$ , we wish to find the coefficients  $\xi_1, \xi_2, \dots$ , that yield via Eq. (488) the unknown wanted function  $u(\mathbf{r})$ .

Choosing in Eq. (481)  $w = w_i$  and using expression (488) for  $u$ , we obtain

$$\sum_{j=1}^{\infty} A_{ij} \xi_j = \eta_i, \quad (489)$$

with the constants

$$\begin{aligned} A_{ij} &= \int_{\Omega} d^n r \left[ \nabla w_i(\mathbf{r}) \right] \cdot \left[ \nabla w_j(\mathbf{r}) \right], \\ \eta_i &= \int_{\Omega} d^n r \left\{ s(\mathbf{r}) w_i(\mathbf{r}) - \left[ \nabla u_{\text{D}}(\mathbf{r}) \right] \cdot \left[ \nabla w_i(\mathbf{r}) \right] \right\} - \int_{\partial\Omega_2} d^{n-1} r g(\mathbf{r}) w_i(\mathbf{r}). \end{aligned} \quad (490)$$

These can be evaluated for any given choice of the functions  $u_{\text{D}}(\mathbf{r})$  and  $w_1(\mathbf{r}), w_2(\mathbf{r}), \dots$ . Then, Eq. (489) provides an infinite system of linear equations for the wanted coefficients  $\xi_1, \xi_2, \dots$ . In matrix form, this system reads

$$A \xi = \eta. \quad (491)$$

### 12.3.3 Summary (3D notation)

- **Given:** A volume region  $\Omega \subset \mathbb{R}^3$  with surface

$$\partial\Omega = \partial\Omega_1 \cup \partial\Omega_2 \quad (\partial\Omega_1 \cap \partial\Omega_2 = \emptyset) \quad (492)$$

and three functions

$$s : \Omega \rightarrow \mathbb{R}, \quad \mathbf{r} \mapsto s(\mathbf{r}), \quad (493)$$

$$f : \partial\Omega_1 \rightarrow \mathbb{R}, \quad \mathbf{r} \mapsto f(\mathbf{r}), \quad (494)$$

$$g : \partial\Omega_2 \rightarrow \mathbb{R}, \quad \mathbf{r} \mapsto g(\mathbf{r}). \quad (495)$$

- **Wanted:** The unique function  $u : \Omega \rightarrow \mathbb{R}$ ,  $\mathbf{r} \mapsto u(\mathbf{r})$  that satisfies:

$$\text{(PDE)} \quad \nabla^2 u(\mathbf{r}) = -s(\mathbf{r}) \quad (\forall \mathbf{r} \in \Omega), \quad (496)$$

$$\text{(D)} \quad u(\mathbf{r}) = f(\mathbf{r}) \quad (\forall \mathbf{r} \in \partial\Omega_1), \quad (497)$$

$$\text{(N)} \quad \mathcal{N}(\mathbf{r}) \cdot \nabla u(\mathbf{r}) = g(\mathbf{r}) \quad (\forall \mathbf{r} \in \partial\Omega_2). \quad (498)$$

Here,  $\mathcal{N}(\mathbf{r})$  is the unit vector normal to  $\partial\Omega_2$  (and pointing outward) at  $\mathbf{r} \in \partial\Omega_2$ .

- **Solution:**

$$u(\mathbf{r}) = u_D(\mathbf{r}) + \sum_{j=1}^{\infty} \xi_j w_j(\mathbf{r}). \quad (499)$$

Here:

- (a) The function  $u_D(\mathbf{r})$  is **some** function  $u_D : \Omega \rightarrow \mathbb{R}$  satisfying (D),

$$u_D(\mathbf{r}) = f(\mathbf{r}) \quad (\forall \mathbf{r} \in \partial\Omega_1), \quad (500)$$

- (b)  $\{w_1, w_2, w_3, \dots\}$  is a basis set of the infinite-dimensional function space<sup>9</sup>

$$\mathcal{B}_1^{\text{hom}} = \left\{ w : \Omega \rightarrow \mathbb{R} \mid w(\mathbf{r}) = 0 \quad (\forall \mathbf{r} \in \partial\Omega_1) \right\}, \quad (502)$$

- (c) the coefficients  $\xi_j$  are fixed by the linear system of equations

$$A\xi = \eta, \quad (503)$$

where the elements  $A_{ij}$  (of the matrix  $A$ ) and  $\eta_i$  (of the column  $\eta$ ) are fixed by: the functions  $s$ ,  $f$  (via  $u_D$ ),  $g$ , and  $\{w_1, w_2, w_3, \dots\}$  via Eqs. (490),

$$A_{ij} = \int_{\Omega} d^n r [\nabla w_i(\mathbf{r})] \cdot [\nabla w_j(\mathbf{r})],$$

$$\eta_i = \int_{\Omega} d^n r \left\{ s(\mathbf{r}) w_i(\mathbf{r}) - [\nabla u_D(\mathbf{r})] \cdot [\nabla w_i(\mathbf{r})] \right\} - \int_{\partial\Omega_2} d^{n-1} r g(\mathbf{r}) w_i(\mathbf{r}).$$

<sup>9</sup>This means: For any function  $w \in \mathcal{B}_1^{\text{hom}}$  there is a unique sequence  $(\xi_1, \xi_2, \xi_3, \dots)$  of coefficients  $\xi_i$  such that

$$w(\mathbf{r}) = \sum_{i=1}^{\infty} \xi_i w_i(\mathbf{r}). \quad (501)$$

### 12.3.4 Finite elements

In the FEM, the vectorspace  $\mathcal{B}_1^{\text{hom}}$  is approximated by a finite-dimensional one, selecting a finite number of basis functions  $w_1(\mathbf{r}), \dots, w_N(\mathbf{r})$ .

For illustration, we consider the 2D problem of Example 2 in section 12.2.3:  $\Omega$  is given by Fig. 21/23, with  $\partial\Omega_1 = \partial\Omega$  and  $\partial\Omega_2 = \emptyset$ ,  $s(\mathbf{r}) \equiv 1$  and  $f(\mathbf{r}) \equiv 0$ ,

$$\nabla^2 u(\mathbf{r}) = -1 \quad (\mathbf{r} \in \Omega), \quad (504)$$

$$u(\mathbf{r}) = 0 \quad (\mathbf{r} \in \partial\Omega). \quad (505)$$

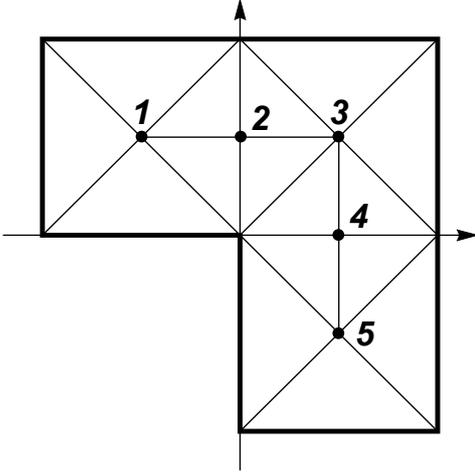


Figure 23: The region  $\Omega$  in the  $xy$ -plane, Fig. 21, divided up into 16 triangles.

For the FEM, we divide  $\Omega$  up into 16 triangles (**finite elements**), as shown in Fig. 23. Five vertices  $\mathbf{r}_1, \dots, \mathbf{r}_5$  of these triangles are lying **inside**  $\Omega$  (dots in Fig. 23),

$$\mathbf{r}_1 = (-\frac{1}{2}|\frac{1}{2}), \quad \mathbf{r}_2 = (0|\frac{1}{2}), \quad \mathbf{r}_3 = (\frac{1}{2}|\frac{1}{2}), \quad \mathbf{r}_4 = (\frac{1}{2}|0), \quad \mathbf{r}_5 = (\frac{1}{2}|-\frac{1}{2}), \quad (506)$$

the remaining eight vertices (with negative subscripts)  $\mathbf{r}_{-1}, \dots, \mathbf{r}_{-8}$  are on the boundary  $\partial\Omega$  (not marked in Fig. 23). For each vertex  $\mathbf{r}_i$ , a function  $w_i(\mathbf{r})$  is introduced as that **continuous** function of  $\mathbf{r} \in \Omega$ , which (i) within each one of the 16 triangles is a **linear function**  $w(x, y) = ax + by + c$  of  $x$  and  $y$  and (ii) on the 13 vertices assumes the values

$$w_i(\mathbf{r}_j) = \delta_{ij} \equiv \begin{cases} 1 & (j = i), \\ 0 & (j \neq i). \end{cases} \quad (507)$$

In other words:  $w_i(\mathbf{r})$  assumes its maximum value 1 at  $\mathbf{r} = \mathbf{r}_i$ , falls linearly from 1 to 0 within each triangle sharing the vertex  $\mathbf{r}_i$ , and is zero within all the remaining triangles.<sup>10</sup> The functions  $w_i(\mathbf{r})$  with  $i < 0$  are needed only for the approximation

$$u_D(\mathbf{r}) \approx U_D(\mathbf{r}) = \sum_{i=-1}^{-8} f(\mathbf{r}_i)w_i(\mathbf{r}) \quad (508)$$

<sup>10</sup>Writing  $\mathbf{r}_i = (x_i, y_i)$  and  $X = x - x_i$ ,  $Y = y - y_i$ , we explicitly have for  $i = 1, 3, 5$

$$w_i(x, y) = \begin{cases} 1 - 2X & (X \in [0, \frac{1}{2}], |Y| \leq X), \\ 1 - 2Y & (Y \in [0, \frac{1}{2}], |X| \leq Y), \\ 1 + 2X & (X \in [-\frac{1}{2}, 0], |Y| \leq -X), \\ 1 + 2Y & (Y \in [-\frac{1}{2}, 0], |X| \leq -Y), \end{cases}$$

to the function  $u_D(\mathbf{r})$  in Eq. (488), while the complete  $u(\mathbf{r})$  there is approximated by

$$U(\mathbf{r}) = U_D(\mathbf{r}) + \sum_{i=1}^5 \xi_i w_i(\mathbf{r}). \quad (509)$$

We see that  $\mathcal{B}_1^{\text{hom}}$  is approximated here by a **5-dimensional** vectorspace. Due to Eq. (507), we also see that  $\xi_i = U(\mathbf{r}_i)$  for  $i > 0$ . Moreover, since  $f(\mathbf{r}) = 0$  here, we have

$$U_D(\mathbf{r}) = 0. \quad (510)$$

Now, with  $\mathbf{r} = (x, y)$ ,  $u_D(\mathbf{r}) \approx U_D(\mathbf{r}) = 0$ ,  $s(\mathbf{r}) = 1$ , and  $\partial\Omega_2 = \emptyset$ , Eqs. (490) read

$$\eta_i = \int_{\Omega} d^2r w_i(\mathbf{r}), \quad A_{ij} = \int_{\Omega} d^2r [\nabla w_i(\mathbf{r})] \cdot [\nabla w_j(\mathbf{r})]. \quad (511)$$

The support  $\Sigma_i = \{\mathbf{r} \in \Omega \mid w_i(\mathbf{r}) \neq 0\}$  of  $w_i(\mathbf{r})$  is a square with area  $S_i = 1$  for  $i = 1, 3, 5$  and  $S_i = \frac{1}{2}$  for  $i = 2, 4$ . Consequently, since  $\langle w_i(\mathbf{r}) \rangle_{\mathbf{r} \in \Sigma_i} = \frac{1}{3}$ , we have

$$\eta_i \equiv S_i \langle w_i(\mathbf{r}) \rangle_{\mathbf{r} \in \Sigma_i} = \frac{S_i}{3} = \begin{cases} \frac{1}{3} & (i = 1, 3, 5), \\ \frac{1}{6} & (i = 2, 4). \end{cases} \quad (512)$$

For  $\mathbf{r} \in \Sigma_i$ , it is easy to verify that  $|\nabla w_i(\mathbf{r})|^2 = 4$  for  $i = 1, 3, 5$ , while  $|\nabla w_i(\mathbf{r})|^2 = 8$  for  $i = 2, 4$ . Consequently, we have

$$A_{ii} \equiv S_i |\nabla w_i(\mathbf{r})|^2 = 4 \quad (i = 1, \dots, 5). \quad (513)$$

In the overlap region  $\Sigma_1 \cap \Sigma_2$  between  $\Sigma_1$  and  $\Sigma_2$ , we have  $\nabla w_1(\mathbf{r}) = -\binom{2}{0}$ , while  $\nabla w_2(\mathbf{r}) = \binom{+2}{-2}$  for  $y > \frac{1}{2}$  and  $\nabla w_2(\mathbf{r}) = \binom{2}{2}$  for  $y < \frac{1}{2}$ . Consequently,

$$A_{12} = \int_{\Sigma_1 \cap \Sigma_2} d^2r \begin{pmatrix} -2 \\ 0 \end{pmatrix} \cdot \begin{pmatrix} 2 \\ \pm 2 \end{pmatrix} = \frac{1}{4} \cdot (-4) = -1, \quad (514)$$

where  $\frac{1}{4}$  is the area of the triangle  $\Sigma_1 \cap \Sigma_2$ . Similarly, we find  $A_{ij} = -1$  for  $|i - j| = 1$  and  $A_{ij} = 0$  for all other matrix elements  $A_{ij}$  with  $i \neq j$ . In summary,

$$A = \begin{pmatrix} 4 & -1 & 0 & 0 & 0 \\ -1 & 4 & -1 & 0 & 0 \\ 0 & -1 & 4 & -1 & 0 \\ 0 & 0 & -1 & 4 & -1 \\ 0 & 0 & 0 & -1 & 4 \end{pmatrix}, \quad \eta = \frac{1}{6} \begin{pmatrix} 2 \\ 1 \\ 2 \\ 1 \\ 2 \end{pmatrix}. \quad (515)$$

The solution  $\xi$  of the corresponding system  $A\xi = \eta$  of 5 linear equations is

$$\xi = \frac{1}{156} \begin{pmatrix} 17 \\ 16 \\ 21 \\ 16 \\ 17 \end{pmatrix} = \begin{pmatrix} 0.1090 \\ 0.1026 \\ 0.1346 \\ 0.1026 \\ 0.1090 \end{pmatrix}. \quad (516)$$

An accurate computation yields  $u(\frac{1}{2}, \frac{1}{2}) = 0.1310$  [Sewell, p. 224], to be compared with the present value  $U(\mathbf{r}_3) \equiv \xi_3 = 0.1346$ .

and for  $i = 2, 4$

$$w_i(x, y) = \begin{cases} 1 - 2X - 2Y & (X \in [0, \frac{1}{2}], Y \in [0, \frac{1}{2} - X]), \\ 1 + 2X - 2Y & (X \in [-\frac{1}{2}, 0], Y \in [0, \frac{1}{2} + X]), \\ 1 + 2X + 2Y & (X \in [-\frac{1}{2}, 0], Y \in [-\frac{1}{2} - X, 0]), \\ 1 - 2X + 2Y & (X \in [0, \frac{1}{2}], Y \in [-\frac{1}{2} + X, 0]), \end{cases}$$

while  $w_i(x, y) = 0$  for all other values of  $X, Y$ .